

# Neutrino Telescopes 2011 Highlights



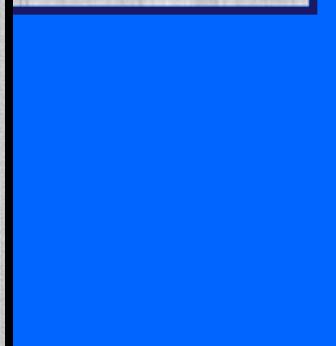
M.Laveder

CSN2 @ Frascati - 11 aprile 2011

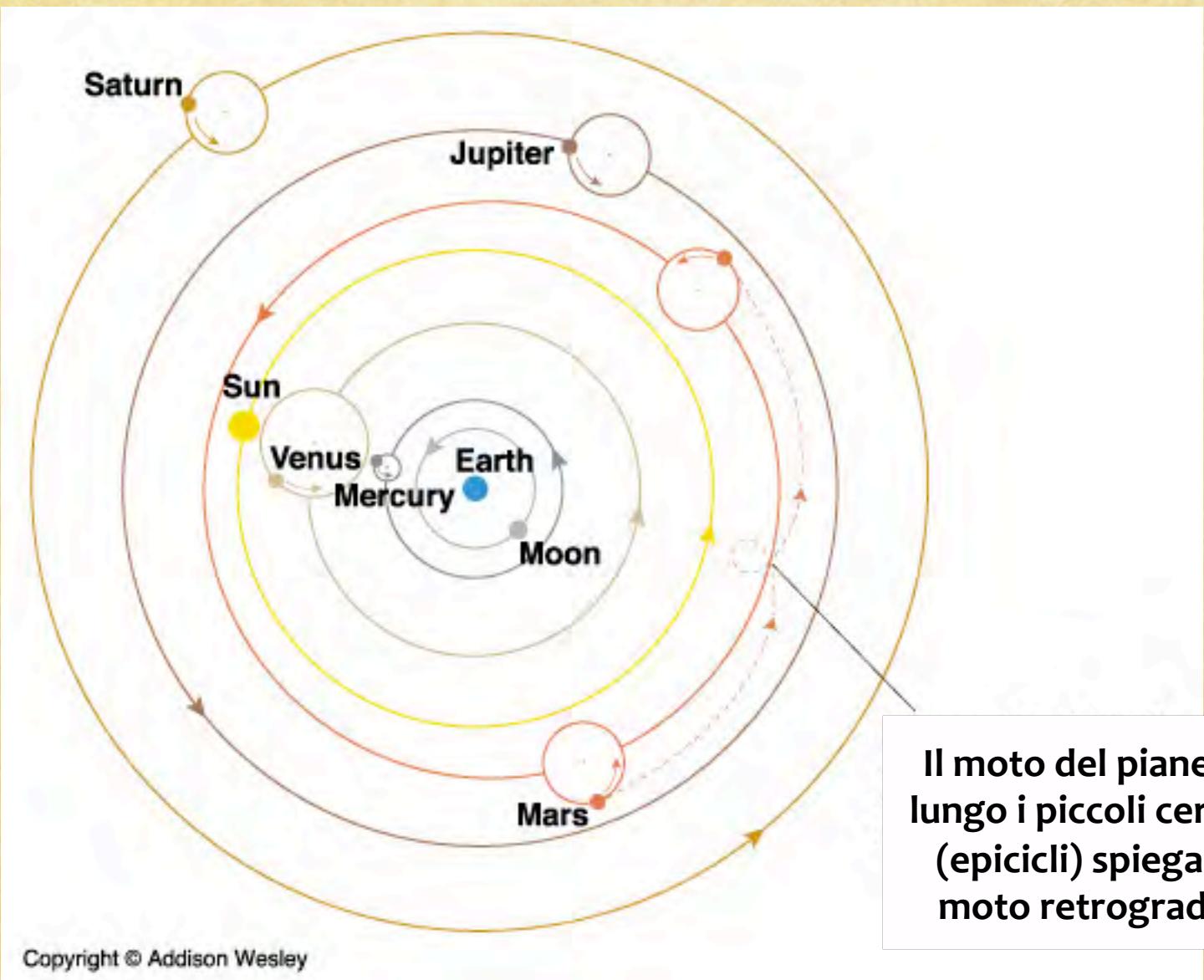
# Il cannocchiale di Galileo sembra infrangere le incorruttibili sfere e rimuovere l'orizzonte



I nuovi  
cannocchiale !

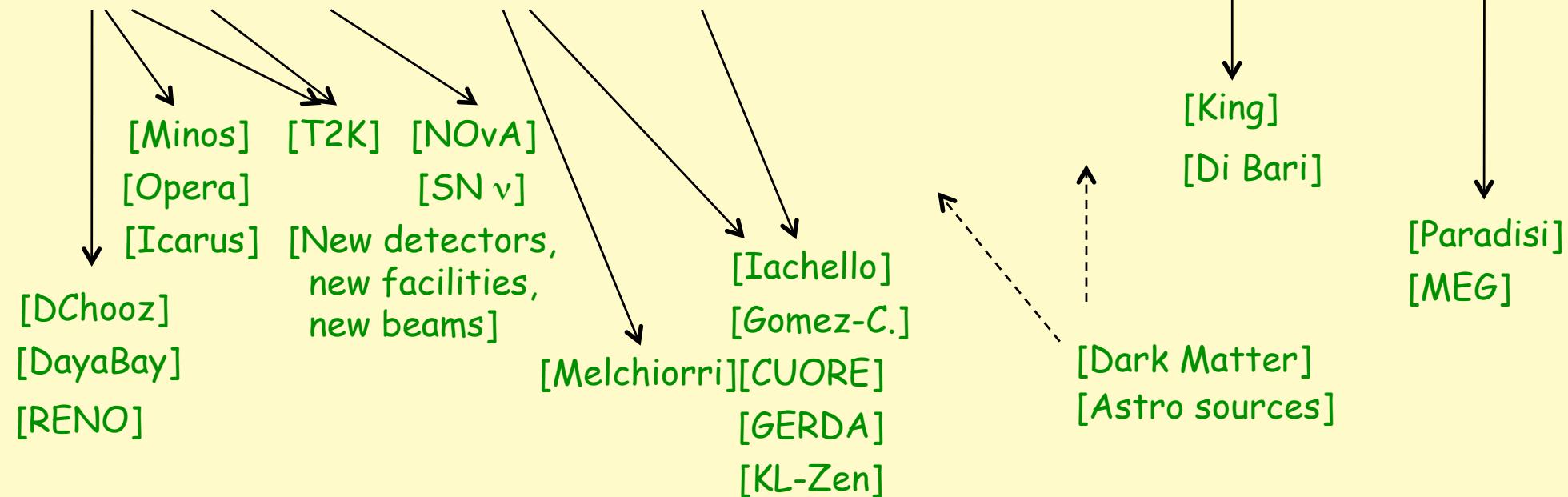


# La soluzione Tolemaica



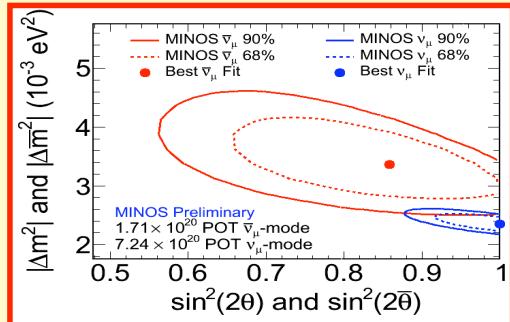
# Unknowns:

$\theta_{13}$ ,  $\delta_{CP}$ ,  $\text{sign}(\Delta m^2)$ ,  $m_\nu$ , Majorana  $\nu$ , ...and a deeper understanding of flavor!

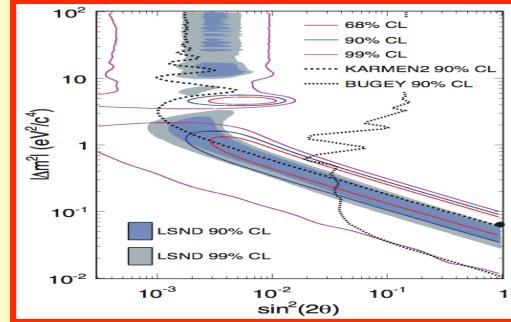


## Some “mismatches:”

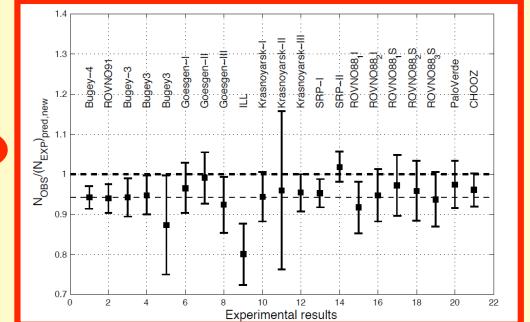
[Corwin] [Maltoni]  
 $\nu$  / anti- $\nu$  (MINOS) ?



[Mills] [Rubbia] [Giunti] [Lasserre]  
 LSND/MiniBooNE ?  
 very SBL reactor dispapp.?



νs?



The rich NEUTEL 2011 program (talks and posters) reflects an evolving field, open to surprises and challenges, both within and ...

## ...beyond the standard $3\nu$ framework

But that's a land with no boundaries...

So, let me discuss just two examples  
of possible surprises and challenges in  
 $\nu$  non-oscillation and oscillation physics

# Results from Super-Kamiokande

Jeffrey Wilkes

University of Washington

For the Super-K collaboration\*

15-Mar-2011, Neutel11

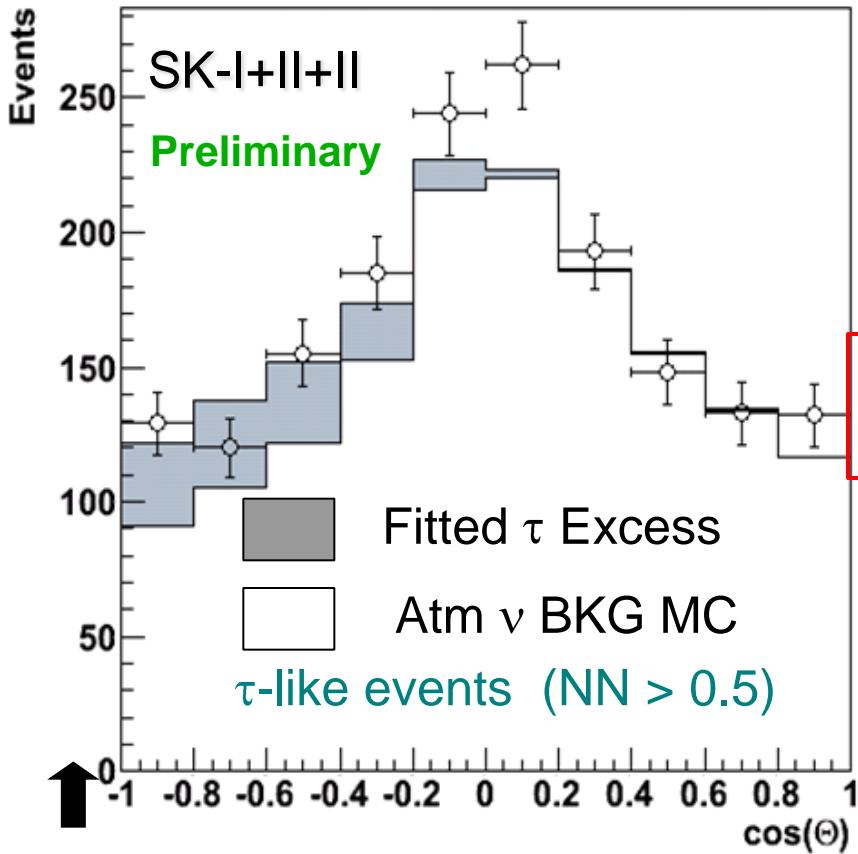


\* Thanks to many  
colleagues for slides;  
errors are all mine

# Fit Results

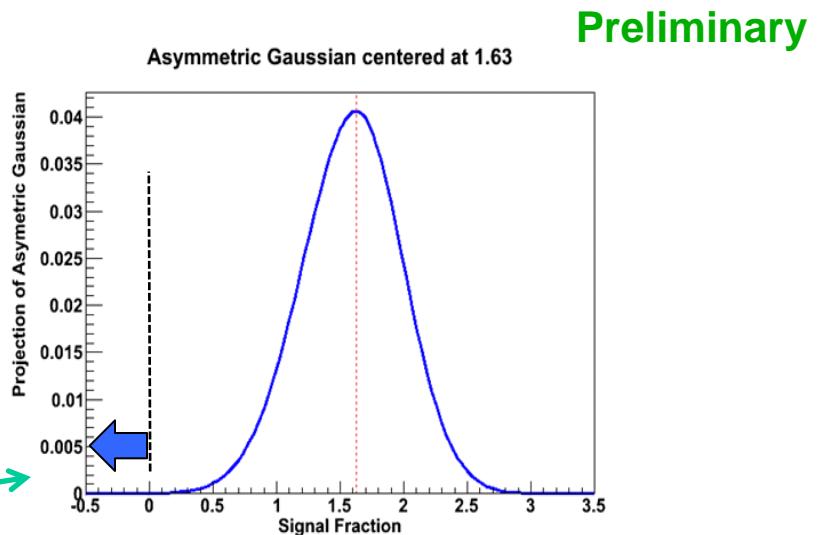
If no  $\tau$  appearance ,  $\beta = 0$

$$Data = \alpha(\gamma) \times bkg + \boxed{\beta(\gamma)} \times signal$$



- » Tau signal clearly appears in upward-going region
- » DIS fits to  $+1\sigma$
- »  $\tau$  normalization fit is  $1.63 \times$  expectation

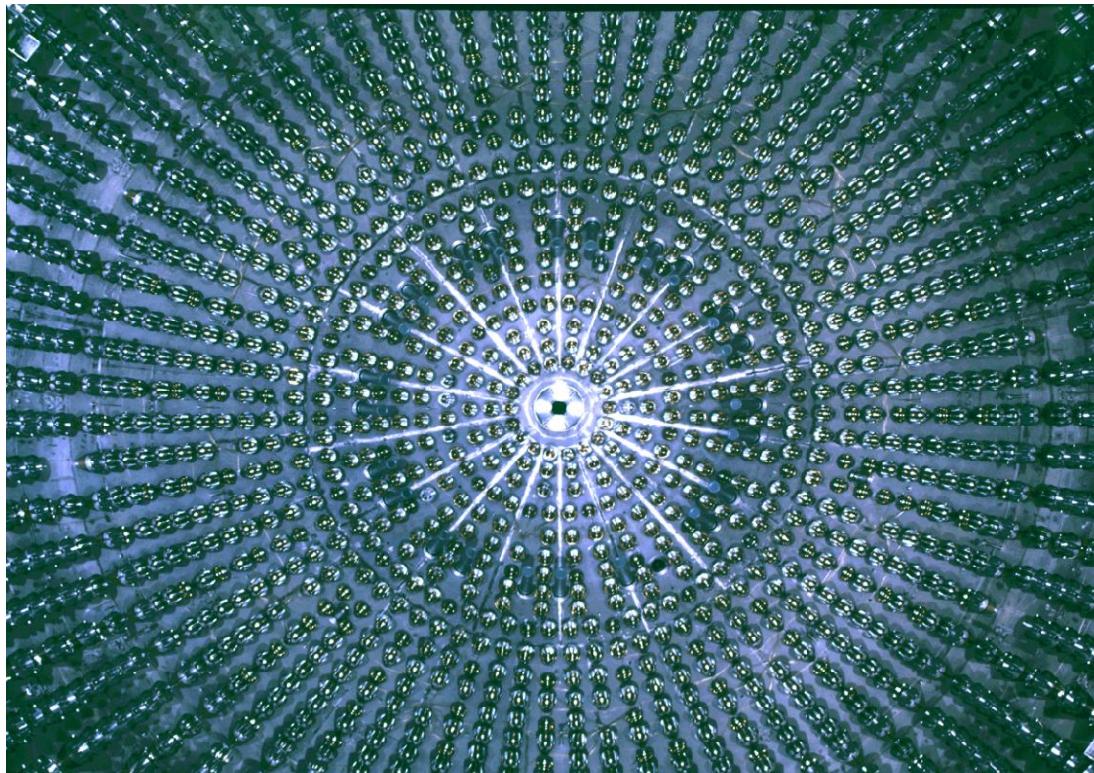
$$\beta = 1.63 \pm 0.35 \begin{array}{l} +0.10 \\ (stat) \end{array} \begin{array}{l} +0.02 \\ (sys) \end{array} \begin{array}{l} -0.08 \\ (3\text{ flav}) \end{array}$$



SK data are *inconsistent* with *no*  $\tau$  appearance at  $3.8\sigma$



# Neutrino Physics with the Borexino experiment

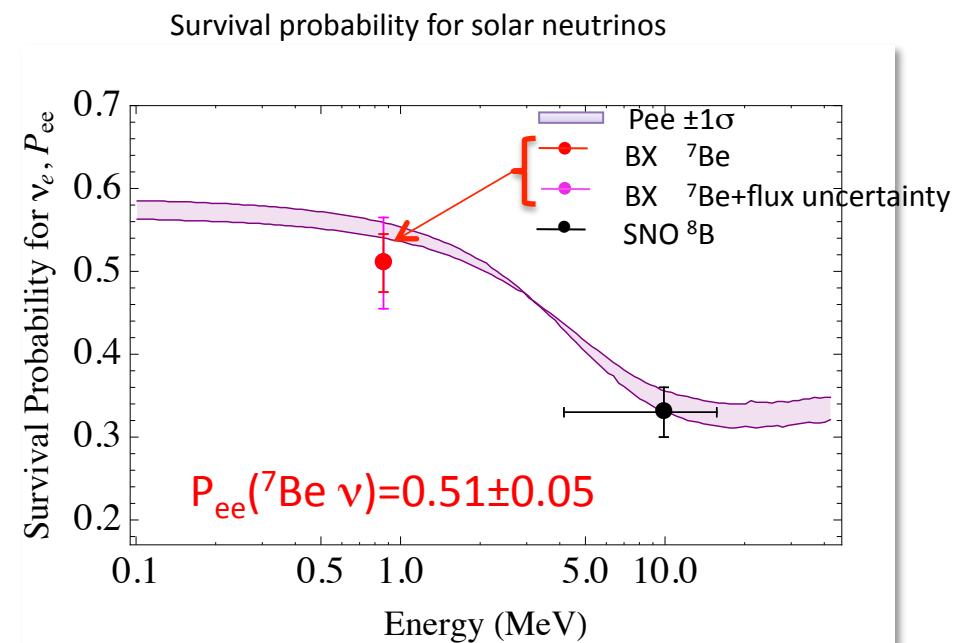


Emanuela Meroni  
On behalf of the Borexino Collaboration

Hypothesis	Expected rate (cpd/100t)
No oscillation + High Metallicity	74±4
No oscillation + Low Metallicity	67±4
Oscillation MSW + High Metallicity	48±4
Oscillation MSW + Low Metallicity	44±4

BX measurement confirms oscillations but cannot discriminate between High and Low metallicity

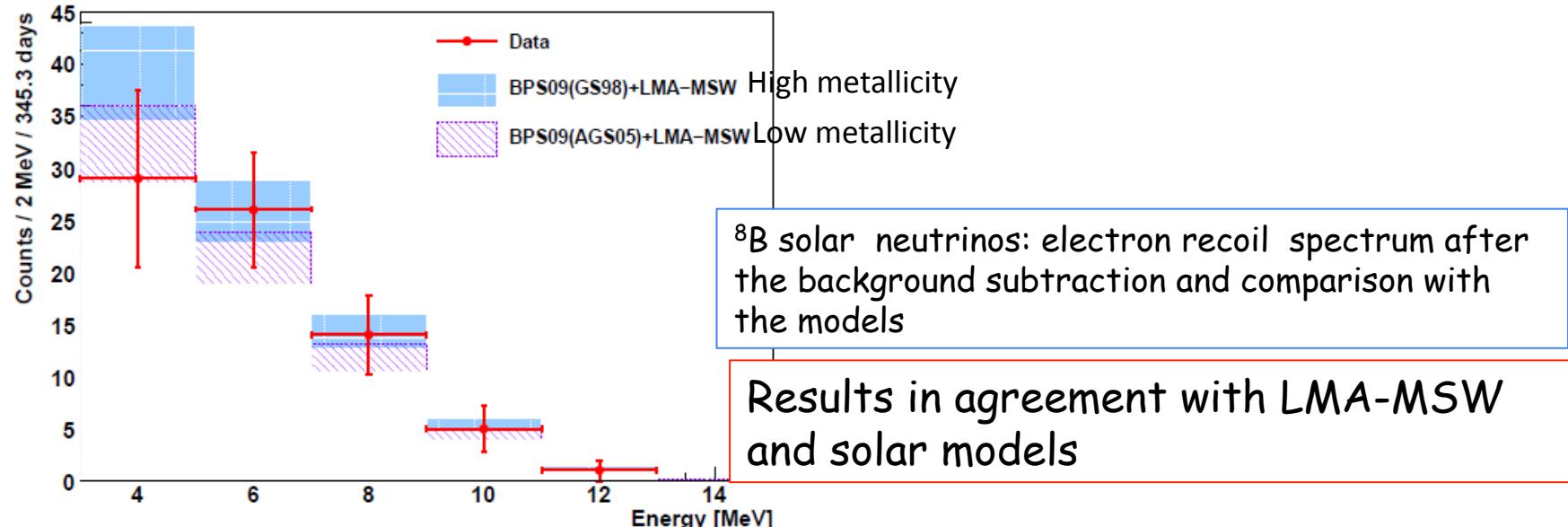
**$46.0 \pm 1.5 \text{ (stat)} \pm 1.3 \text{ (sys)}$**



