



# Observation of Electron Antineutrino Disappearance at Daya Bay

Yifang Wang Institute of High Energy Physics CERN, March 20, 2012

F.P. An et al., Daya Bay Coll., "A side-by-side comparison of Daya Bay anti-neutrino detectors", arXiv: 1202.6181[physics.ins-det], submitted to NIM

F.P. An et al., Daya Bay Coll., "Observation of electron anti-neutrino disappearance at Daya Bay", arXiv: 1203.1669[hep-ex], submitted to PRL

# **Neutrinos & Neutrino Oscillation**

**Fundamental building blocks of matter:** 

$$\begin{pmatrix} e & \mu & \tau \\ v_e & v_\mu & v_\tau \end{pmatrix} \quad \begin{pmatrix} u & c & t \\ d & s & b \end{pmatrix}$$

Neutrino mass: the central issue of neutrino physics

- ⇒ Tiny mass but huge amount
- ⇒ Influence to Cosmology: evolution, large scale structure, ...
- ⇒ Only evidence beyond the Standard Model
- Neutrino oscillation: a great method to probe the mass

$$\frac{v_e}{v_e} \xrightarrow{v_\mu} \frac{v_e}{v_e} \xrightarrow{v_\mu} \frac{v_e}{v_\mu}$$
Oscillation  
probability:  
$$P(v_e \rightarrow v_\mu) = \sin^2(2\theta) \sin^2(1.27\Delta m^2 L/E)$$
Oscillation  
amplitude  
$$Oscillation frequency$$

2012-03-08



• Neutrino mixing matrix:

$$\mathbf{V} = egin{pmatrix} 1 & 0 & 0 \ 0 & \mathbf{c_{23}} & \mathbf{s_{23}} \ 0 & -\mathbf{s_{23}} & \mathbf{c_{23}} \end{pmatrix} egin{pmatrix} \mathbf{c_{13}} & 0 & \mathbf{s_{13}} \ 0 & \mathbf{e^{-i\delta}} & 0 \ -\mathbf{s_{12}} & \mathbf{c_{12}} & \mathbf{s_{12}} & 0 \ -\mathbf{s_{12}} & \mathbf{c_{12}} & \mathbf{c_{12}} & 0 \ 0 & \mathbf{e^{i\sigma}} & 0 \ 0 & \mathbf{e^{i\sigma}} & 0 \ 0 & 0 & 1 \end{pmatrix} egin{pmatrix} \mathbf{e^{i\rho}} & 0 & 0 \ 0 & \mathbf{e^{i\sigma}} & 0 \ 0 & 0 & 1 \end{pmatrix}$$

Unknown mixing parameters:  $\theta_{13}$ ,  $\delta$  + 2 Majorana phases

Need sizable  $\theta_{13}$  for the  $\delta$  measurement

2012-03-08

# **Two ways to measure** $\theta_{13}$

## **Reactor experiments:**

$$\begin{split} P_{ee} &\approx 1 - \sin^2 2\theta_{13} \sin^2 (1.27 \Delta m_{13}^2 L/E) - \\ &\quad \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 (1.27 \Delta m_{12}^2 L/E) \end{split}$$
  $\begin{aligned} \text{Long baseline accelerator experiments:} \\ P_{\mu e} &\approx \sin^2 \theta_{23} \sin^2 2\theta_{13} \sin^2 (1.27 \Delta m_{23}^2 L/E) + \\ &\quad \cos^2 \theta_{23} \sin^2 2\theta_{12} \sin^2 (1.27 \Delta m_{12}^2 L/E) - \end{aligned}$ 

 $A(\rho) \bullet \cos^2 \theta_{13} \sin \theta_{13} \bullet \sin(\delta)$ 



## At reactors:

- $\blacktriangleright$  Clean signal, no cross talk with  $\delta$  and matter effects
- Relatively cheap compared to accelerator based experiments
- Provides the direction to the future of neutrino physics

# **Direct Searches in the Past**



## Double Chooz: 1.7 σ

 $\sin^2 2\theta_{13} = 0.086 \pm 0.041 (\text{stat}) \pm 0.030 (\text{sys})$ 

# Measuring $\theta_{13}$ with Reactor Experiments





# Measuring $\theta_{13}$ with Reactor Experiments





Absolute Reactor Flux Largest uncertainty in previous measurements

Relative Measurement Removes absolute uncertainties!