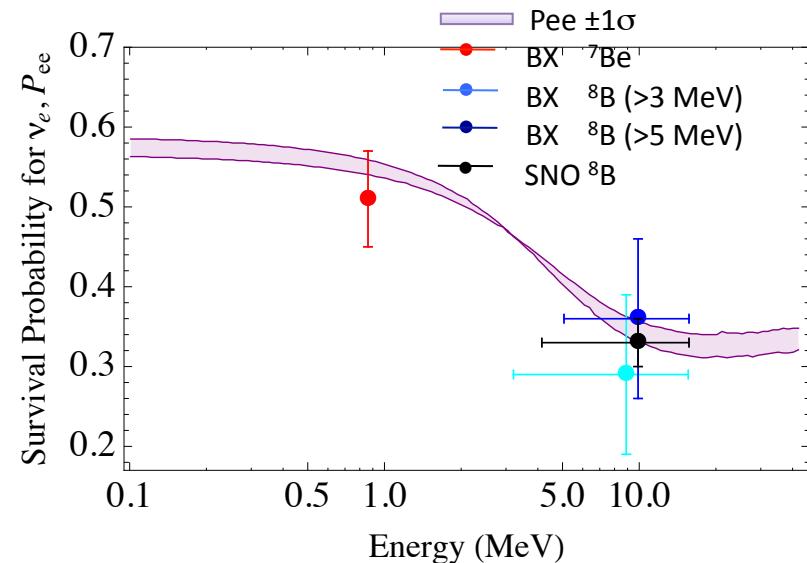
## The day night preliminary result

- The  ${}^7\text{Be}$  flux is obtained from the separated full fit of the day and night spectra
- Preliminary (and conservative) result:  $A_{DN} = 2 \frac{\Phi_n - \Phi_d}{\Phi_n + \Phi_d} = 0.007 \pm 0.073(\text{stat})$
- $A_{DN}$  is well consistent with zero: further confirmation of the LMA!
- Unique measurement for solar  ${}^7\text{Be}$  neutrinos
- Result not sensitive to many systematic effect influencing the  ${}^7\text{Be}$  absolute measurement
- ➡ More refined analysis is in progress aiming to reduce the error

# ${}^8\text{B} \nu$ with 3 MeV energy threshold in Borexino

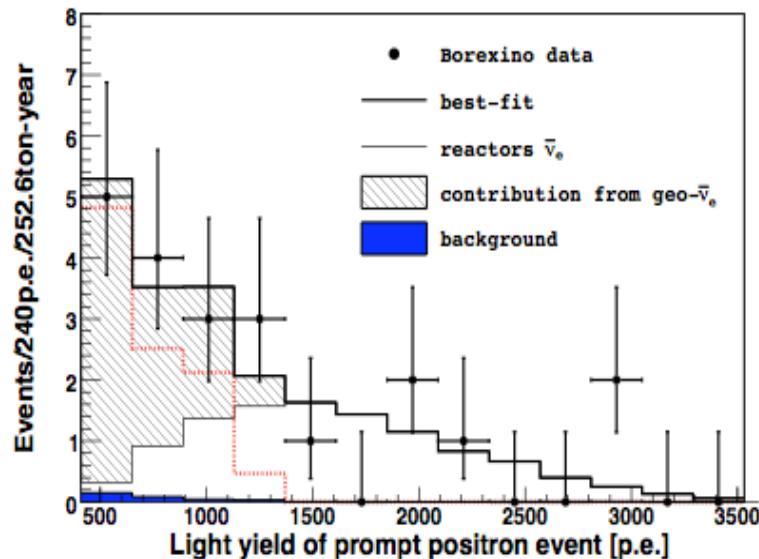


Probing for the first time with the same experiment the Pee in the vacuum regime ( ${}^7\text{Be}$  neutrinos) and in the matter-enhanced regime ( ${}^8\text{B}$  neutrinos);



## GeoNeutrino Results:

- 21 candidates selected
- exposure= 483 live days (252.6 ton-year after all cuts) December 07 - December 09



- Extract signal with an unbinned maximum likelihood fit using reference MonteCarlo shapes for both geo-neutrinos and reactor neutrinos, since small statistics;
- just the result is plot in a binned spectrum;
- result of the fit: amplitudes of the geo and reactor anti- $\nu$  spectra;

$$N_{geo} = 9.9^{+4.1 \text{ green} +14.6 \text{ magenta}}_{-3.4 \text{ green} -8.2 \text{ magenta}} \quad 68.3\% \text{ green} \quad 99.7\% \text{ magenta}$$

$$N_{react} = 10.7^{+4.3 \text{ green} +15.8 \text{ magenta}}_{-3.4 \text{ green} -8.0 \text{ magenta}} \quad 68.3\% \text{ green} \quad 99.7\% \text{ magenta}$$

- the first clear observation of geoneutrinos at  $4.2\sigma$  ;
- the rate is measured with 40% precision;
- confirmation/exclusion of geological models limited by the statistics;
- confirmation of oscillations (reactor antinu) at 1000 km @  $2.9\sigma$ ;
- georeactor in the Earth core with  $> 3$  TW rejected at 95% C.L.;



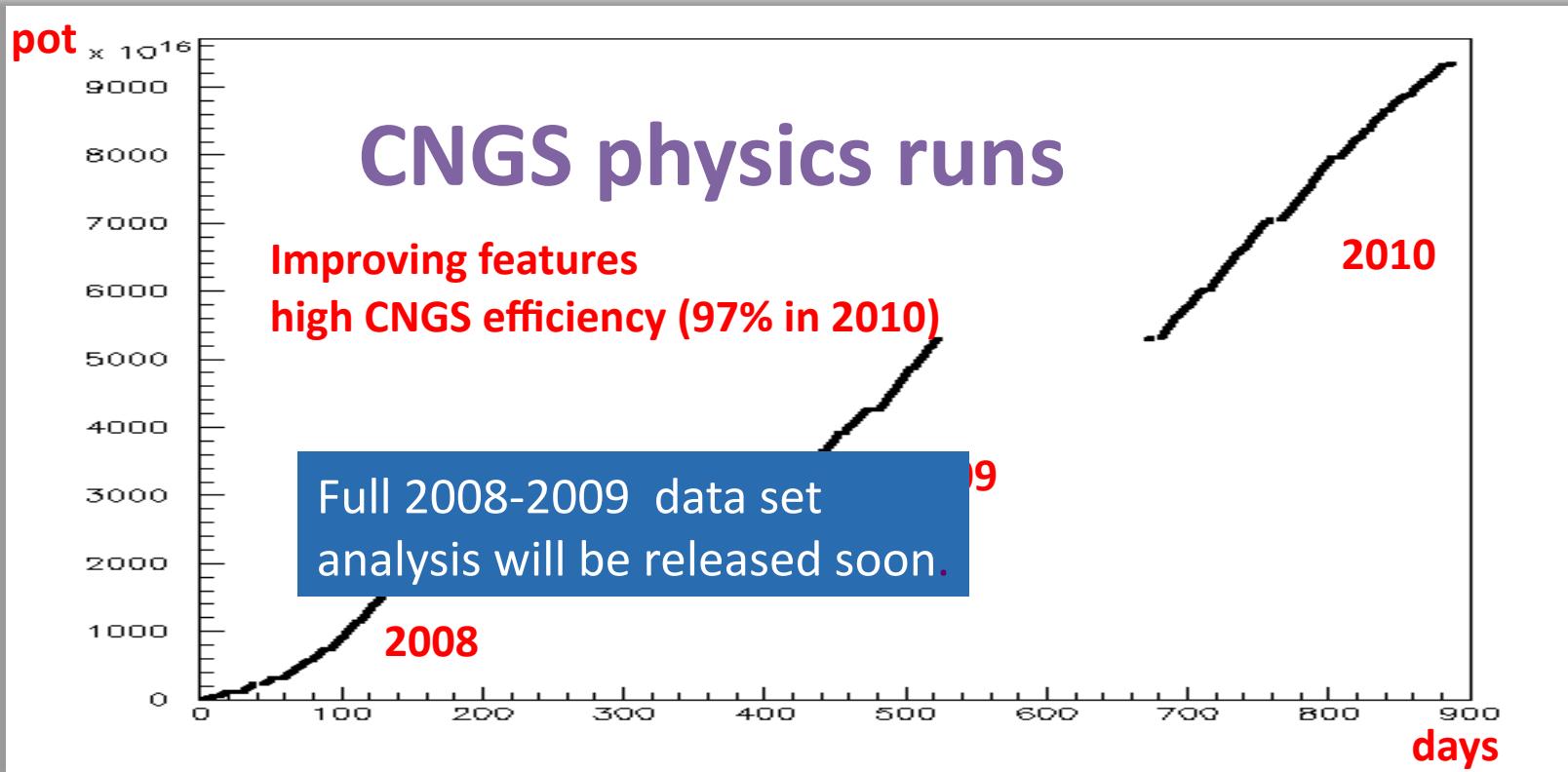
March 15, 2011

- The OPERA experiment
  - The physics case
  - Detector description
- Experimental results
  - Oscillation physics
    - First  $\nu_\tau$  candidate event [Physics Letters B 691 (2010) 138]
    - $\nu_\mu \leftrightarrow \nu_e$  ??
  - Non-Oscillation physics
    - Atmospheric muon charge ratio [EPJC 67 (2010) 25]
    - Atmospheric neutrinos??

## Summary & Outlook

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L Patrizii on behalf of the OPERA Collaboration



year	beam days	protons on target	SPS eff.	events in the bricks
2008	123	$1.78 \times 10^{19}$	61%	1698
2009	155	$3.52 \times 10^{19}$	70%	3693
2010	187	$4.04 \times 10^{19}$	81%	4248
TOTAL	465	$9.34 \times 10^{19}$	<71%>	9639

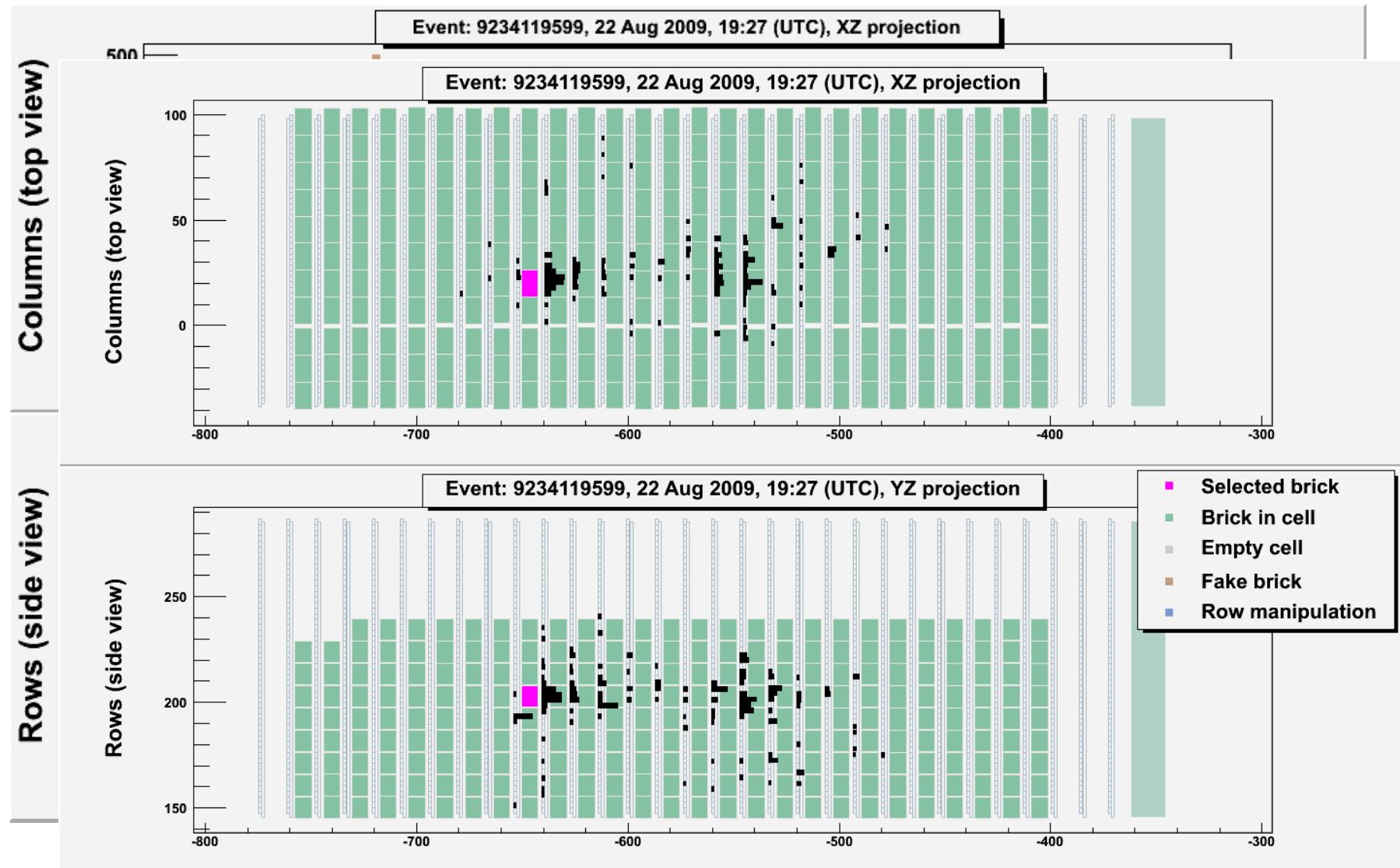
↳ 2.1 nominal CNGS years

NB. In what follows results refer to data released in physics publications **1088 (187 NC)**  
 $1.85 \times 10^{19}$  p.o.t. , 35% of 2008-'09 statistics, 20% of the total)

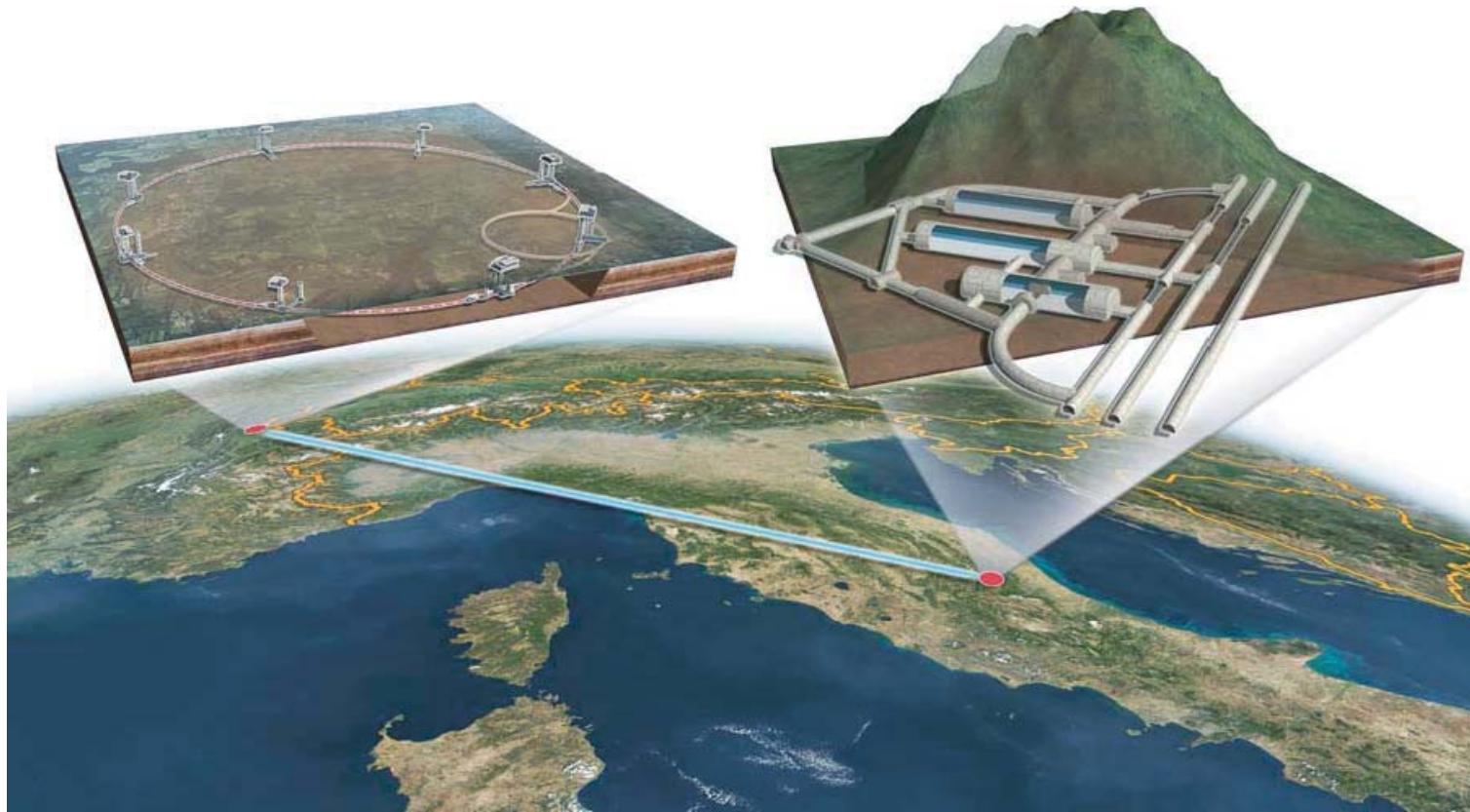
With that limited statistics, for  $\Delta m_{23}^2 = 2.5 \times 10^{-3} \text{ eV}^2$  and full mixing OPERA expected  $\sim 0.5 \nu_\tau$  events

# The first $\nu_\tau$ candidate event

Phys. Lett. B 691 (2010) 138

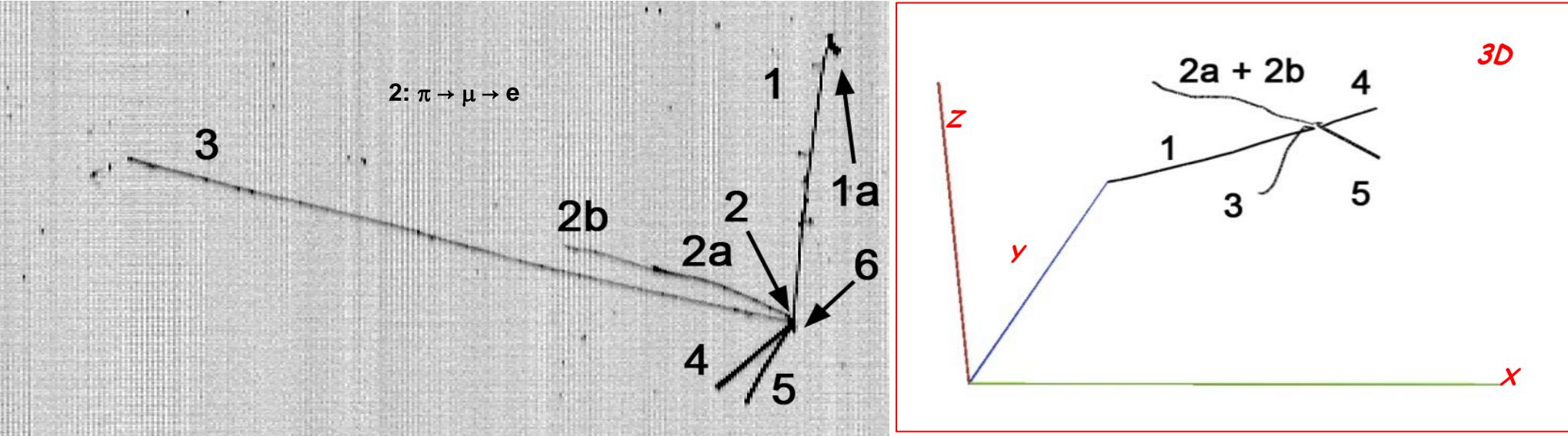


# ICARUS and Status of Liquid Argon Technology



F. Pietropaolo (INFN-PD)  
for the ICARUS Collaboration

# Run 9392 Event 106



Track	$E_k$ [MeV]	Range [cm]
1 (prob. $\pi$ , decays in flight)	136.1	55.77
2 ( $\pi$ )	26	3.3
2a ( $\mu$ )	79.1	17.8
2b ( $e$ )	24.1	10.4
3 ( $\mu$ )	231.6	99.1
4 (p)	168	19.2
5 (p)	152	16.3
6 (?) (merged with vtx)		2.9

- ⌚ Total deposited energy: 887 MeV
- ⌚ Total reconstructed momentum: 929 MeV/c at about 35° away from the CNGS beam direction

# Preliminary results of first CNGS 2010 run

- Analyzed sample: 1494 CNGS triggers, i.e.  $4.54 \cdot 10^{18}$  pot = 78 % out of whole sample. Classified by visual scanning into fiducial volume of 434 t.
- Number of collected interactions compared with number of interactions predicted ( $(2.6 v_{CC} + 0.86 v_{NC}) 10^{-17}/\text{pot}$ ), in the whole energy range up to 100 GeV, corrected by fiducial volume and DAQ dead-time.

Event type	Collected	Expected
$v_\mu CC$	94	98
$v NC$	32	31
$v XC^*$	6	-
Total	132	129

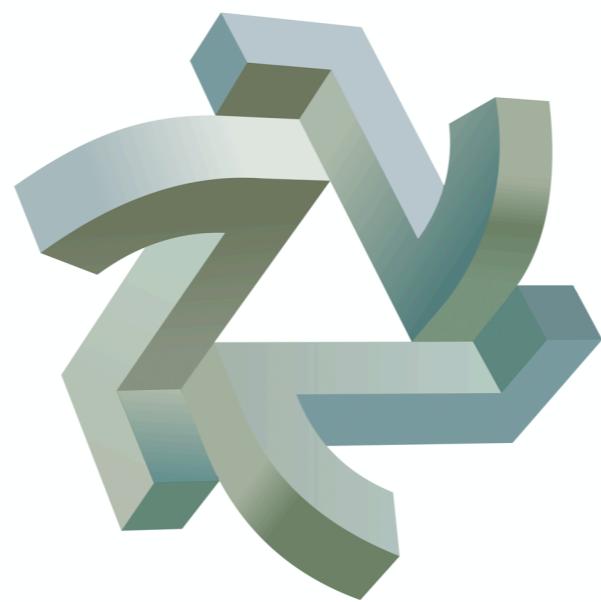
\* Events at edges, with  $\mu$  track too short to be visually recognized: further analysis needed.

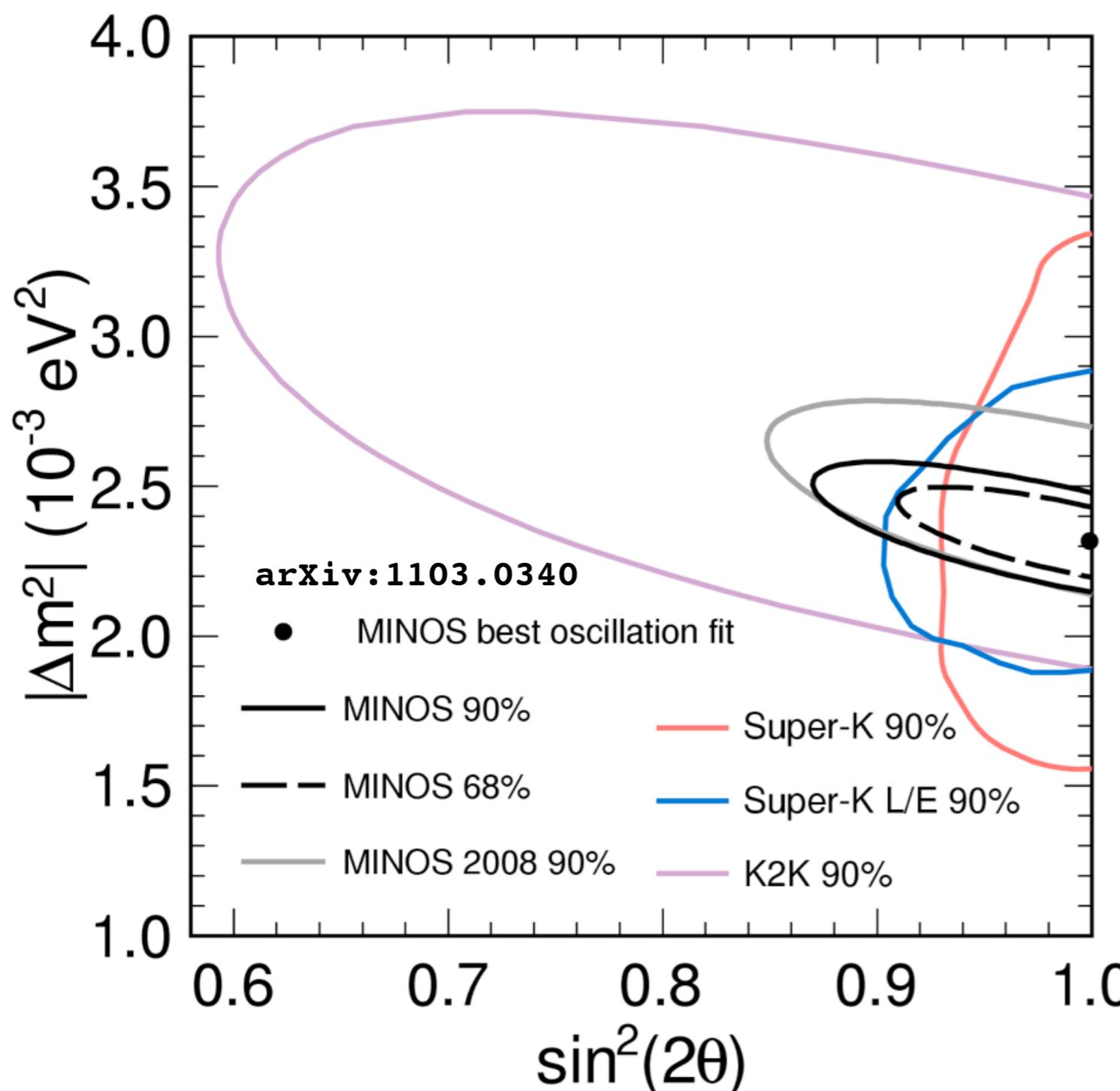
On overall statistics **in agreement with expectations**.

# MINOS

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Luke A. Corwin, for MINOS Collaboration  
Indiana University  
XIV International Workshop On Neutrino Telescopes  
2011 March 15

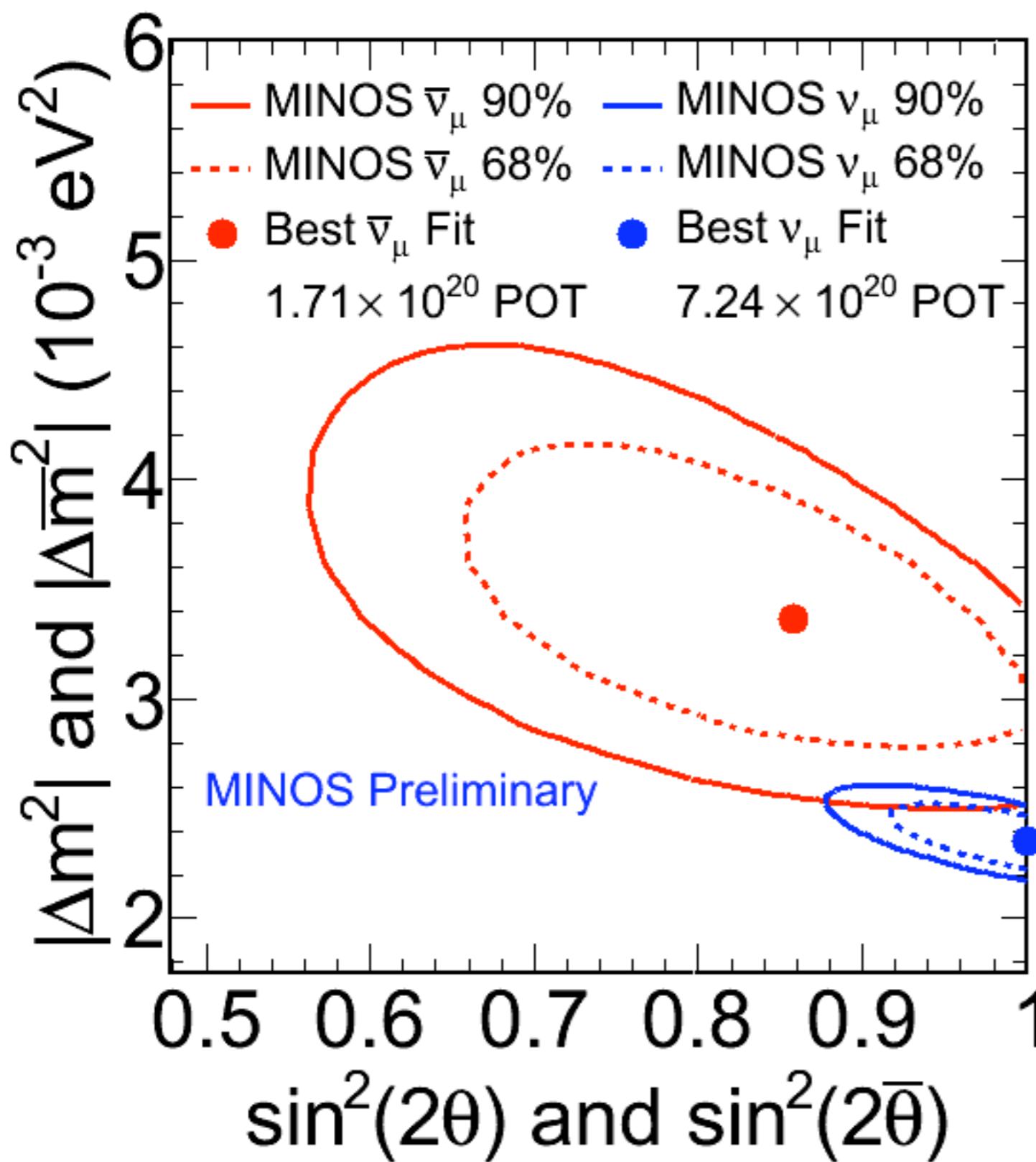




$$|\Delta m^2| = 2.32^{+0.12}_{-0.08} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta) > 0.90 \text{ (90\% C.L.)}$$

- Pure decoherence disfavored at  $9\sigma$
- Pure decay at  $7\sigma$
- World's most precise  $|\Delta m^2|$  measurement
- Included Samples
  - Fiducial Events
  - Events outside fiducial volume
  - Muons from neutrino events in rock



- Interesting Tension (2.3 $\sigma$  difference)
- Plan to have at least double current data set by this Summer.
  - Data taking interrupted by target failure on Feb. 26
  - Plan to have new target in April

# Results from the MiniBooNE Experiment

Geoffrey Mills

Los Alamos National Laboratory

For the MiniBooNE Collaboration

NeuTel2011

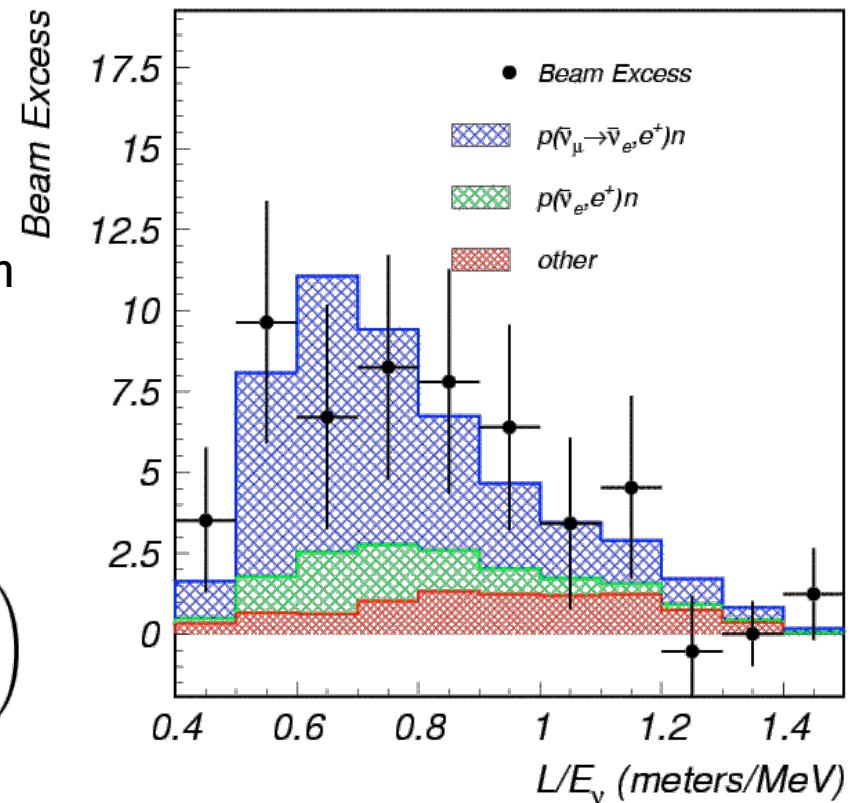
Venezia, Italia

Motivation....

## Excess Events from LSND still remain:

- LSND found an excess of  $\bar{\nu}_e$  in  $\bar{\nu}_\mu$  beam
- Signature: Cerenkov light from  $e^+$  with delayed n-capture (2.2 MeV)
- Excess:  $87.9 \pm 22.4 \pm 6.0$  (3.8s)
- The data was analysed under a two neutrino mixing hypothesis\**

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = \sin^2(2\theta) \sin^2 \left( \frac{1.27 L \Delta m^2}{E} \right)$$
$$= 0.245 \pm 0.067 \pm 0.045 \%$$

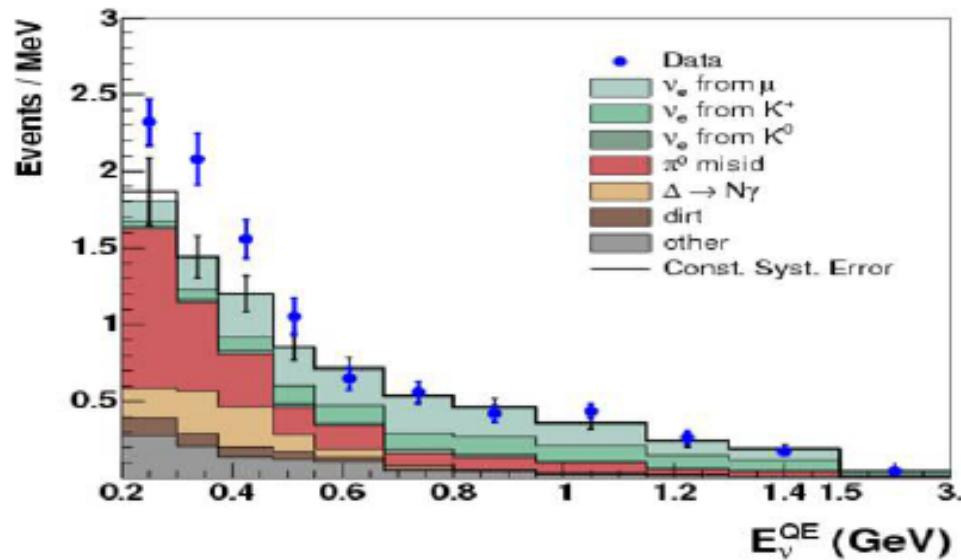


KARMEN at a distance of 17 meters saw no evidence for oscillations → low  $\Delta m^2$

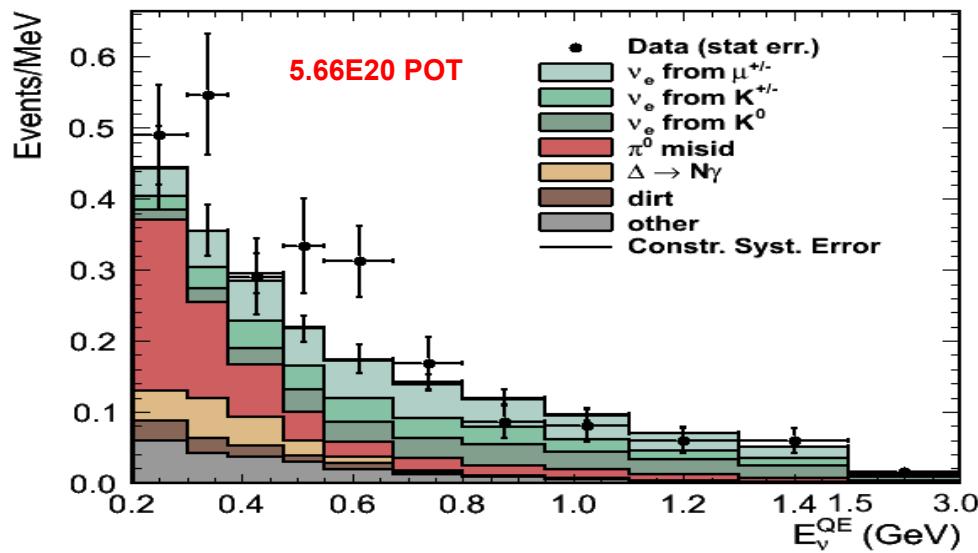
\*3 active + ≥2 sterile vs needed to fit all appearance and disappearance

# MiniBooNE $\nu_e$ and $\bar{\nu}_e$ Data

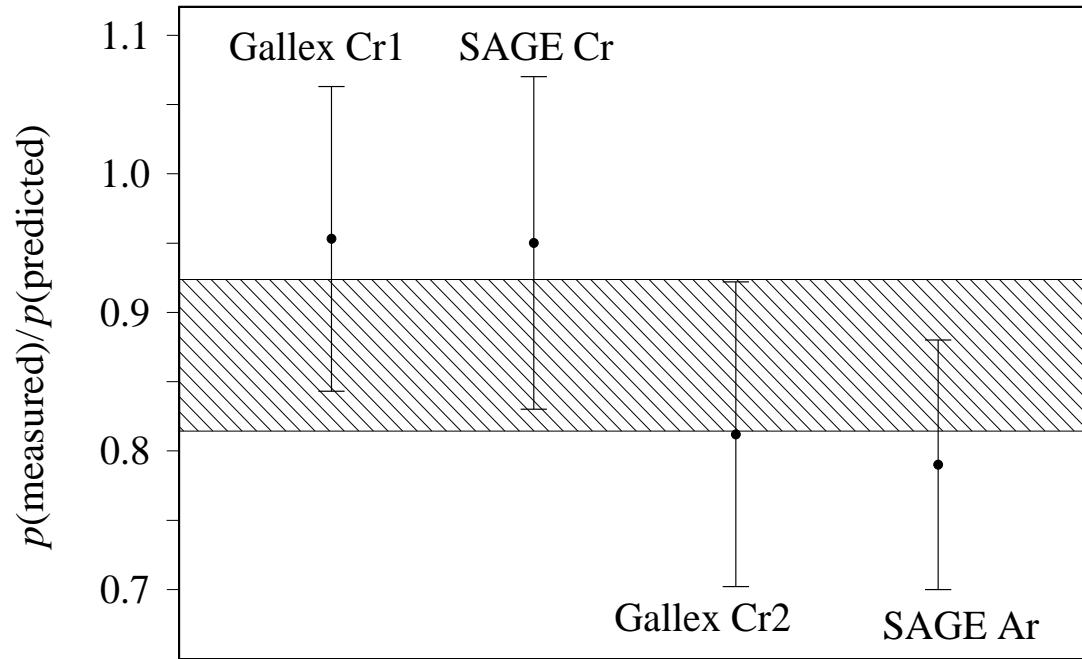
$\nu$  Mode



$\bar{\nu}$  Mode



## $\nu_e$ Disappearance in Gallium radioactive source experiments



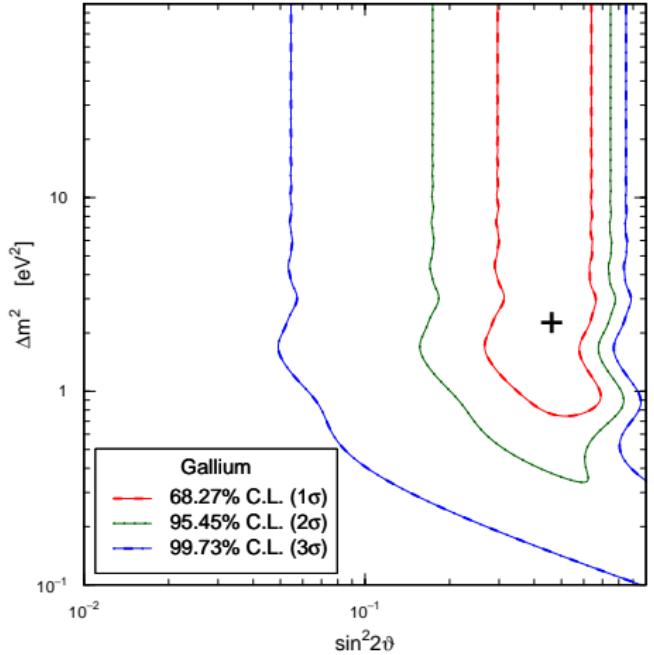
$R \equiv$  weighted average value of the ratio of measured and predicted  $^{71}Ge$  production rates (p) :

$$R \equiv \frac{p(\text{measured})}{p(\text{predicted})} = 0.87 \pm 0.05$$

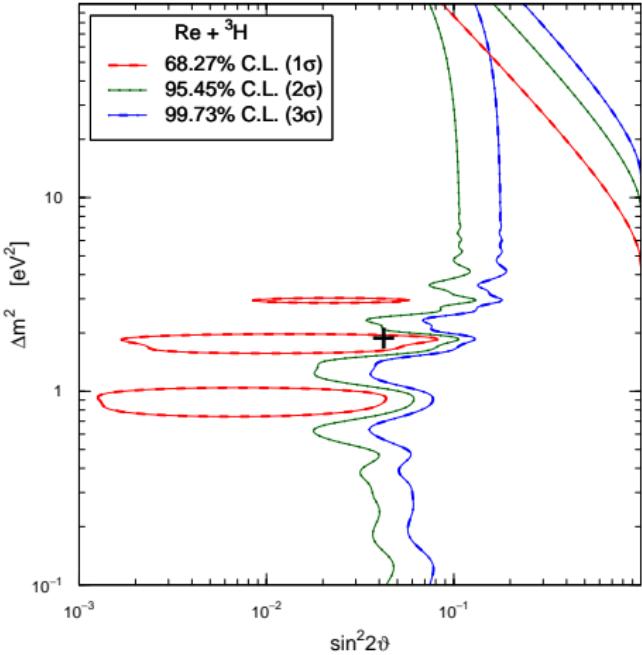
arXiv:0901.2200[nucl-ex]

Ga radioactive source exp. results may be interpreted as an indication of the disappearance of  $\nu_e$  due to active-sterile oscillations!

hep-ph/0610352 Carlo Giunti & ML



[Giunti, Laveder, arXiv:1006.3244]



[Giunti, Laveder, PRD 82 (2010) 053005, arXiv:1005.4599]

$$\Delta m_{SBL}^2 \gtrsim 1 \text{ eV}^2 \quad \text{is OK, but} \quad \sin^2 2\theta_\nu > \sin^2 2\theta_{\bar{\nu}}$$

Parameter Goodness of Fit = 0.2%

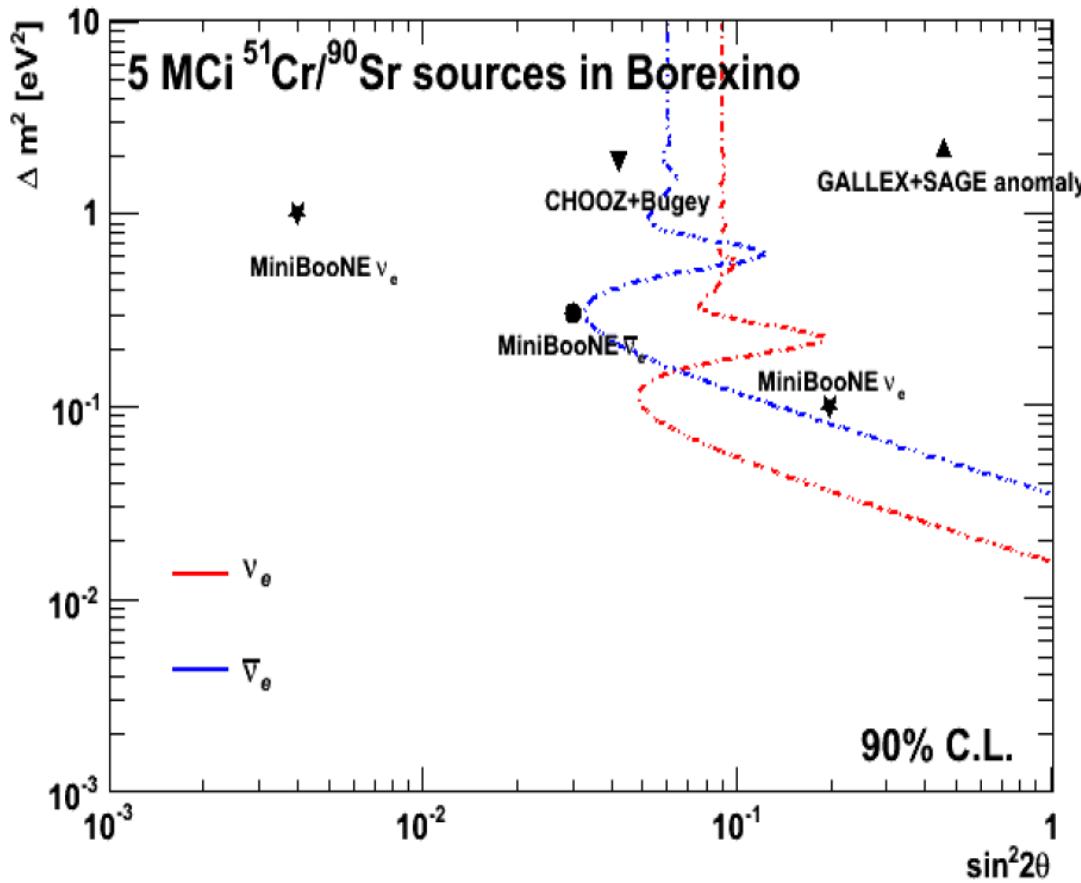
CPT violation?