First proposed by L. A. Mikaelyan and V.V. Sinev, Phys. Atomic Nucl. 63 1002 (2000)



# Reactor Experiment: comparing observed/expected neutrinos



## **Our design goal:** a precision of ~ 0.4%

2012/3/20





Karsten Heeger, Univ. of Wisconsin

# **Daya Bay Experiment: Layout**



- Relative measurement to cancel Corr. Syst. Err.
  - ⇒ 2 near sites, 1 far site
- Multiple AD modules at each site to reduce Uncorr. Syst. Err.
  - ⇒ Far: 4 modules, near: 2 modules
    Cross check; Reduce errors by 1/√N
- Multiple muon detectors to reduce veto eff. uncertainties
  - ➡ Water Cherenkov: 2 layers
  - ⇒ **RPC:** 4 layers at the top + telescopes

# **Underground Labs**



	Overburden (MWE)	$\frac{R_{\mu}}{(Hz/m^2)}$	E <sub>μ</sub> (GeV)	D1,2 (m)	L1,2 (m)	L3,4 (m)
EH1	250	1.27	57	364	857	1307
EH2	265	0.95	58	1348	480	528
EH3	860	0.056	137	1912	1540	1548

2012/3/20



Karsten Heeger, Univ. of Wisconsin

EWNP Symposium, March 8, 2012

Hall 1: began 2 AD operation on Sep. 23, 2011

D1

**D2** 

Daya Bay NPP

## **Daya Bay Antineutrino Detectors**





Karsten Heeger, Univ. of Wisconsin

EWNP Symposium, March 8, 2012

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 $\left[\frac{P_{\rm sur}(E, L_{\rm f})}{P_{\rm sur}(E, L_{\rm n})}\right]$ 

192 8"-PMTs

 $\left| \frac{\epsilon_{\rm f}}{\epsilon_{\rm n}} \right|$ 



Prompt energy (MeV)

Delayed energy (MeV)

## Antineutrino Detector Installation - Near Hall





## Antineutrino Detector Installation - Far Hall





# March 8, 2012 : Daya Bay results



hall. Comparing with the prediction based on the near-hall measurements, a deficit of 6.0% was found. A rate-only analysis yielded  $\sin^2 2\theta_{13} = 0.092 \pm 0.016(\text{stat}) \pm 0.005(\text{syst})$ . The neutrino mixing angle  $\theta_{13}$  is non-zero with a significance of 5.2 standard deviations.

 $\chi^2$  Analysis



 $\chi^2$  Analysis



# **Uncertainty Summary**



Detector				1	For near/far oscillation.		
	Efficiency	Correlated	Uncorrelated		only uncorrelated		
Target Protons		0.47%	0.03%		uncertainties are used.		
Flasher cut	99.98%	0.01%	0.01%				
Delayed energy cut	90.9%	0.6%	0.12%				
Prompt energy cut	99.88%	0.10%	0.01%		Largest systematics are		
Multiplicity cut		0.02%	< 0.01%	>	smaller than far site statistics		
Capture time cut	98.6%	0.12%	0.01%	1	(~1%)		
Gd capture ratio	83.8%	0.8%	<0.1%				
Spill-in	105.0%	1.5%	0.02%				
Livetime	100.0%	0.002%	< 0.01%				
Combined	78.8%	1.9%	0.2%				
Reactor							
Correlate	d	Uncorrelated			Influence of uncorrelated		
Energy/fission	0.2%	Power	0.5%		reactor systematics reduced		
$\overline{\nu}_{e}$ /fission	3%	Fission fraction	0.6%	/	$(\sim 1/20)$ by far vs. near		
		Spent fuel	0.3%		measurement.		
Combined	3%	Combined	0.8%				

Karsten Heeger, Univ. of Wisconsin

# Future plan

- Assembly of AD7 and AD8 is underway now, to be completed before summer
- Current data taking will continue until the summer
- Summer activities:
  - ➡ Installation of AD7 & AD8
  - ⇒ Detector calibration
- Re-start data taking after summer

# **The Daya Bay Collaboration**

#### Political Map of the World, June 1999

Europe (2) JINR, Dubna, Russia Charles University, Czech Republic

## North America (16)

BNL, Caltech, LBNL, Iowa State Univ.,
Illinois Inst. Tech., Princeton, RPI,
UC-Berkeley, UCLA, Univ. of Cincinnati,
Univ. of Houston, Univ. of Wisconsin,
William & Mary, Virginia Tech.,
Univ. of Illinois-Urbana-Champaign, Siena

## ~250 Collaborators

IHEP, Beijing Normal Univ., Chengdu Univ. of Sci. and Tech., CGNPG, CIAE, Dongguan Polytech. Univ., Nanjing Univ., Nankai Univ., NCEPU, Shandong Univ., Shanghai Jiao tong Univ., Shenzhen Univ.,

**Asia** (20)

Tsinghua Univ., USTC, Zhongshan Univ., Univ. of Hong Kong, Chinese Univ. of Hong Kong, National Taiwan Univ., National Chiao Tung Univ., National United Univ.