[Giunti, Laveder, PRD 82 (2010) 113009, arXiv:1008.4750]

# Borexino test exp

► Borexino:

[Ianni, Montanino, Scioscia, EPJC 8 (1999) 609, arXiv:hep-ex/9901012]



[A. Ianni, Private Communication]

# Two-zone Ga source experiment

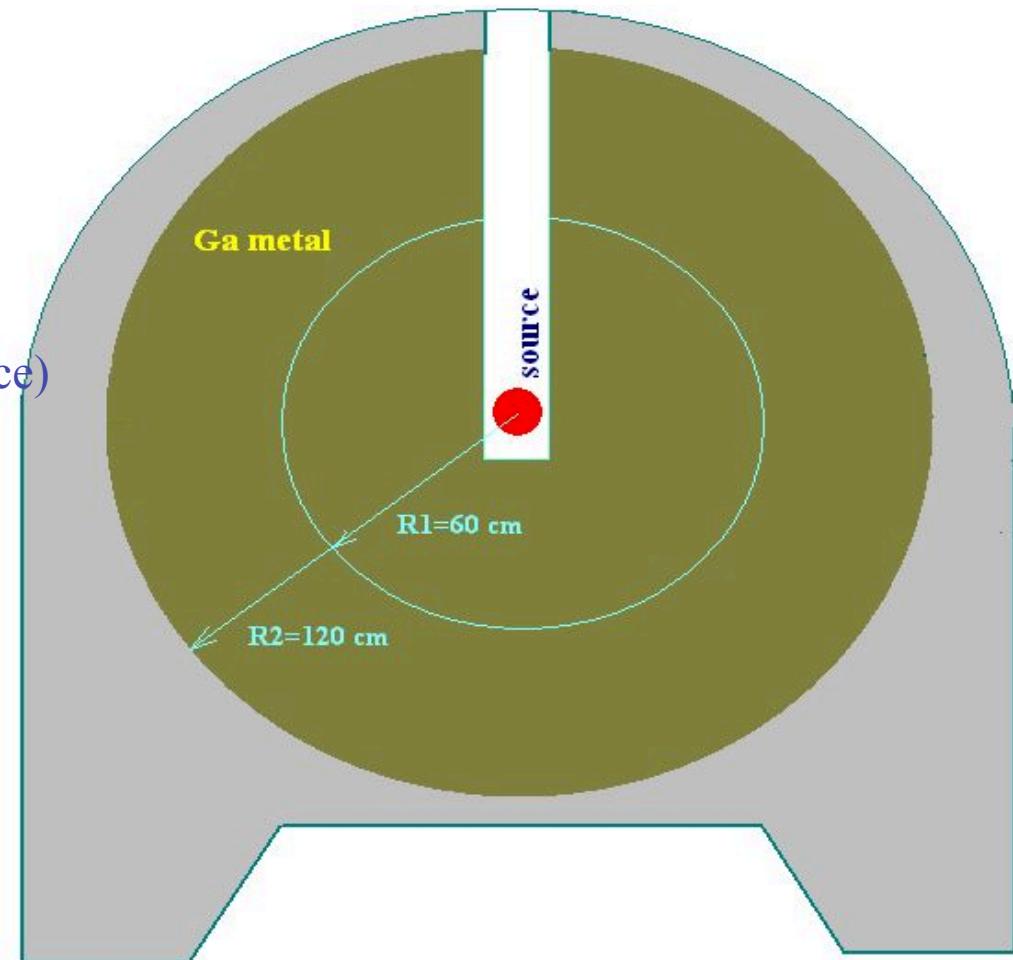
To get additional information in source experiment we separate the SAGE Ga target (50 tons) on two independent spherical zones.

It gives:

dependence on distance of source (test of  $\nu_e$  disappearance)

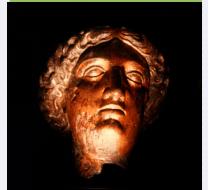
additional possibilities for statistical analysis

For two zones of target with thickness 60 cm each, the total uncertainty for each zone will be 5-5.5%, and statistical error of combined result will be about 3%



# MINERvA

Prospects and  
Status



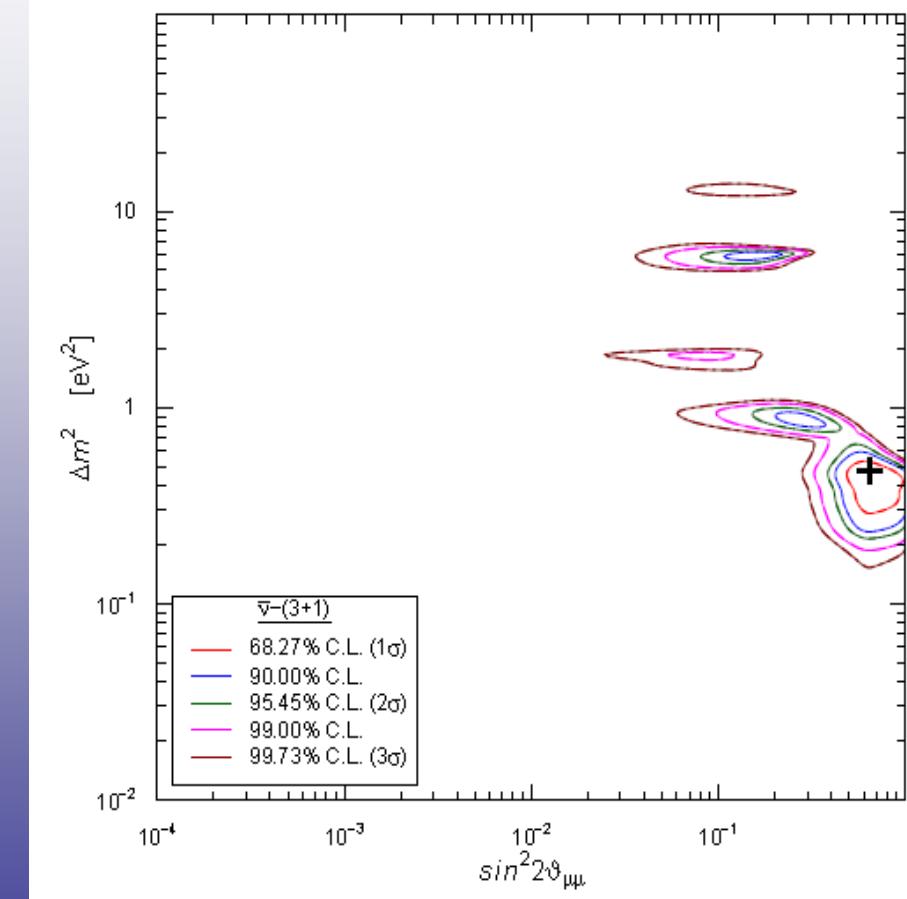
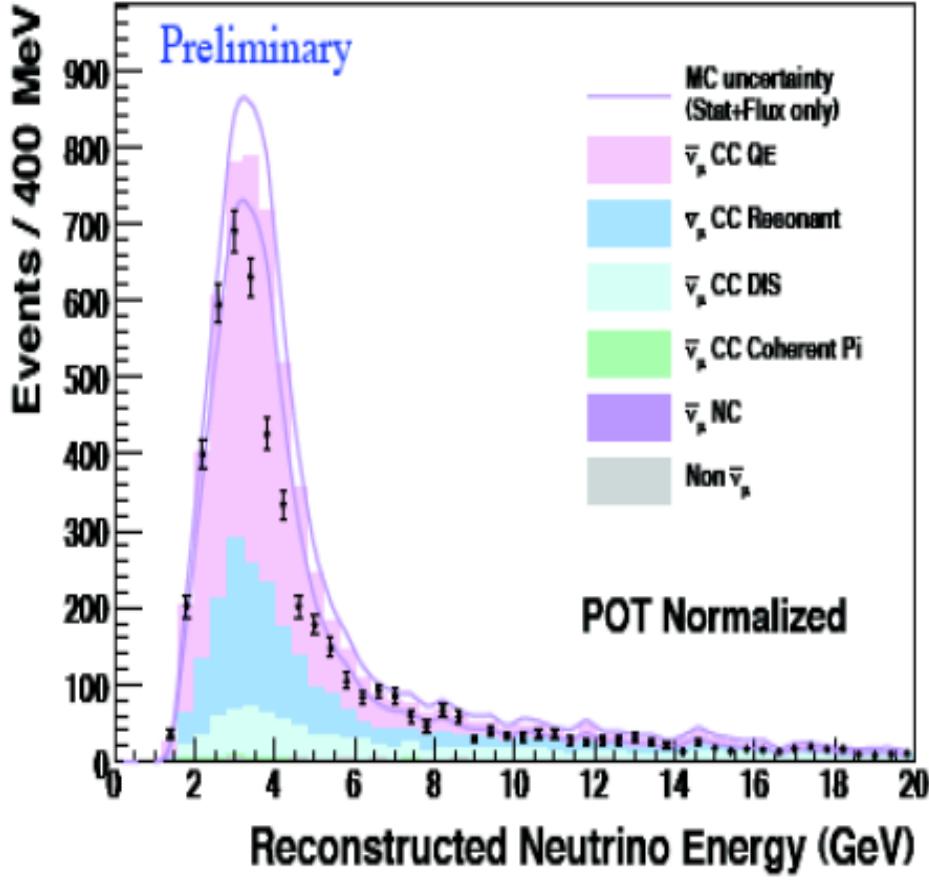
**NEUTEL 2011**  
**XIV International Workshop on**  
**Neutrino Telescopes**  
**15 - 18 March 2011**  
**Venice, Italy**



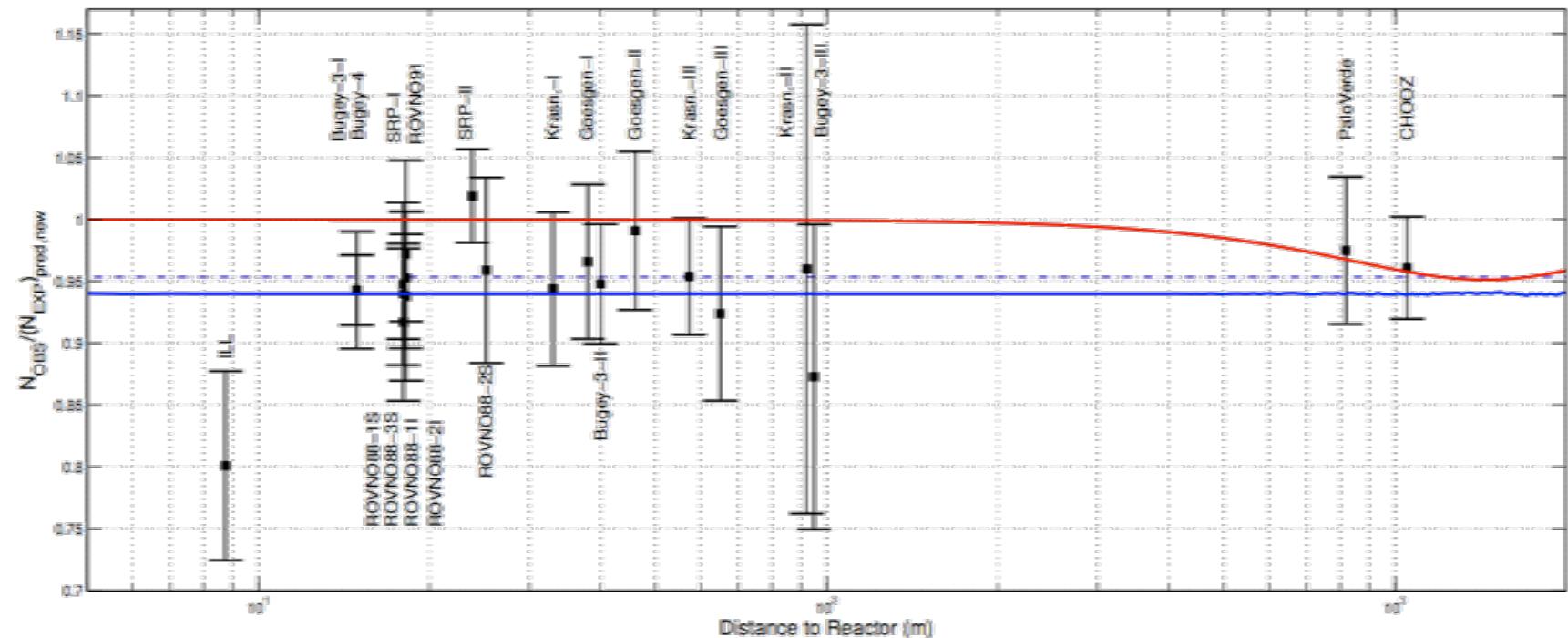
Vittorio Paolone  
University of Pittsburgh  
(Representing the MINERvA collaboration)



# Minerva anomaly - new



# The Reactor Antineutrino Anomaly and implications



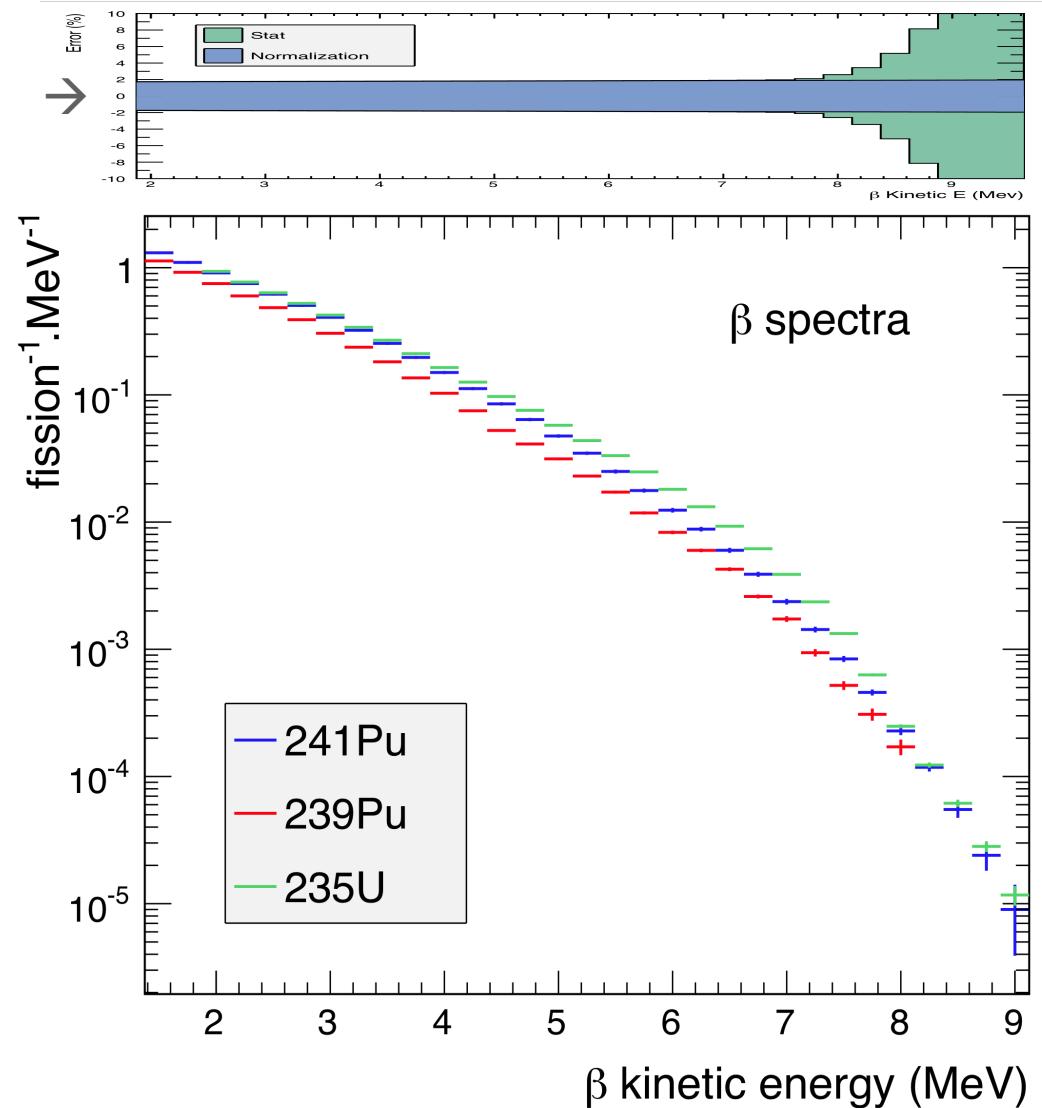
Th. Lasserre (CEA-Saclay, Irfu SPP & APC)

# The ILL electron Data Anchorage

Unique reference to be met by any other measurement or calculation

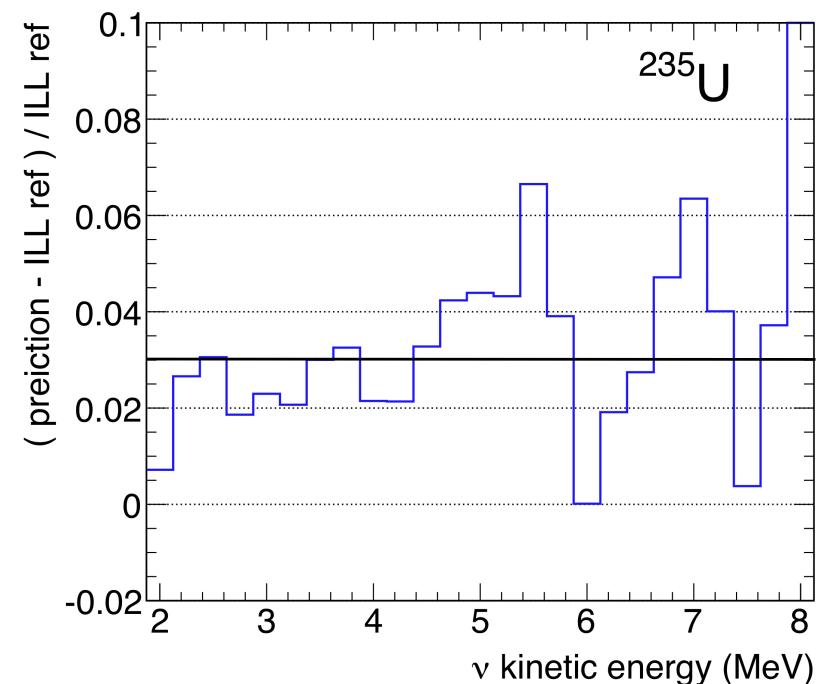
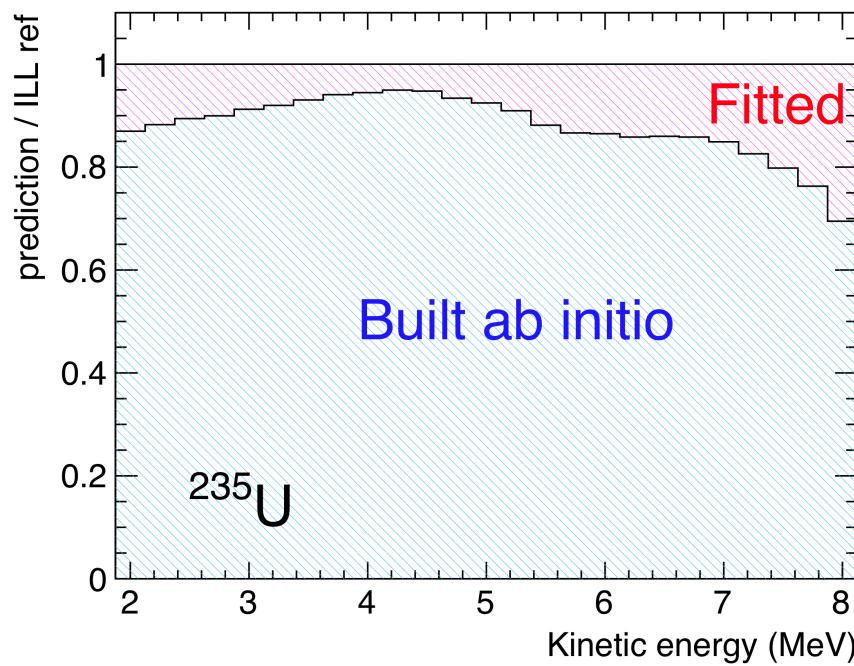
uncertainty

- Accurate  $e^-$  measurements  
@ ILL' (1980-89):
  - High resolution magn. spectrometer
  - Intense and pure thermal n spectrum from the core
  - Extensive use of reference internal conversion electron lines  
 $\rightarrow$  Normalization (1.8%)



# The New Mixed Conversion Approach

1. SAME ILL e- data Anchorage
2. Ab-Initio: “true” distribution of  $\beta$ -branches reproduces >90% of ILL e<sup>-</sup> data.
3. Old-procedure: five effective anchorage-branches to the remaining 10%.



- +3% normalization shift with respect to old  $v$  spectrum
- Similar result for all isotopes ( $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$ )
- Stringent Test Performed – Origin of the bias identified

# 19 Experimental Results Revisited (L<100m)

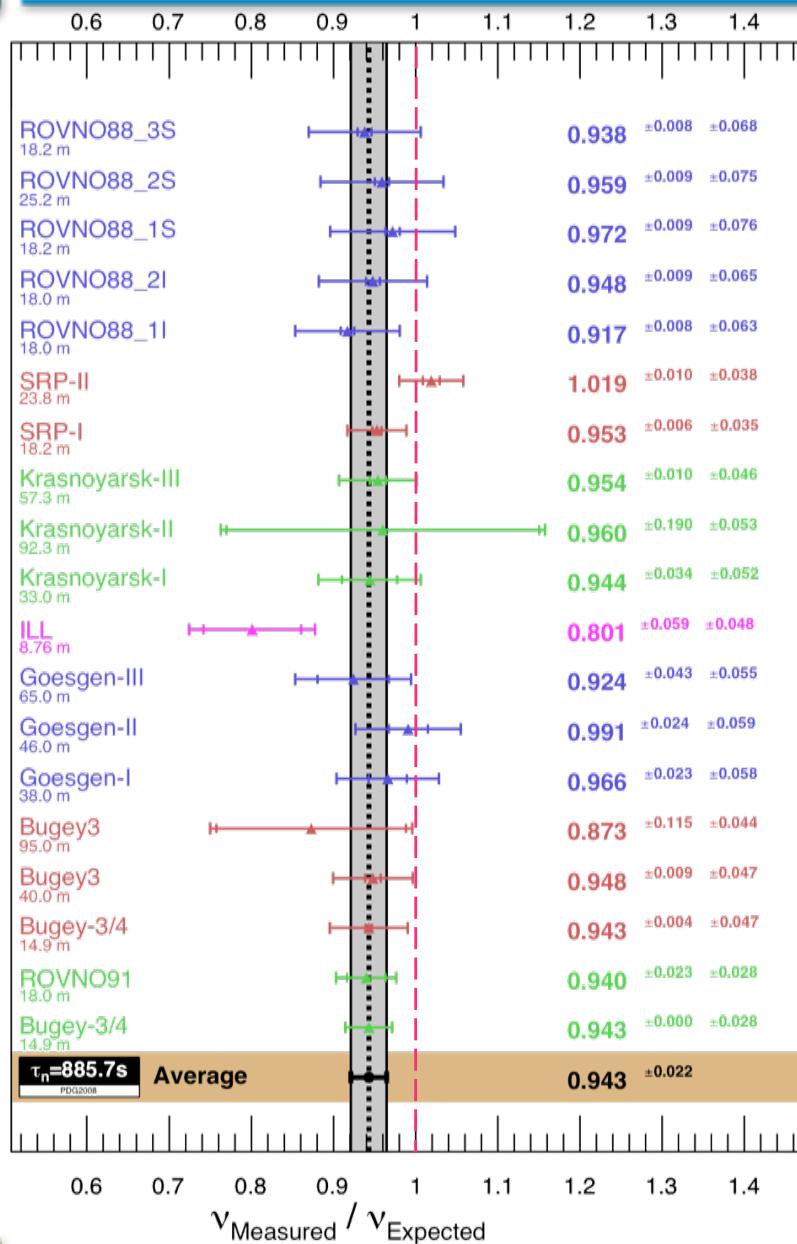
#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	$^3\text{He} + \text{H}_2\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.943	3.0	3.0	15
2	ROVNO91	$^3\text{He} + \text{H}_2\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
3	Bugey-3-I	$^6\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.988	0.943	5.0	5.0	15
4	Bugey-3-II	$\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.994	0.948	5.1	5.0	40
5	Bugey-3-III	$\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.915	0.873	14.1	5.0	95
6	Goesgen-I	$^3\text{He} + \text{LS}$	897	0.6198	0.274	0.074	0.042	1.018	0.966	6.5	6.0	38
7	Goesgen-II	$^3\text{He} + \text{LS}$	897	0.584	0.298	0.068	0.050	1.045	0.991	6.5	6.0	45
8	Goesgen-II	$^3\text{He} + \text{LS}$	897	0.543	0.329	0.070	0.058	0.975	0.924	7.6	6.0	65
9	ILL	$^3\text{He} + \text{LS}$	889	$\approx 1$	<0.01	<0.01	<0.01	0.832	0.801	9.5	6.0	9
10	Krasn. I	$^3\text{He} + \text{PE}$	899	$\approx 1$	<0.01	<0.01	<0.01	1.013	0.944	5.1	4.1	33
11	Krasn. II	$^3\text{He} + \text{PE}$	899	$\approx 1$	<0.01	<0.01	<0.01	1.031	0.960	20.3	4.1	92
12	Krasn. II	$^3\text{He} + \text{PE}$	899	$\approx 1$	<0.01	<0.01	<0.01	0.989	0.954	4.1	4.1	57
13	SRP I	Gd-LS	887	$\approx 1$	<0.01	<0.01	<0.01	0.987	0.953	3.7	3.7	18
14	SRP II	Gd-LS	887	$\approx 1$	<0.01	<0.01	<0.01	1.055	1.019	3.8	3.7	24
15	ROVNO88-1I	$^3\text{He} + \text{PE}$	898.8	0.607	0.277	0.074	0.042	0.969	0.917	6.9	6.9	18
16	ROVNO88-2I	$^3\text{He} + \text{PE}$	898.8	0.603	0.276	0.076	0.045	1.001	0.948	6.9	6.9	18
17	ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.8	18
18	ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.8	25
19	ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18

# 19 Experimental Results Revisited (L<100m)

OBSERVED/PREDICTED ratios: OLD & NEW (this work)

#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	$^3\text{He} + \text{H}_2\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.943	3.0	3.0	15
2	ROVNO91	$^3\text{He} + \text{H}_2\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
3	Bugey-3-I	$^6\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.988	0.943	5.0	5.0	15
4	Bugey-3-II	$\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.994	0.948	5.1	5.0	40
5	Bugey-3-III	$\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.915	0.873	14.1	5.0	95
6	Goesgen-I	$^3\text{He} + \text{LS}$	897	0.6198	0.274	0.074	0.042	1.018	0.966	6.5	6.0	38
7	Goesgen-II	$^3\text{He} + \text{LS}$	897	0.584	0.298	0.068	0.050	1.045	0.991	6.5	6.0	45
8	Goesgen-II	$^3\text{He} + \text{LS}$	897	0.543	0.329	0.070	0.058	0.975	0.924	7.6	6.0	65
9	ILL	$^3\text{He} + \text{LS}$	889	$\simeq 1$	<0.01	<0.01	<0.01	0.832	0.801	9.5	6.0	9
10	Krasn. I	$^3\text{He} + \text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	1.013	0.944	5.1	4.1	33
11	Krasn. II	$^3\text{He} + \text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	1.031	0.960	20.3	4.1	92
12	Krasn. II	$^3\text{He} + \text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	0.989	0.954	4.1	4.1	57
13	SRP I	Gd-LS	887	$\simeq 1$	<0.01	<0.01	<0.01	0.987	0.953	3.7	3.7	18
14	SRP II	Gd-LS	887	$\simeq 1$	<0.01	<0.01	<0.01	1.055	1.019	3.8	3.7	24
15	ROVNO88-1I	$^3\text{He} + \text{PE}$	898.8	0.607	0.277	0.074	0.042	0.969	0.917	6.9	6.9	18
16	ROVNO88-2I	$^3\text{He} + \text{PE}$	898.8	0.603	0.276	0.076	0.045	1.001	0.948	6.9	6.9	18
17	ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.8	18
18	ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.8	25
19	ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18

# The reactor antineutrino anomaly

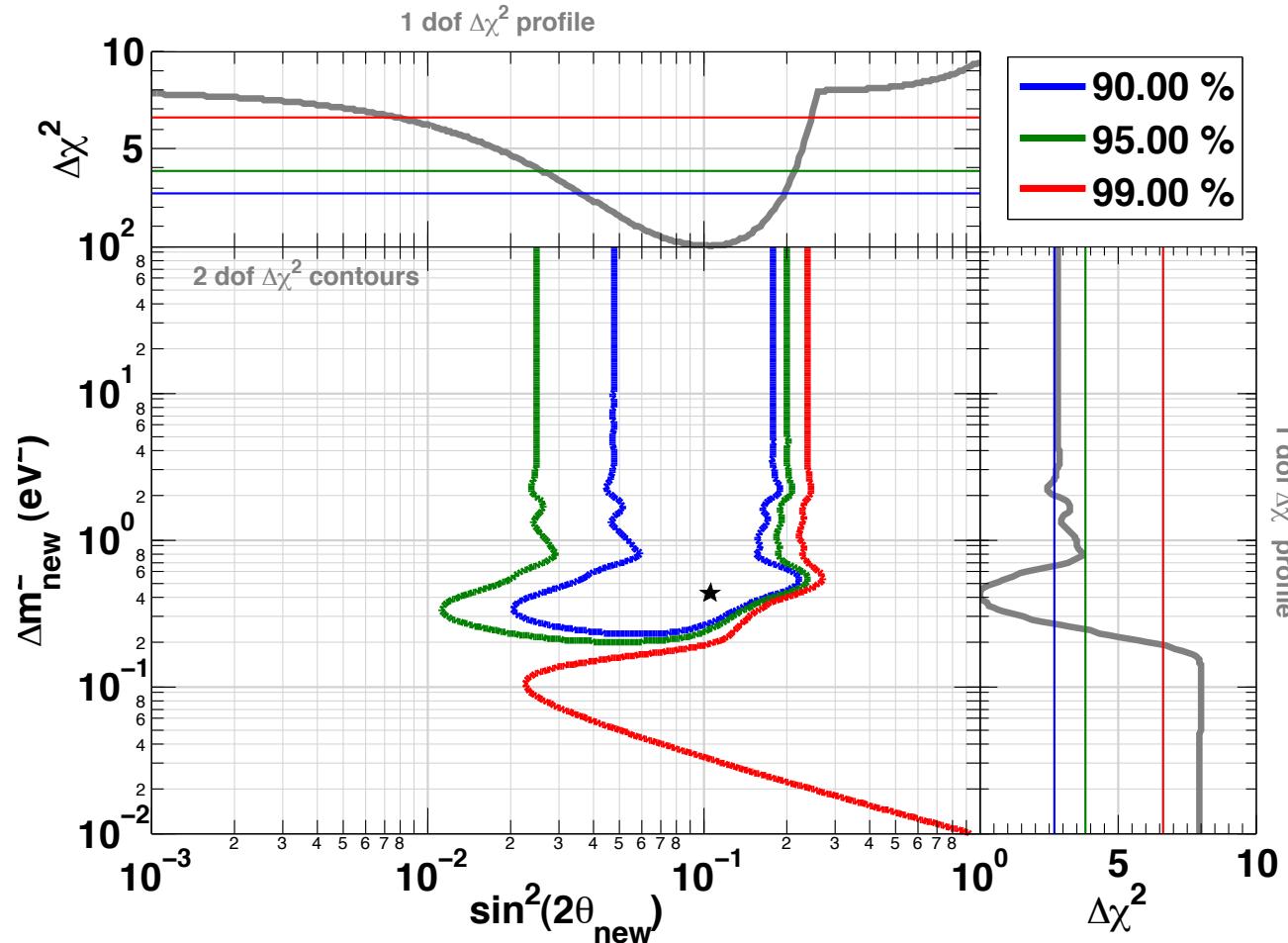


$$\chi^2 = \left( \mathbf{r} - \overrightarrow{\mathbf{R}} \right)^T W^{-1} \left( \mathbf{r} - \overrightarrow{\mathbf{R}} \right)$$

- Best fit :  $\mu = 0.943$
- Uncertainty : 0.023
- $\chi^2 = 19.6/19$
- Deviation from unity
- Naïve Gaussian : 99.3% C.L.
- Toy MC: 98.6% C.L. ( $10^6$  trials)
- No hidden covariance
- 18% of Toy MC have  $\chi^2_{\min} < 19.6$

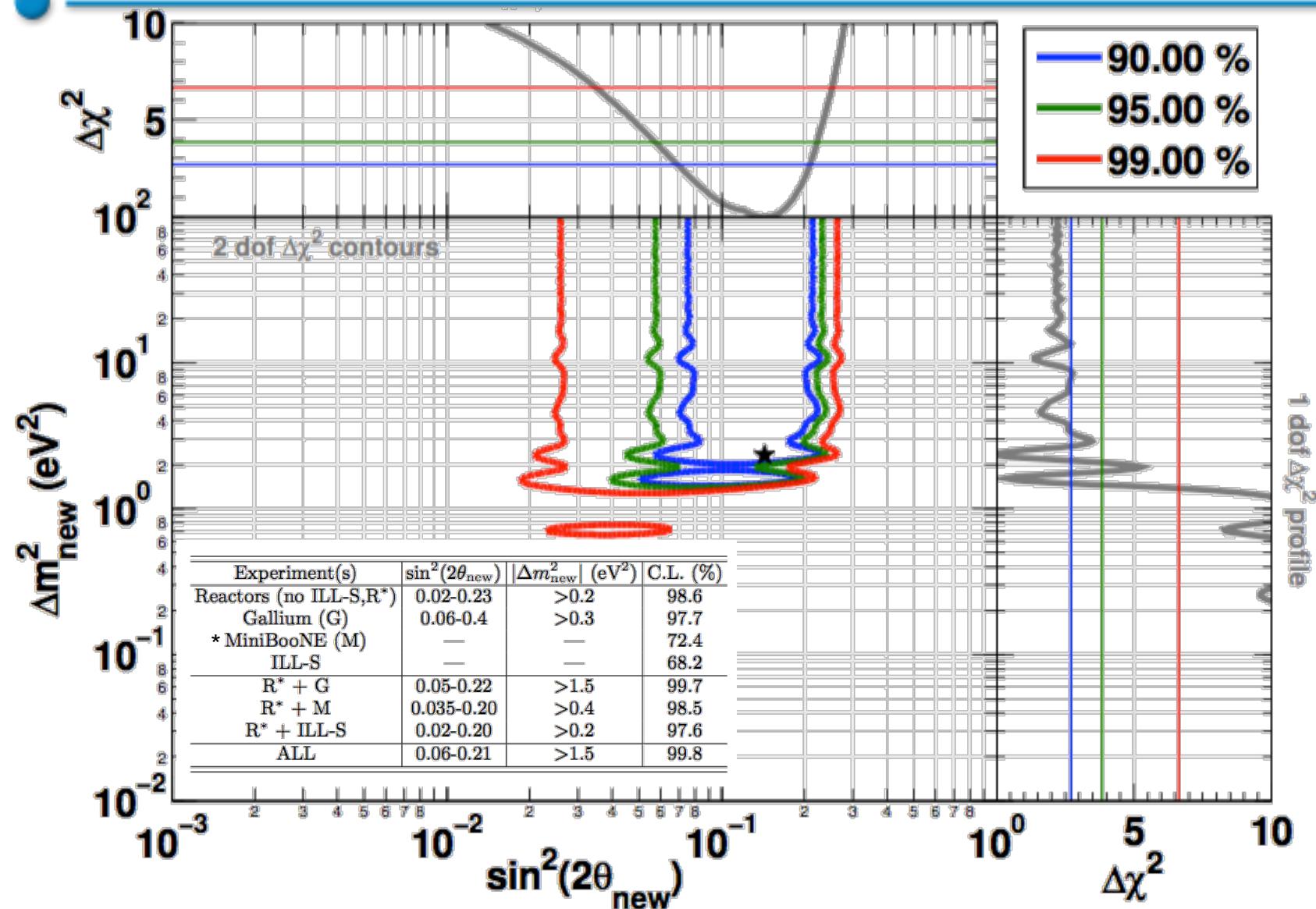
# The 4<sup>th</sup> neutrino hypothesis

- Combine all rate measurements, no spectral-shape information
- Fit to anti- $\nu_e$  disappearance hypothesis



- Absence of oscillations disfavored at 98.6% C.L.

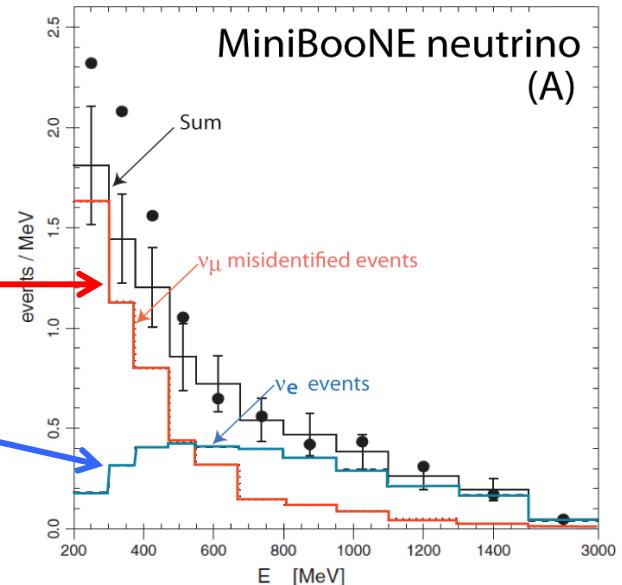
# Putting it all together: reactor rates + shape + Gallium + (MB)



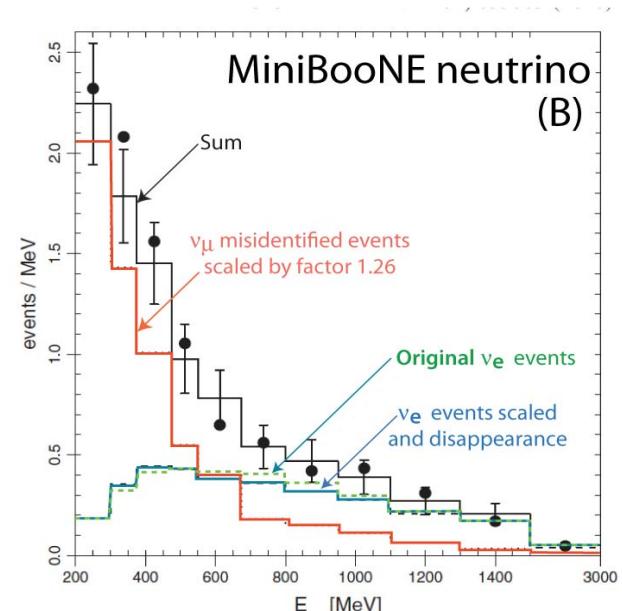
The no-oscillation hypothesis is disfavored at 99.8% CL

# The neutrino run

- (A) electron-like neutrino data. Comparison between the data (black dots) and the calculated distributions due to misidentified  $\nu_\mu$  events (red) and genuine  $\nu_e$  events (blue). The sum is indicated in black. One notices an anomaly at low energies, which is incompatible with LNSD predictions.

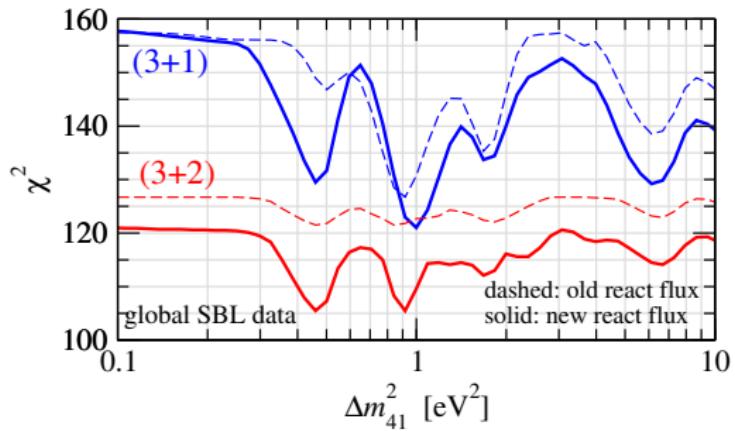


- (B) according to Giunti & Laveder scaling of the events is applied with a factor 1.26, within the permitted uncertainty of  $F = 1.24 \pm 0.21$  and gives an acceptable fit to the data. The  $\nu_e$  with and without scaling and disappearance are also shown.