 Electron anti-neutrino disappearance is observed at Daya Bay,

 $R = 0.940 \pm 0.011 \text{ (stat)} \pm 0.004 \text{ (syst)},$ 

together with a spectral distortion

• A new type of neutrino oscillation is thus discovered

Sin<sup>2</sup>2 $\theta_{13}$ =0.092± 0.016 (stat)±0.005(syst)  $\chi^2$ /NDF = 4.26/4 5.2  $\sigma$  for non-zero  $\theta_{13}$ 

- $\theta_{13}$  was one of the few standard model parameters still unknown.
- It is one of the most discriminant parameters to select neutrino mass matrixes, a key ingredient to decide grand unified theories (if any).
- Non-zero  $\theta_{13}$  is necessary to build-up leptonic CP violation. The value (order of magnitude) of  $\theta_{13}$  is necessary to optimize new facilities to measure leptonic CP violation.

# A refresher on neutrino mixing

The flavor state of the neutrino,  $v_{\alpha}$ , is related to the mass states,  $v_i$ , by a mixing matrix,  $U_{\alpha i}$ 

$$v_i > = \sum U_{\alpha i} | v_{\alpha} >$$

Since there are three observed flavors of neutrinos ( $v_e$ ,  $v_\mu$ ,  $v_\tau$ ), U contains three mixing angles ( $\theta_{12}$ ,  $\theta_{23}$ ,  $\theta_{13}$ ) and a CP violating phase  $\delta$ .

$$\begin{pmatrix} v_{e} \\ v_{\mu} \\ v_{\tau} \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} v_{1} \\ v_{2} \\ v_{3} \end{pmatrix}$$

$$c_{ij} = \cos\theta_{ij}, s_{ij} = \sin\theta_{ij}$$

$$C_{ij} = \cos\theta_{ij}, s_{ij} = \sin\theta_{ij}$$

$$(C_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$(CP \text{ sector'': } \theta_{13} < 11^{\circ} \\ (CP \text{ solar'': } \theta_{12} ~ 34^{\circ} )$$

Rather different from quark mixing:

- Nearly diagonal unitary matrix
- Small angles: θ<sup>CKM</sup><sub>12</sub>~13.0°, θ<sup>CKM</sup><sub>23</sub>~2.3°, θ<sup>CKM</sup><sub>13</sub>~0.2°

Open questions:

sector?

- Is 0<sub>23</sub> exactly 45 degrees, or not?
- Is  $\theta_{13}$  zero, or just small?
- Is there CP violation in the neutrino

### Reactors vs Accelerators

Accelerators:  $\nu_e$  appearance

$$P_{\nu_{\mu} \to \nu_{e}} = 4c_{13}^{2}s_{13}^{2}s_{23}^{2}\sin^{2}\frac{\Delta m_{13}^{2}L}{4E} \times \left[1 \pm \frac{2a}{\Delta m_{13}^{2}}(1 - 2s_{13}^{2})\right] \qquad \theta_{13} \text{ driven}$$

$$+ 8c_{13}^{2}s_{12}s_{13}s_{23}(c_{12}c_{23}cos\delta - s_{12}s_{13}s_{23})\cos\frac{\Delta m_{23}^{2}L}{4E}\sin\frac{\Delta m_{13}^{2}L}{4E}\sin\frac{\Delta m_{12}^{2}L}{4E}\text{ CPer}$$

$$\mp 8c_{13}^{2}c_{12}c_{23}s_{12}s_{13}s_{23}\sin\delta\sin\frac{\Delta m_{23}^{2}L}{4E}\sin\frac{\Delta m_{13}^{2}L}{4E}\sin\frac{\Delta m_{12}^{2}L}{4E} \text{ CPodd}$$

$$+ 4s_{12}^{2}c_{13}^{2}\{c_{13}^{2}c_{23}^{2} + s_{12}^{2}s_{23}^{2}s_{13}^{2} - 2c_{12}c_{23}s_{12}s_{23}s_{13}cos\delta\}\sin\frac{\Delta m_{12}^{2}L}{4E} \text{ solar driven}$$

$$\mp 8c_{12}^{2}s_{13}^{2}s_{23}\cos\frac{\Delta m_{23}^{2}L}{4E}\sin\frac{\Delta m_{13}^{2}L}{4E}(1 - 2s_{13}^{2}) \text{ matter effect (CP odd)}$$