# 3+2 global fit

$\Delta m_{41}^2$	$ U_{e4} $	$ U_{\mu 4} $	$\Delta m_{51}^2$	$ U_{e5} $	$ U_{\mu 5} $	$\delta/\pi$	$\chi^2/\text{dof}$
0.47	0.131	0.170	0.93	0.135	0.142	1.62	105.9/130



- $\Delta\chi^2$  (old vs new fluxes) = 15.5
- $\Delta\chi^2$  (3+1 vs 3+2) = 14.1 (99.3% CL, 4 dof)

# Intermezzo sulla fisica di Majorana



# Natura del neutrino



Dirac



Majorana

# *Nuovo Cimento* 14 (1937) 171-184

## TEORIA SIMMETRICA DELL'ELETTRONE E DEL POSITRONE

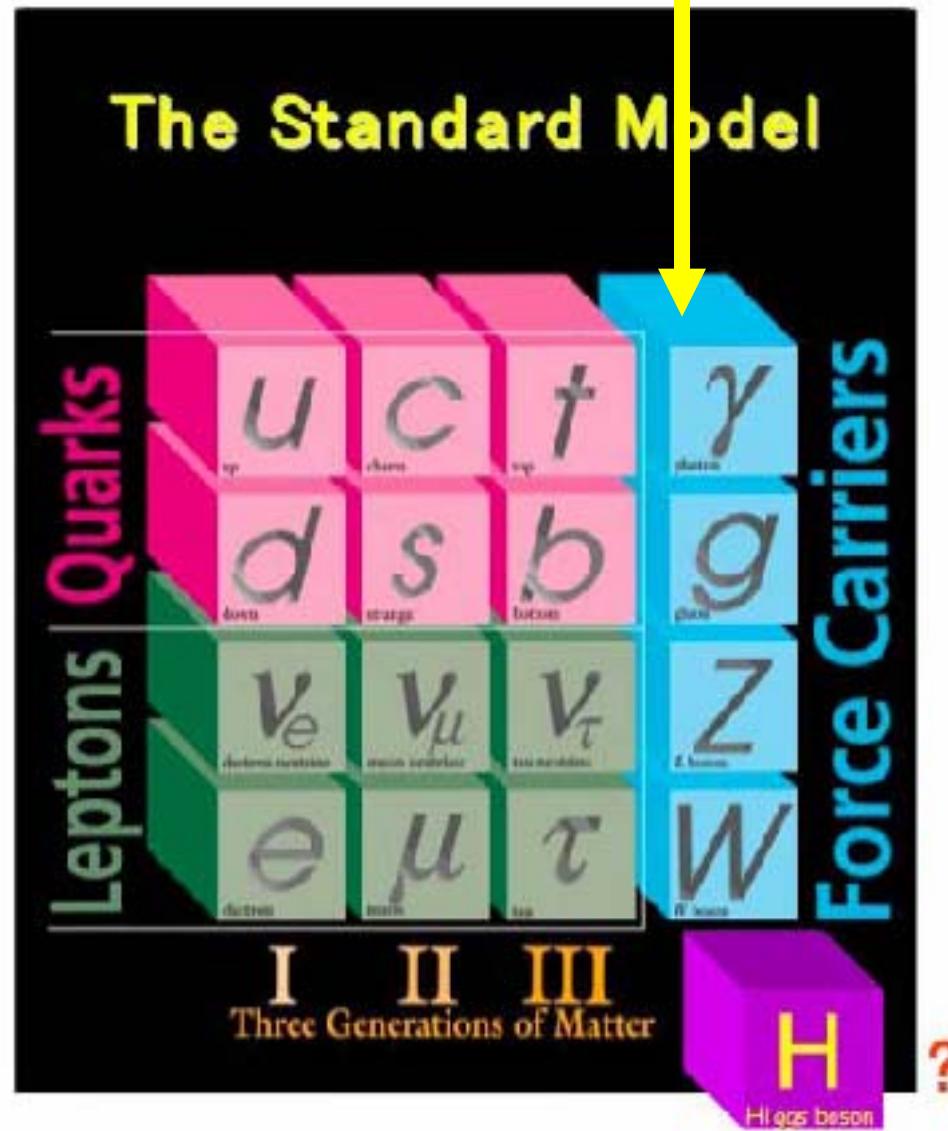
Nota di ETTORE MAJORANA

**Sunto.** - *Si dimostra la possibilità di pervenire a una piena simmetrizzazione formale della teoria quantistica dell'elettrone e del positrone facendo uso di un nuovo processo di quantizzazione. Il significato delle equazioni di DIRAC ne risulta alquanto modificato e non vi è più luogo a parlare di stati di energia negativa; nè a presumere per ogni altro tipo di particelle, particolarmente neutre, l'esistenza di « antiparticelle » corrispondenti ai « vuoti » di energia negativa.*

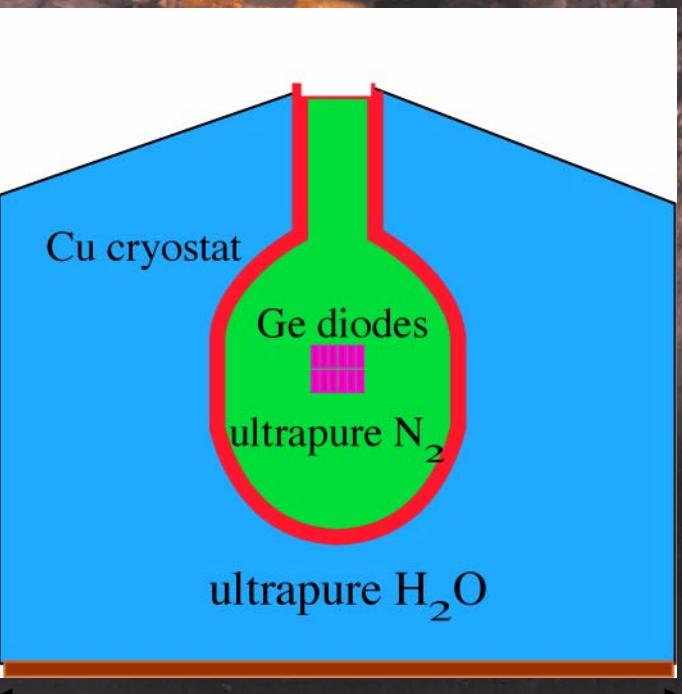
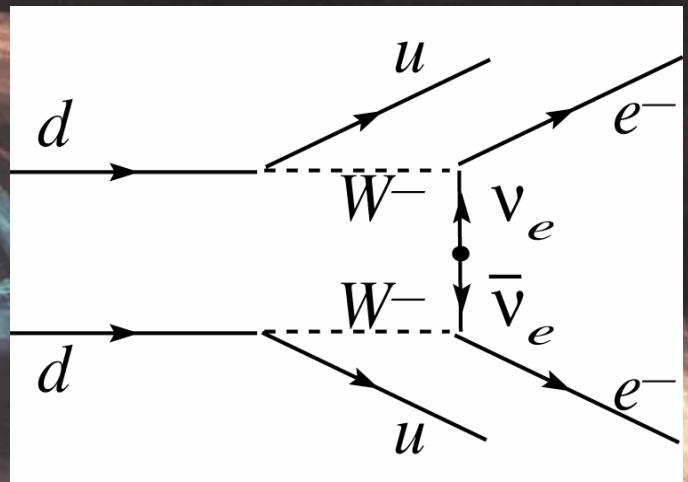
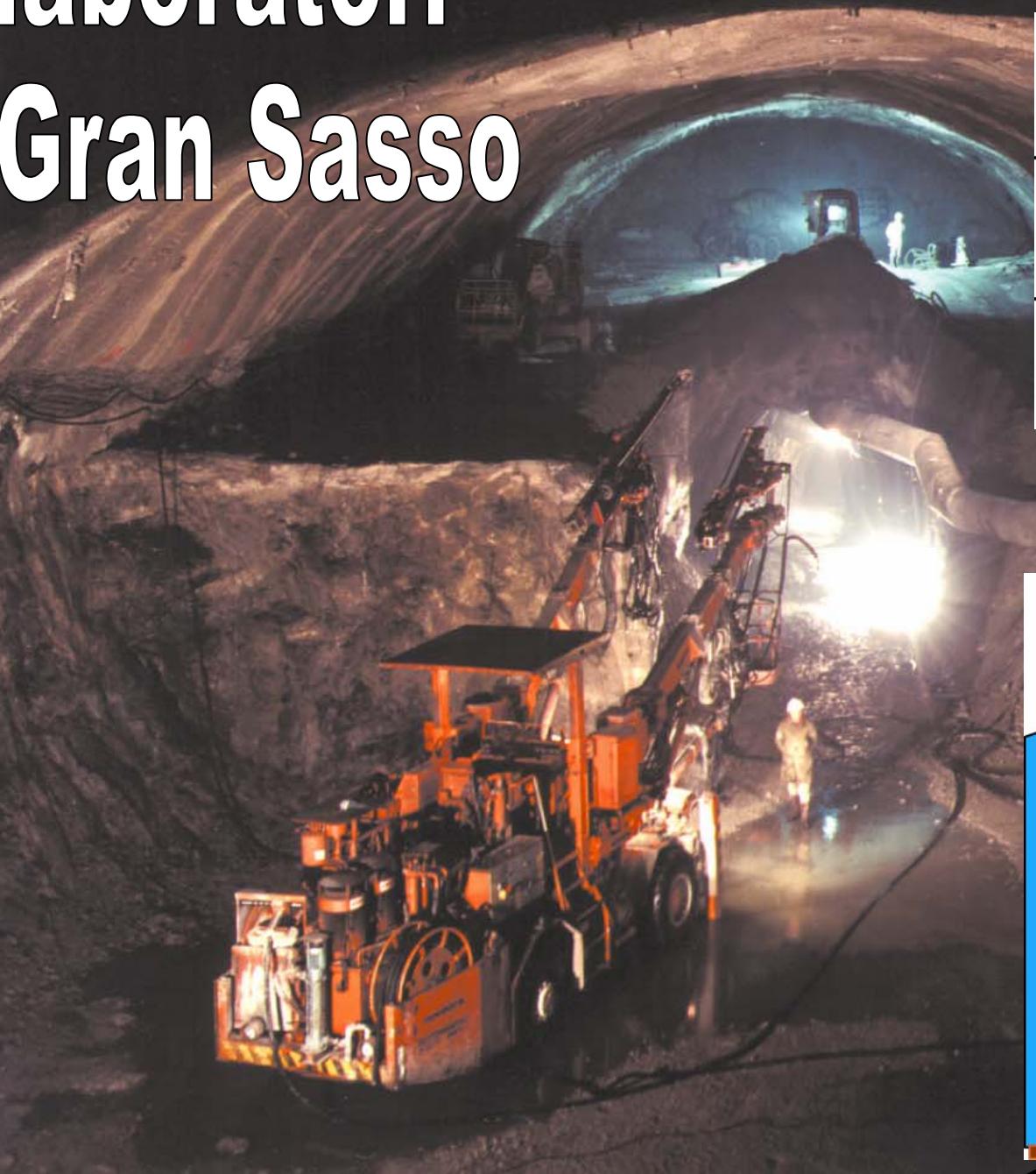
*We show that it is possible to achieve complete formal symmetrization in the electron and positron quantum theory by means of a new quantization process. The meaning of Dirac equations is somewhat modified and it is no more necessary to speak of negative-energy states; nor to assume, for any other type of particles, especially neutral ones, the existence of antiparticles, corresponding to the “holes” of negative energy.*

# Majorana Neutrino

Gravity  
?



# Laboratori Gran Sasso

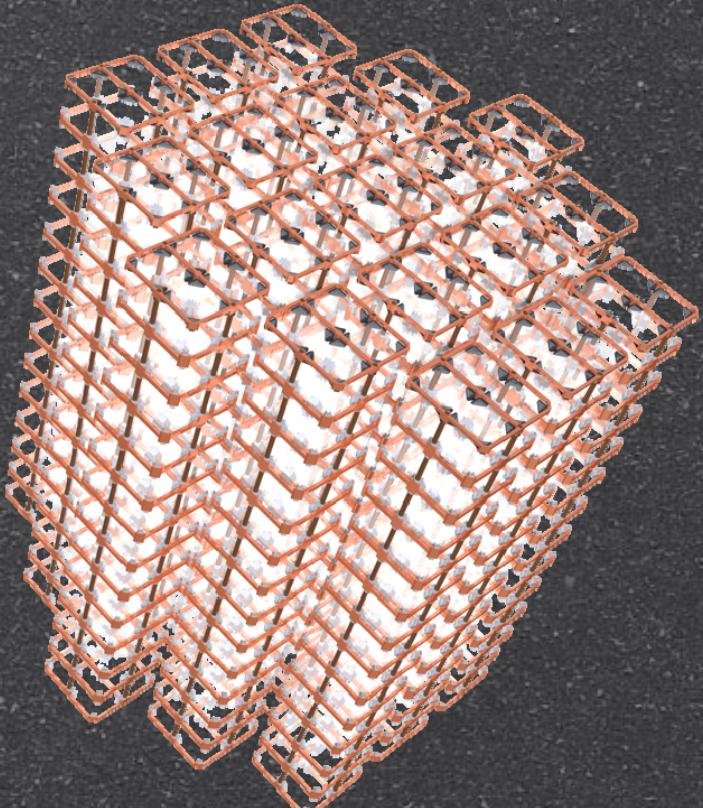


XIV International Workshop on "Neutrino Telescopes"

Venice, March 15-18, 2011



# CUORICINO, CUORE-0 AND CUORE: AN UPDATE



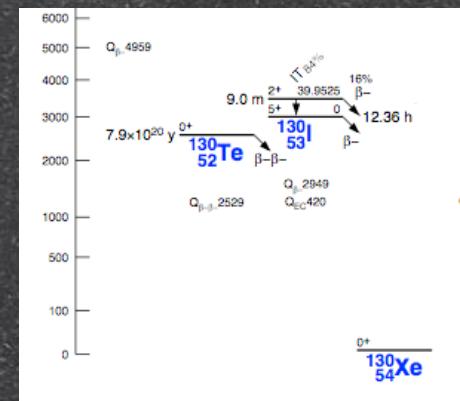
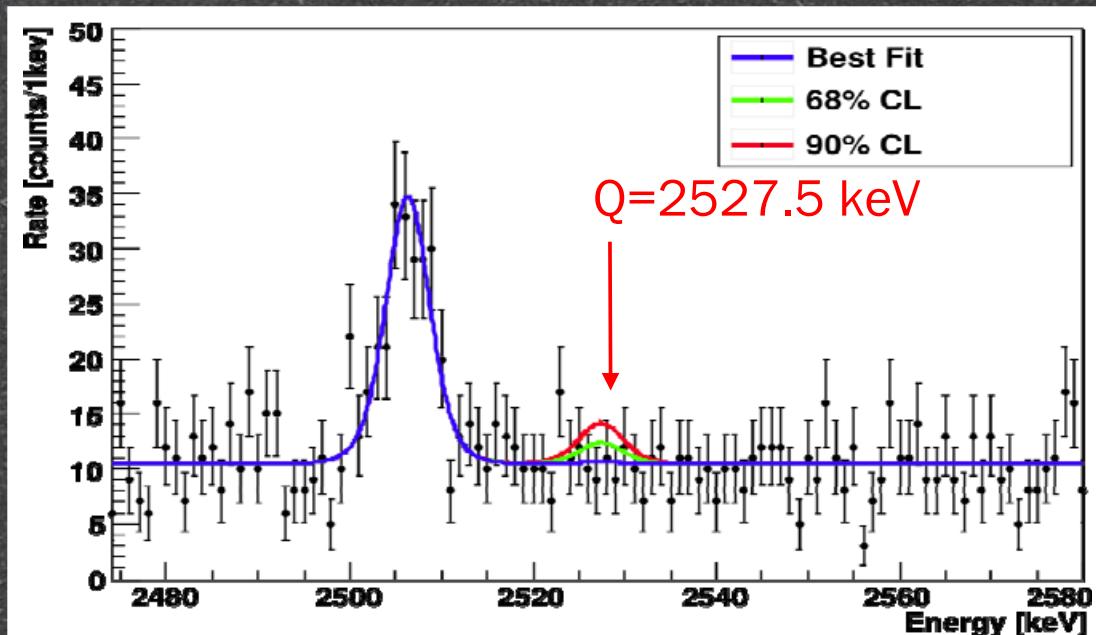
C. Brofferio, University of Milano Bicocca  
on behalf of the CUORE Collaboration



# CUORICINO: $0\nu\text{DBD}$ RESULT

Astropart. Phys. (2011), doi:10.1016/j.astropartphys.2011.02.002

TOTAL: **19.6 kg · yr**  $^{130}\text{Te}$  exposure  
 collected in 2 runs (2003-2004, 2004-2008)  
 (II Run, Big Crystals alone: **15.8 kg · y**)



## NME bibliography:

- 1 Šimkovic et al.,  
PRC 77 (2008) 045503
- 2 Civitarese et al.,  
JoP:Conference series  
173 (2009) 012012
- 3 Menéndez et al.,  
NPA 818 (2009) 139
- 4 Barea and Iachello,  
PRC 79 (2009) 044301

Background Big Crystals, II run:

$0.153 \pm 0.006$  counts/keV/kg/y

Lower limit, half-life:

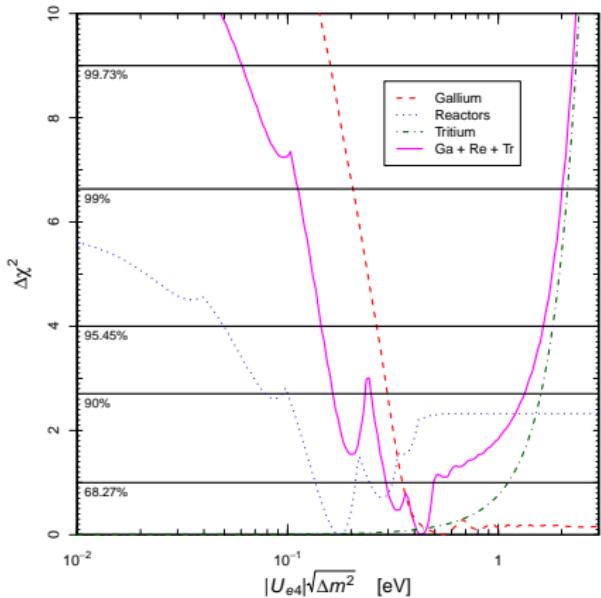
$T_{1/2}^{0\nu} ({}^{130}\text{Te}) > 2.8 \times 10^{24} \text{ y}$  (90% C.L.)

Upper limit, Majorana mass:

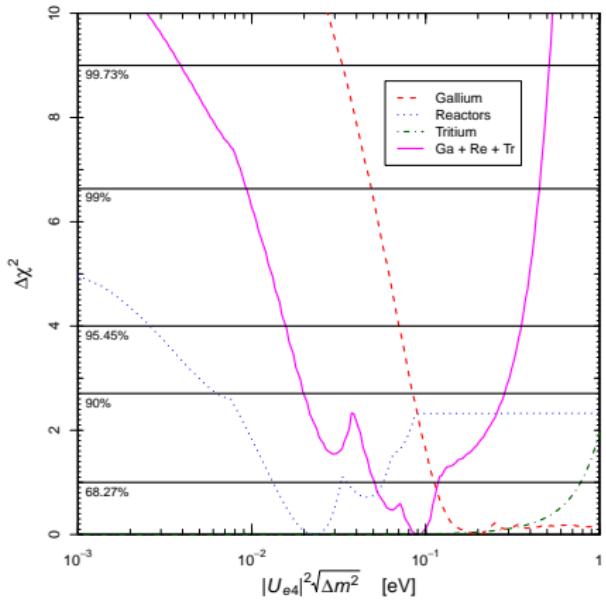
$m_{\nu_e} < 0.3 - 0.7 \text{ eV}$

# Implications of Gallium and Reactor Anomalies

$\beta$  Decay



$(\beta\beta)_{0\nu}$  Decay

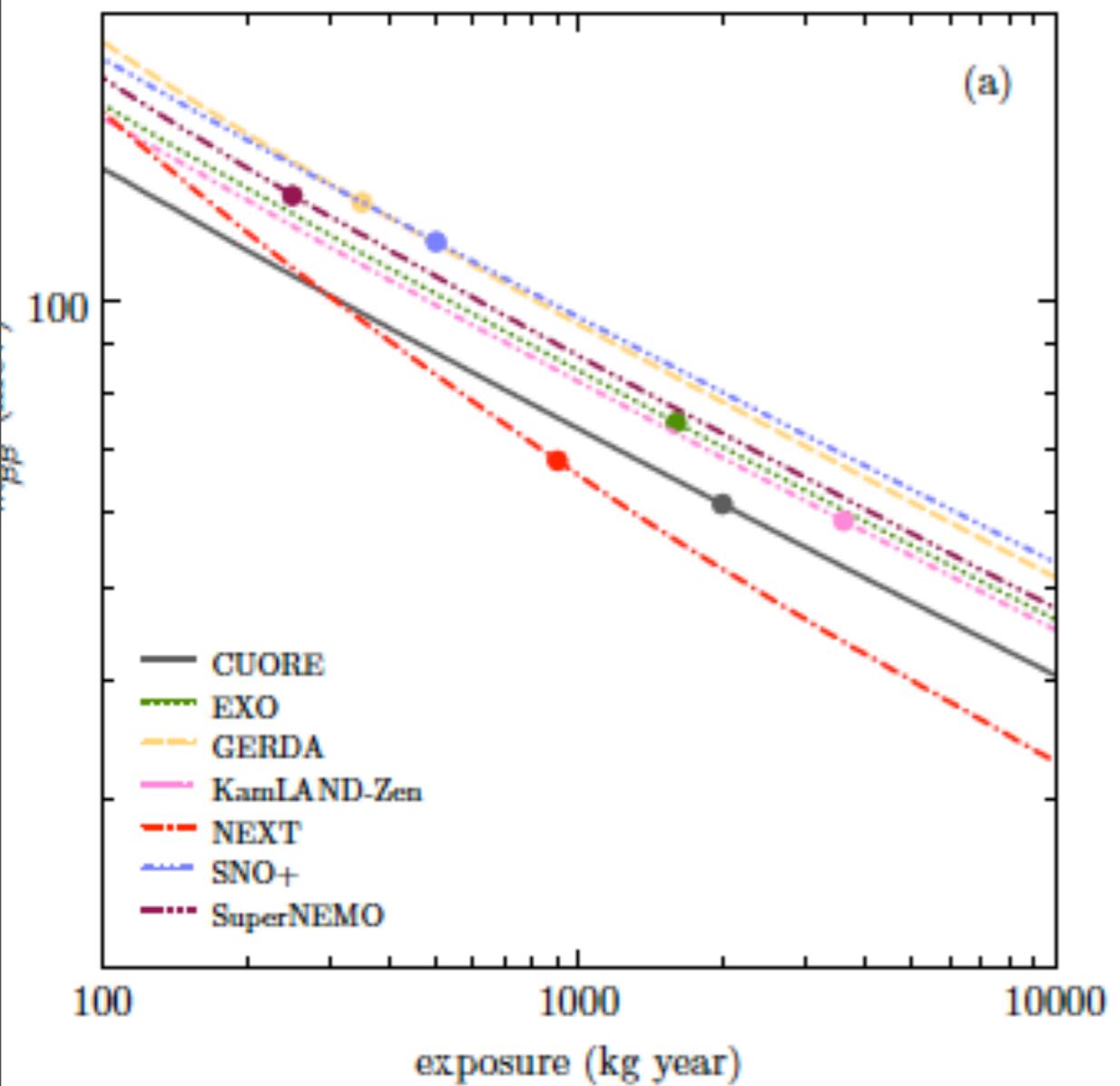


$$m_\beta^2 = \sum_k |U_{ek}|^2 m_k^2$$

$$m_{\beta\beta} = \left| \sum_k U_{ek}^2 m_k \right|$$

[Giunti, Laveder, In Preparation]

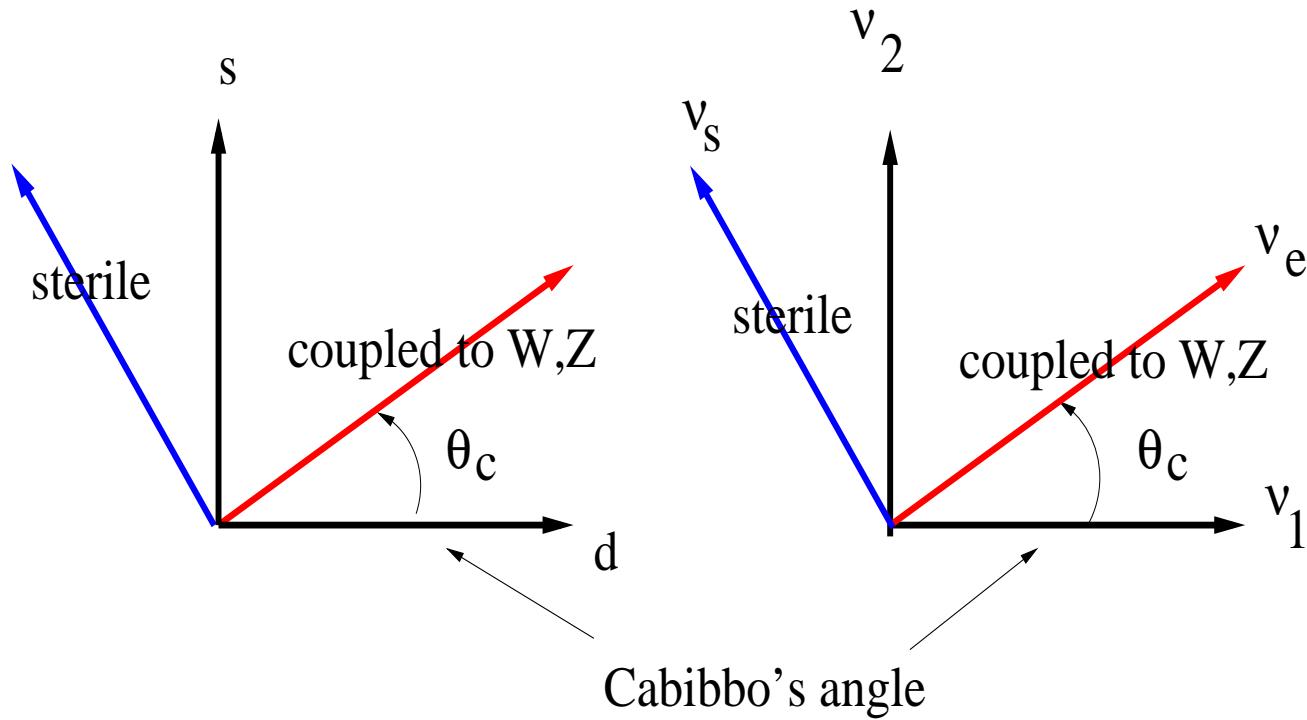
# Sensitivity of proposals



- Does not take into account “t0”
- Uses “conservative” background model (that can be more conservative for some experiments than for others)
- Uses central values for PMRs
- (should we use rather lower and upper limits of PMR?).
- Dots mean (10 years x mass)

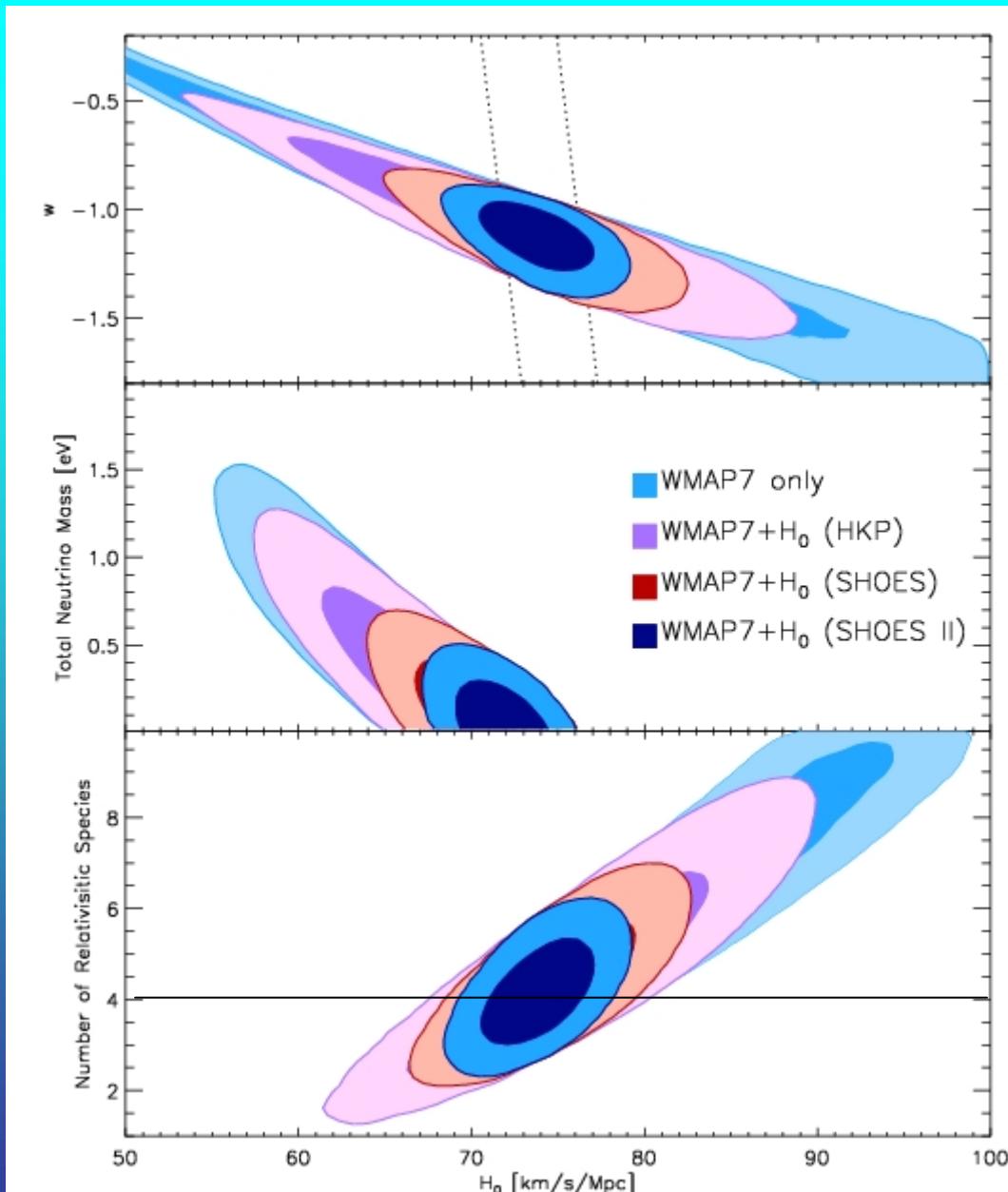
QUARKS

NEUTRINOS



COMPLEMENTARITY relation :

$$\theta_{12} \sim 32^\circ \quad \theta_{es} \sim 13^\circ \quad \theta_{12} + \theta_{es} = 45^\circ$$



... if they are roses they'll flower...



... A BRIGHT FUTURE for Majorana  $\nu$  physics !!!

*A proposed search for Sterile Neutrinos  
with the ICARUS detector at the CERN-PS*

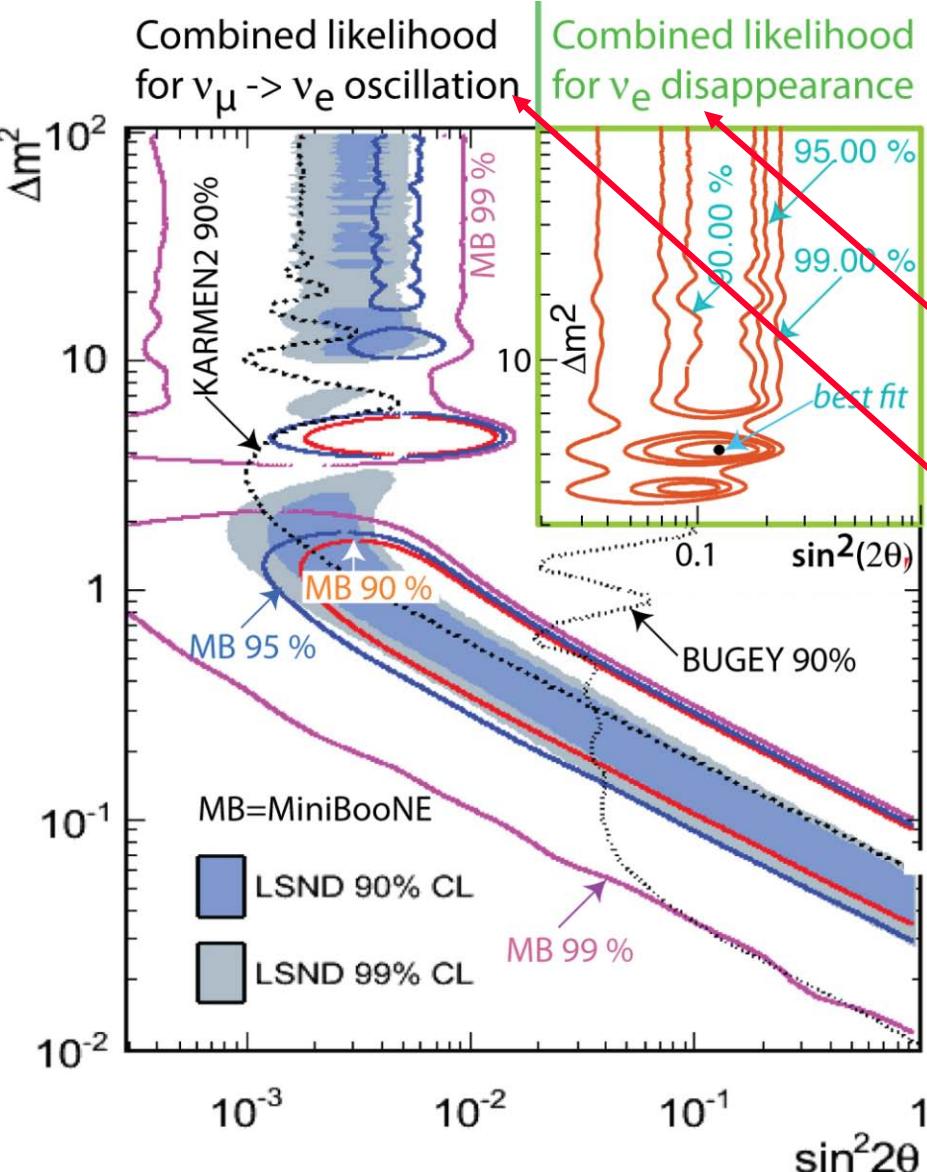
Carlo Rubbia

*INFN/LNGS, Assergi, Italy*

and

*CERN, Geneva, Switzerland*

# An unified approach ?



Allowed regions in the plane for combined results:

the  $\nu e$  disappearance rate (right)

the LSND /MiniBooNE anti- $\nu e$  anomaly (left).

While the values of  $\Delta m^2_{\text{new}}$  may indeed have a common origin, the different values of  $\sin^2(2\theta_{\text{new}})$  may reflect within the four neutrino hypothesis the structure of  $U_{(4,k)}$  mass matrix, with  $k = \mu$  and  $e$ .

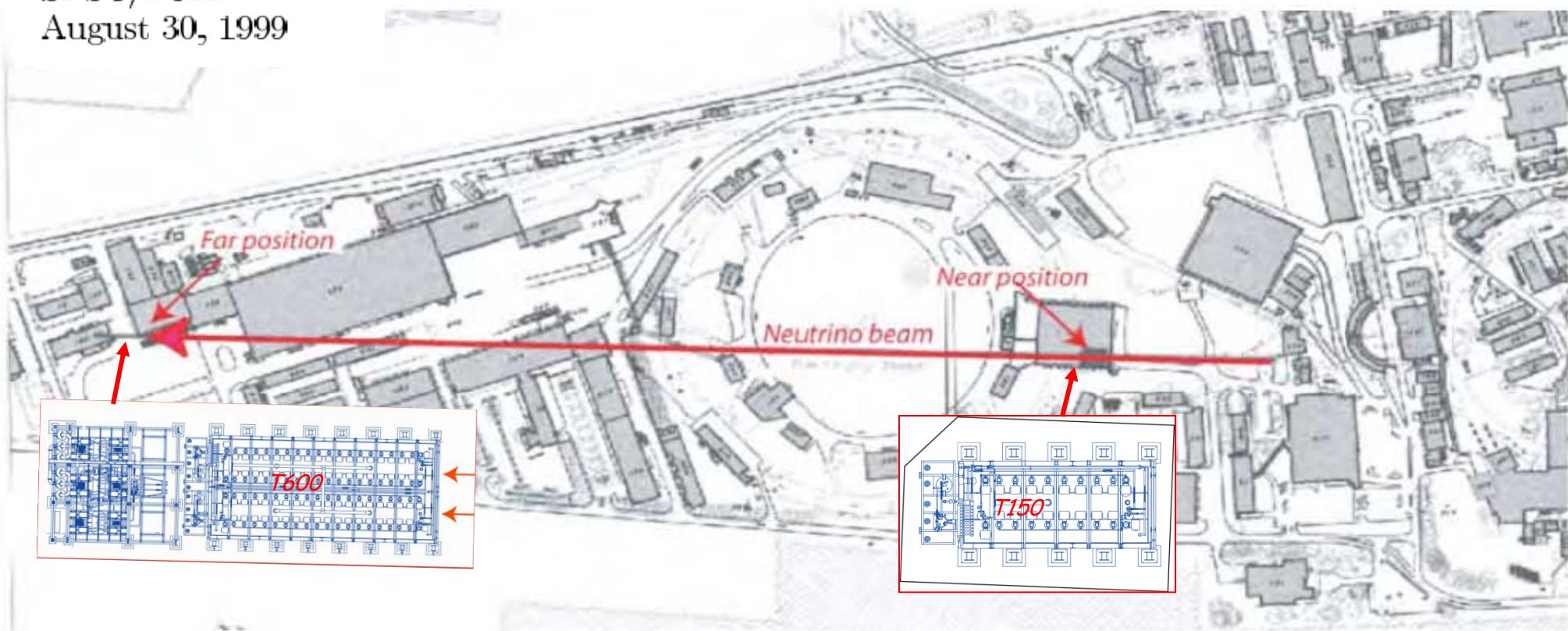
# Basic features of the proposed experiment

- Our proposed experiment, collecting a large amount of data both with neutrino and antineutrino focussing, may be able to give a likely definitive answer to the 4 following queries:
  - the LSND/+MiniBooNe both antineutrino and neutrino  $\nu\mu \rightarrow \nu e$  oscillation anomalies;
  - The Gallex + Reactor oscillatory disappearance of the initial  $\nu - e$  signal, both for neutrino and antineutrinos
  - an oscillatory disappearance maybe present in the  $\nu - \mu$  signal, so far unknown.
  - Accurate comparison between neutrino and antineutrino related oscillatory anomalies
- In absence of these “anomalies”, the signals of the detectors should be a precise copy of each other for all experimental signatures and without any need of Monte Carlo comparisons.

# Two detectors at the CERN-PS neutrino beam

CERN-SPSC/99-26  
SPSC/P311  
August 30, 1999

SEARCH FOR  $\nu_\mu \rightarrow \nu_e$  OSCILLATION  
AT THE CERN PS

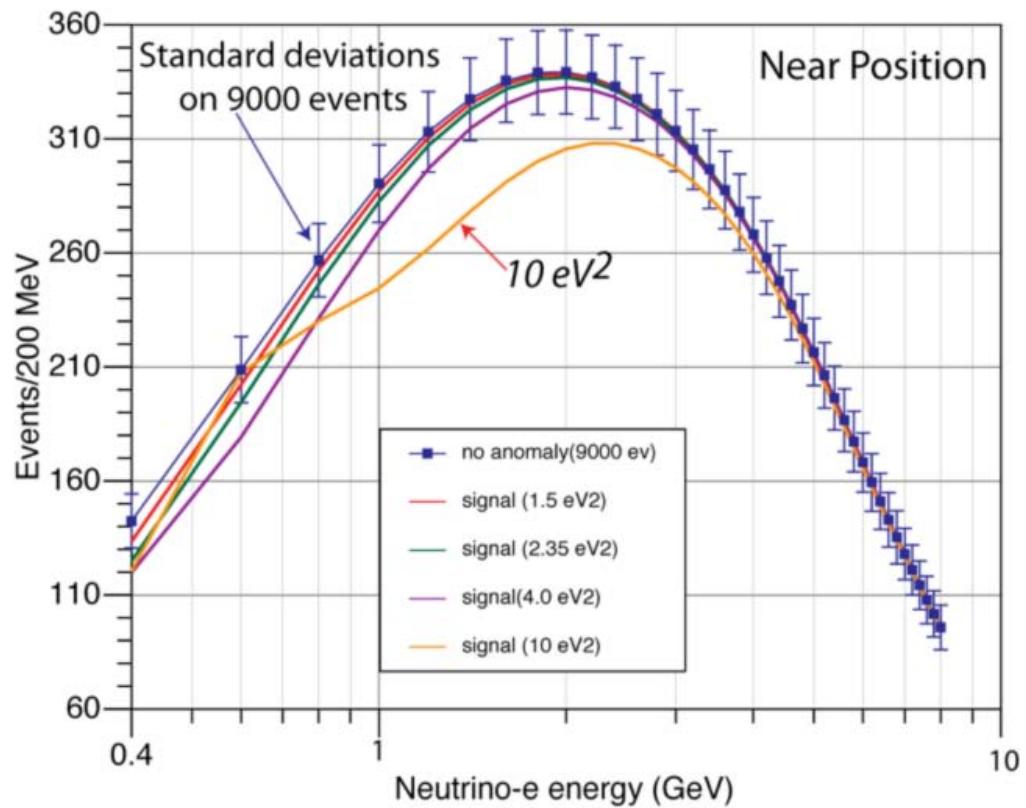
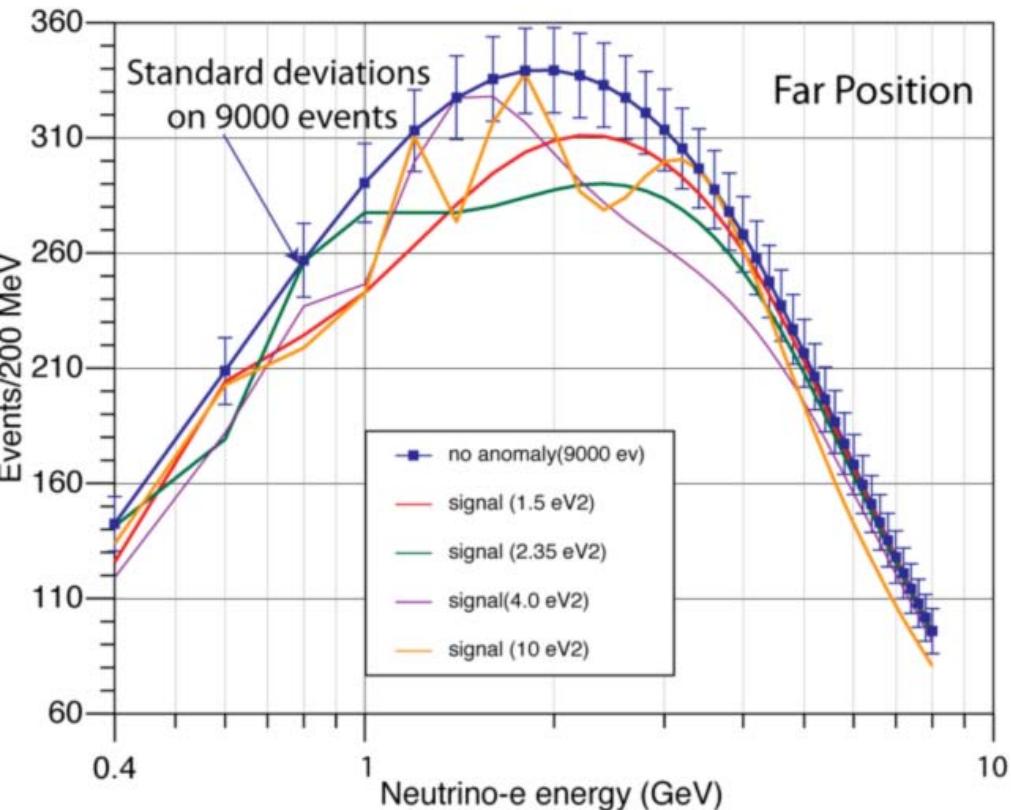


Two positions are foreseen for the detection of the neutrinos  
The far (T600) location at 850 m from the target:  $L/E \sim 1 \text{ km/GeV}$ ;  
The new location at a distance of 127 m from the target:  $L/E 0.15 \text{ km/GeV}$

# A new experiment at the CERN-PS

- The present proposal at the CERN-PS is based on the search for spectral differences of electron like specific signatures *in two identical detectors but at two different neutrino distances*, at the "Far" and the "Near" locations, respectively at 850 m and 127 m away from the source.
- The "Far" detector is the ICARUS T600, now perfectly operational in the underground Hall B of the LNGS in a neutrino beam from the CERN-SPS, collecting data as CNGS2 experiment. The T600 detector is the largest liquid Argon TPC ever built, with a size of about 600 t of imaging mass.
- The "Near" detector has to be constructed anew and it is as far as possible identical to the T600 but with a mass of 150 t, namely a clone of a single T300 half-module with the length reduced by a factor 2.