#### Reactors: $\overline{\nu}_e$ disappearance

 $1 - P_{\overline{\nu}_e - \overline{\nu}_e} \simeq \sin^2 2\theta_{13} \sin^2(\Delta m_{31}^2 L/4E) + (\Delta m_{21}^2/\Delta m_{31}^2)^2 (\Delta m_{31}^2 L/4E)^2 \cos^4 \theta_{13} \sin^2 2\theta_{12}$ 



### **CHOOZ** final results

- Analysis A  $\overline{\nu}_e$  spectrum after background subtraction. Both the absolute rate and the spectrum are used.
- Analysis B Uses the different baseline

 $(\Delta L = 117.7 \text{ m})$  of the two reactors. Many systematic errors cancel, but statistical errors are bigger and the  $\Delta m^2$  sensitivity is reduced by the shorter baseline.

• Analysis C Only spectrum information is used.

#### 1450 citations: the top cited null result in hep ever !

### The T2K Experiment



## T2K result, PRL 107 (2011) 041801



## MINOS, PRL 107 (2011) 181802



pot	MINOS 8.2 10 <sup>20</sup>	T2K 1.45 10 <sup>20</sup>
tjoule	1.57	0.07
tjoule kton	7.85	1.57



### The three reactor players

Setup	$P_{\mathrm{Th}}$ [GW]	<i>L</i> [m]	$m_{ m Det}$ [t]	Events/year	Backgrounds/day
Daya Bay	17.4	1700	80	$10 \cdot 10^4$	0.4
Double Chooz	8.6	1050	8.3	$1.5\cdot 10^4$	3.6
RENO	16.4	1400	15.4	$3 \cdot 10^4$	2.6





## **Double Chooz**

#### Talk by J. Dawson



#### 2 cores - 1 site - 8.5 GW<sub>th</sub>

#### 1 near position, 1 far

- target: 2 x 8.3 t
   Civil engineering
- 1 near lab ~ Depth 40 m, Ø 6 m
- 1 available lab

#### Statistics (including ɛ)

- far: ~ 40 evts/day
- near: ~ 460 evts/day

#### Systematics

- reactor : ~ 0.2%
- detector : ~ 0.5%

#### Backgrounds

- $\sigma_{b2b}$  at far site: ~ 1%
- $\sigma_{h2h}$  at near site: ~ 0.5%

#### Planning

- 1. Far detector only
- Sensitivity (1.5 ans) ~ 0.06
- 2. Far + Near sites
  - available from 2010
  - Sensitivity (3 years) ~ 0.025

### RENO



Mauro Mezzetto (INFN Padova)

Rassegna Sperimentale sulle Oscillazioni dei Neutrini

### Daya Bay



### **Reactor detectors**



### **Experimental Results**

T2K ( $\theta_{13} > 0$  @ 2.5 $\sigma$ ) Expected events: 1.5, Detected 6

**Double Chooz (1.3** $\sigma$ ) Expected events: 4344, Detected 4101  $R_{DC} = 0.944 \pm 0.016(\text{stat}) \pm 0.040(\text{syst})$ 

Daya Bay (5.2 $\sigma$ ) Expected events: 85506, Detected 80376  $R_{DB} = 0.940 \pm 0.011(\text{stat}) \pm 0.004(\text{syst})$ 

#### **RENO (4.**9 $\sigma$ ) Expected events:149905, Detected 137912 $R_R = 0.920 \pm 0.009(\text{stat.}) \pm 0.014(\text{syst.})$

### Spectral information

Not used in the fit



Summary of  $\theta_{13}$  results Computed for  $\Delta m^2_{23} = 2.4 \cdot 10^{-3} \text{ eV}^2$ 



### Conclusions

The measurement of  $\theta_{13}$  solves one of the few question marks still left in the standard model. Among the many fondamental consequences, it opens the door to future long-baseline neutrino experiments addressing leptonic CP violation.

Five experimental results in the past 9 months, coming from accelerators and reactors, provided exciting information about  $\theta_{13}$ .

Leptonic CP violation, measurable only at accelerators, will require challenging experimental improvements. The optimization of future facilities is now possible by knowing the  $\theta_{13}$  value.

A worldwide effort is ongoing with multiple proposals in three different continents.