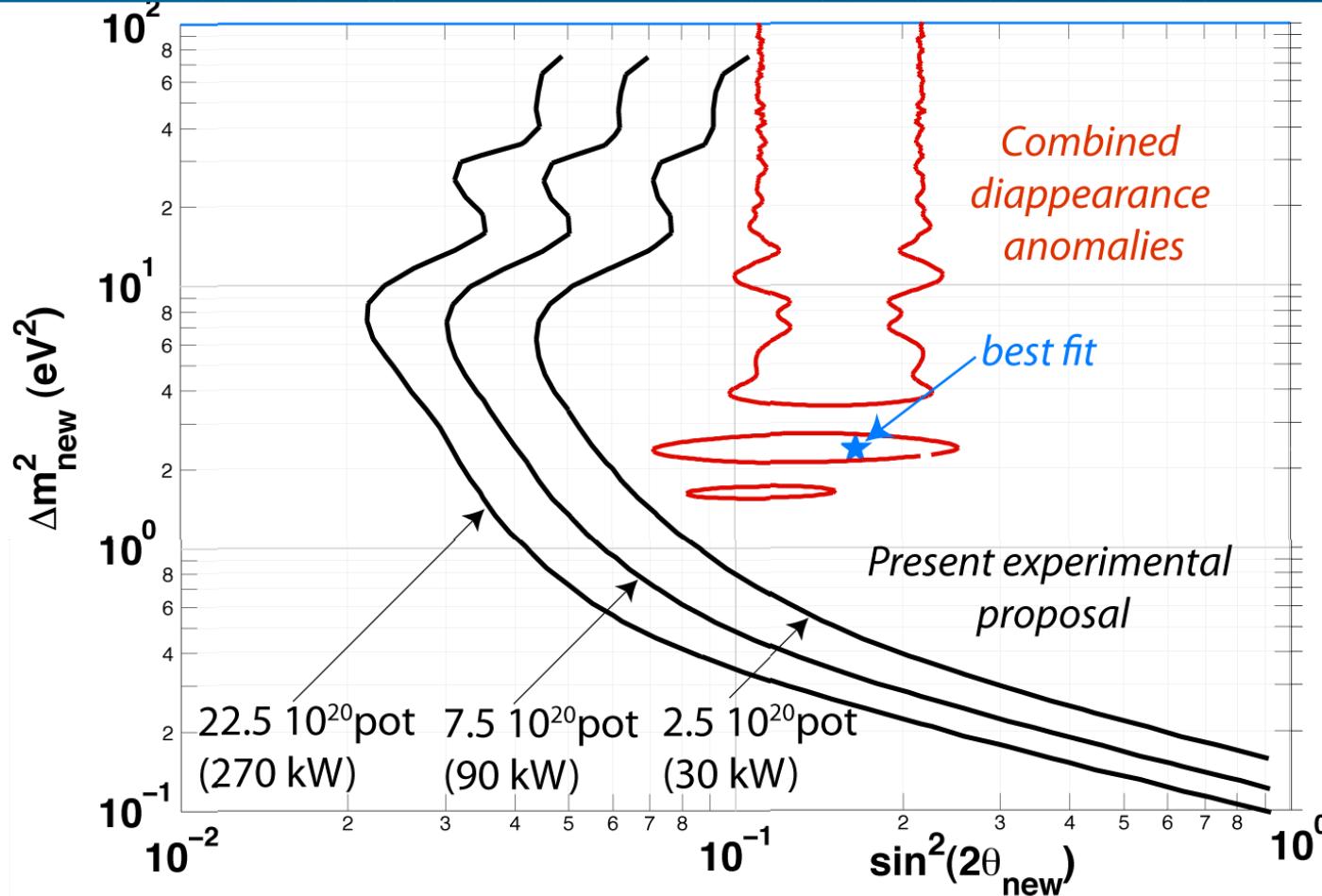
# Sensitivity to $\nu e$ (and $\nu \mu$ ) disappearance signals



The energy distributions of the electron neutrino events is shown in (a) and (b) respectively for the "Far" and "Near" and a number of possible values in the region of  $\Delta m^2 > 1\text{eV}^2$  and  $\sin^2(2\theta) \approx 0.16$ .

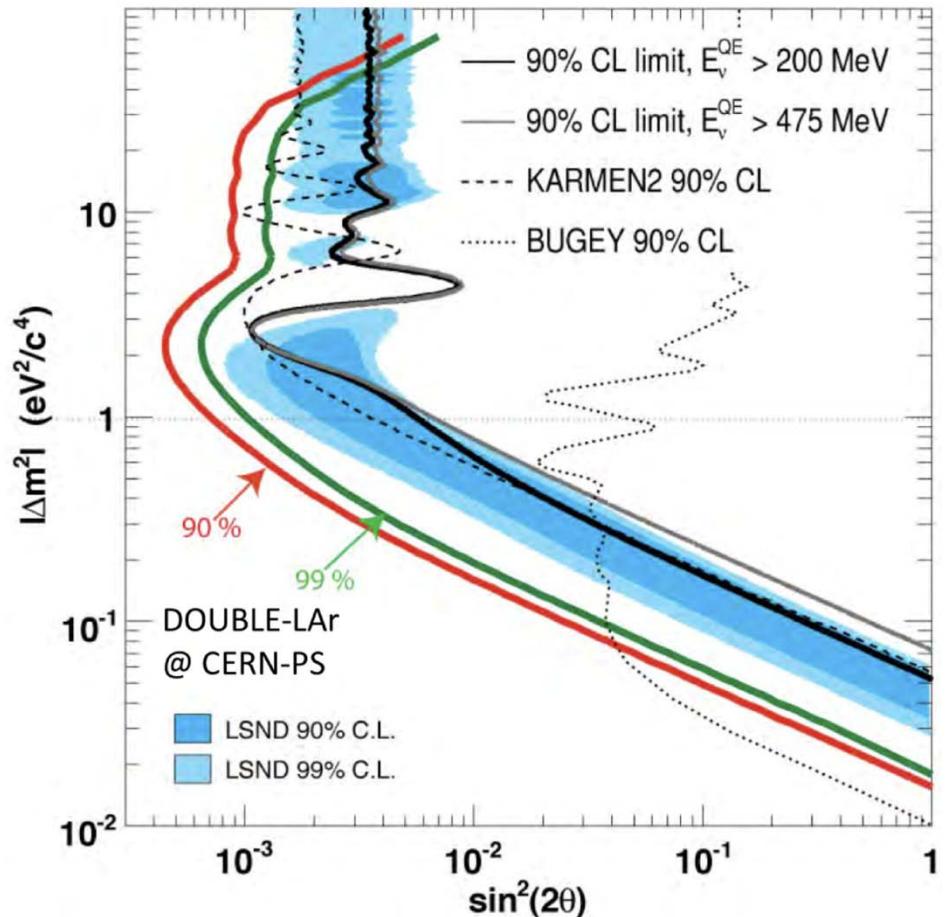
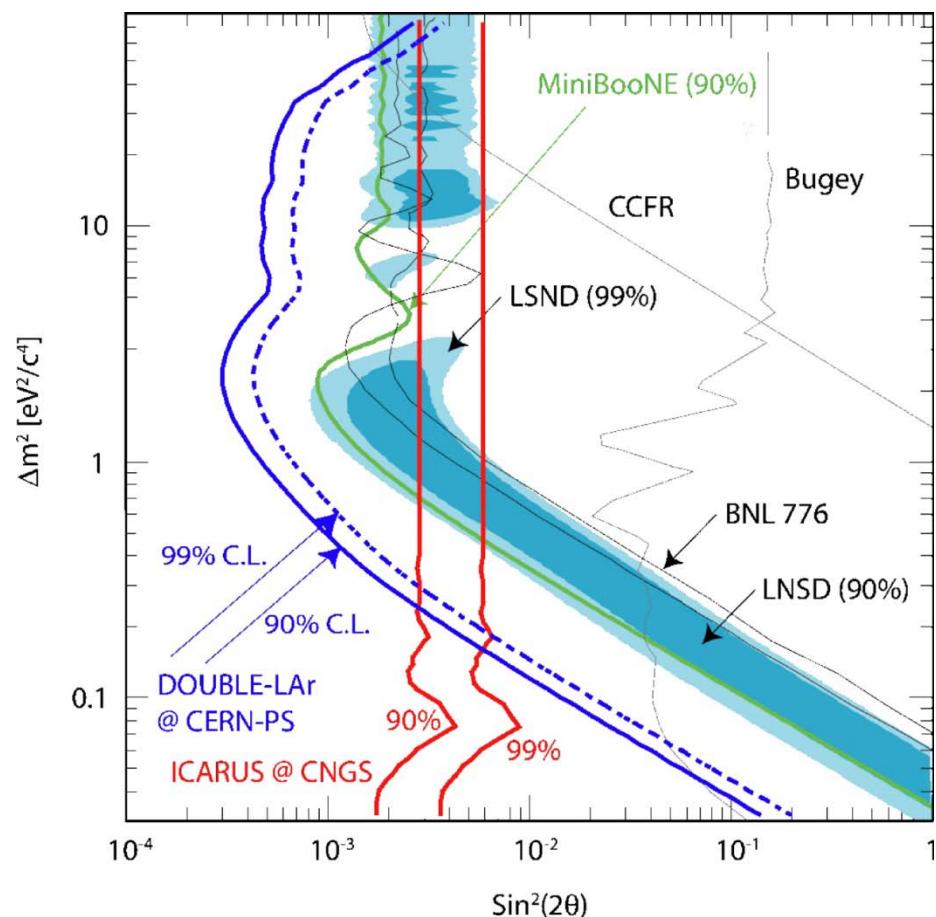
If confirmed without any doubt such a large mass difference will have an important role in the explanation of the existence of the Dark Mass in the Universe.

# Sensitivity to disappearance anomaly



- Sensitivities (90% CL) in the  $\sin^2(2\theta_{\text{new}})$  vs.  $\Delta m_{\text{new}}^2$  for an integrated intensity of (a) at the 30 kWatt beam intensity of the previous CERN/PS experiments, (b) the newly planned 90 kWatt neutrino beam and (c) a 270 kWatt curve. They are compared (in red) with the "anomalies" of the reactor + Gallex and Sage experiments. A 1% overall and 3% bin-to-bin systematic uncertainty is included (for 100 MeV bins).

# Comparing sensitivities (arXiv:0909.0355)



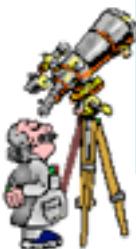
Expected sensitivity for the proposed experiment exposed at the CERN-PS neutrino beam (left) for  $2.5 \cdot 10^{20}$  pot and twice as much for anti-neutrino (right). The LSND allowed region is fully explored both for neutrinos. The expectations from one year of at LNGS are also shown.



# Possibilities for future ν beams at CERN



XIV International Workshop on “Neutrino Telescopes”  
March 15-18, 2011



“Istituto Veneto di Scienze, Lettere ed Arti”, Palazzo Franchetti - Campo Santo Stefano, Venice

Many thanks to: M. Dracos, R. Garoby, E. Gschwendtner,, A. Guglielmi, K. Long, F. Pietropaolo, A. Rubbia, R. Steerenberg, E. Wildner

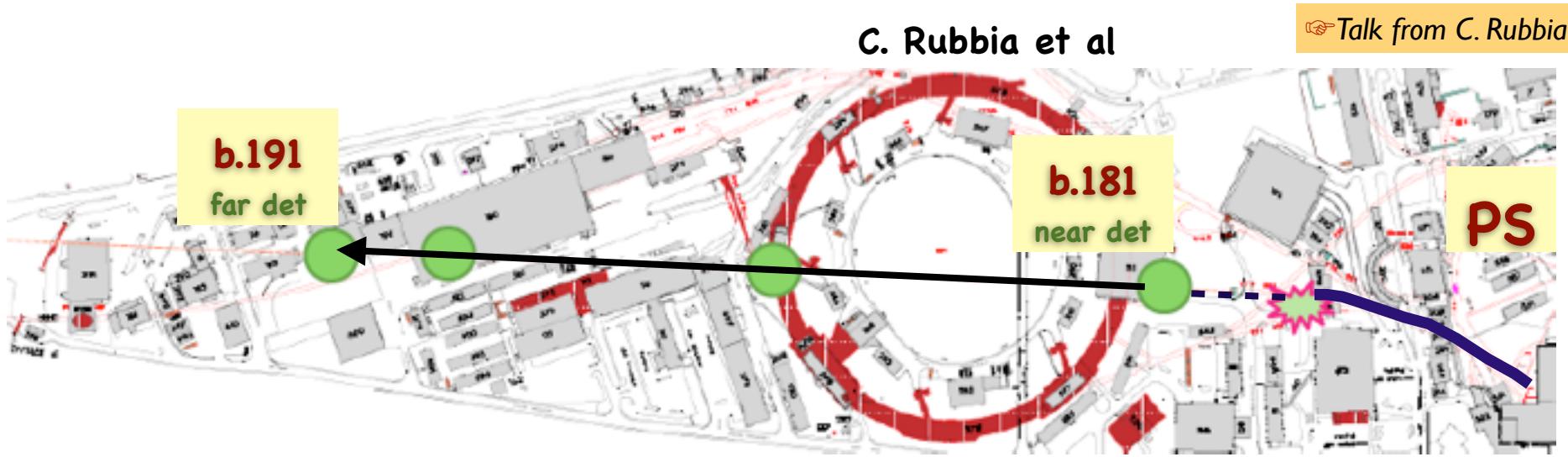


I. Efthymiopoulos – CERN

XIV International Workshop on  
“Neutrino Telescopes”  
Venice - March 17, 2011

# PS - Short Baseline ν-beam

- A search for anomalous neutrino  $\nu_\mu \rightarrow \nu_e$  oscillations at the CERN PS with LAr-TPC detectors



- Beam line originally operated in early 80's for PS169, PS181, PS180(BEBC) experiments
- **PS beam possibilities (180, 85% efficiency) :**
  - $6.13 \cdot 10^{19} \div 2.02 \cdot 10^{20}$  from zero to max impact to PS users

	Old neutrino facility		New neutrino facility		
	PS dedicated Feb-Mar 1983	PS parallel 1983 - 1984	PS dedicated	PS parasitic	PS ultimate <sup>2</sup>
Proton Momentum	19.2 GeV/c	19.2 GeV/c	20 GeV/c	20 GeV/c	26 GeV/c
Protons/pulse	$1.25 \times 10^{13}$	$1.2 \times 10^{13}$	$3 \times 10^{13}$	$2.6 \times 10^{13}$	$4 \times 10^{13}$
Max. rep. rate	1.2 s	14.4 s	1.2 s	1.2 s	1.2
Beam energy	38 kJ	38 kJ	96 kJ	84 kJ	166 kJ
Average beam power	32 kW	2.5 kW	80 kW	70 W	140 kW

Courtesy: R. Steerenberg – CERN

# First Neutrino Oscillation Results from the T2K Experiment

André Rubbia (ETH Zurich)  
for the T2K Collaboration



XIV International Workshop on Neutrino Telescopes  
Venice, March 15-18<sup>th</sup> 2011



# Message from KEK



Japan experienced very severe earthquake on March 11th 2011 at 14:46 JST. J-PARC facility suffered damages for some extent. There are no reports of casualties and all staff, graduate students, and foreign visitors have been located and as of evening Sunday March 13th all T2K members have been evacuated from Tokai area.

Fortunately enough, the Tsunami tidal wave did not hit J-PARC. We will start the investigation of the facilities. We will update the announcement as we learn the detail of the entire damage.

Our present priority is to restore life-supporting infrastructure such as electricity, water supply and gas at J-PARC. It may take some time, but we promise the full recovery of the J-PARC accelerator and T2K experiment in the near future.

I thank you for the messages of solidarity and sympathy.

Director of the Institute of Particle and Nuclear Studies, KEK

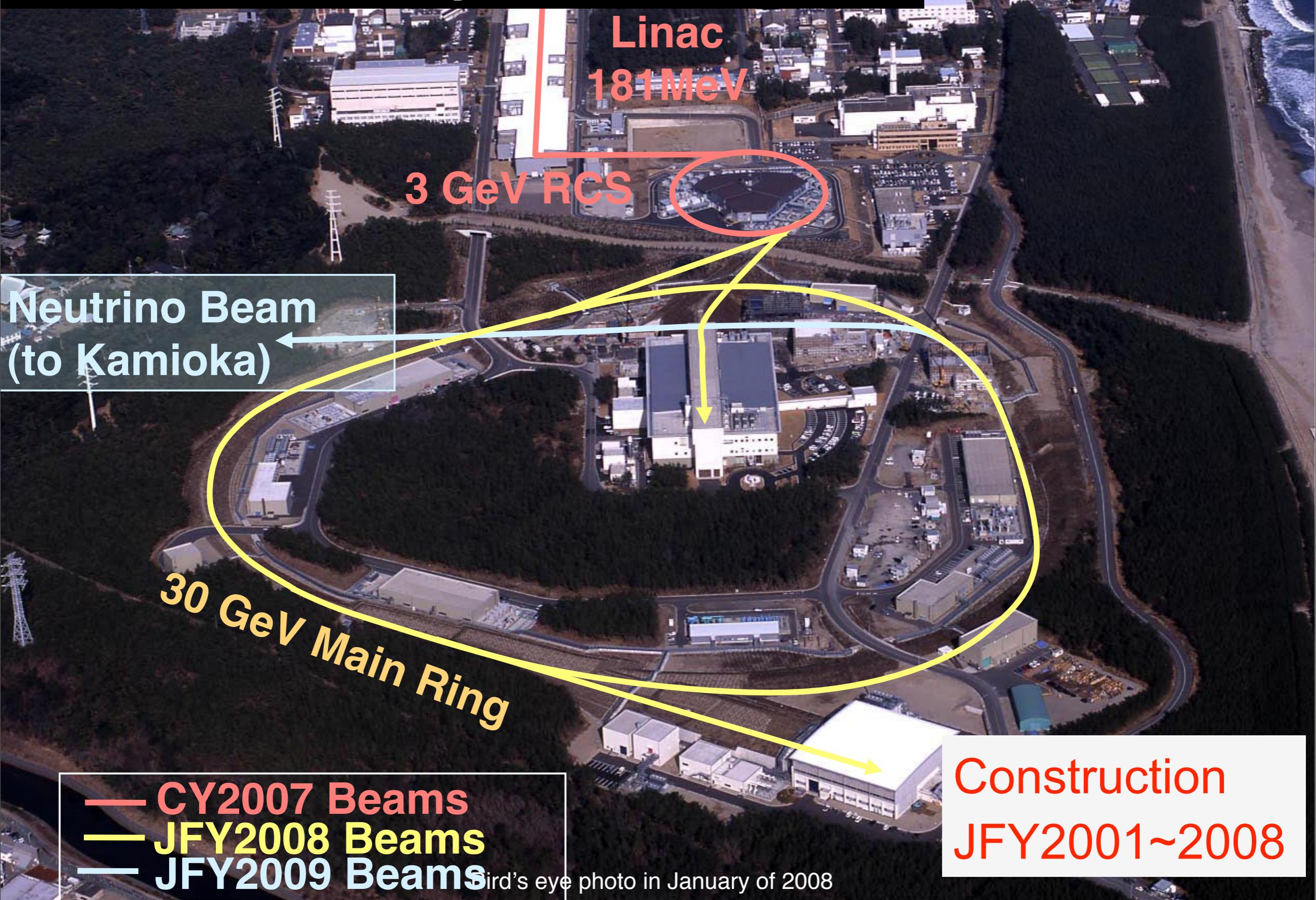
Koichiro Nishikawa

Spokesperson of the T2K experiment

Takashi Kobayashi

# J-PARC Facility (KEK/JAEA)

North ↑



# ND280 off-axis detector overview



**Two main target regions:**

- *Pi-0 Detector (P0D)*: optimised for (NC)  $\pi^0$  events
- *Tracker*: optimised for charged particle final states

**Both regions have passive water planes**

*P0D, Barrel and DownStream ECAL*

**Scintillator planes with radiator**

*Measure EM showers from inner detector  
( $\gamma$  for NC  $\pi^0$ , bremstrahlung in  $\nu_e$  measurement)  
Sand muon rejection*

**UA1 magnet (0.2T) Inner volume 3.5x3.6x7m<sup>3</sup>**

*Yoke Fe mass ~ 900 tons*

**SMRD (Side Muon Range Detector)**

**Scintillator planes in magnet yoke.**  
*Detect muons from inner detector  
(neutrino rate, side muon veto, cosmic trigger)  
Momentum measurement*

**P0D ( $\pi^0$  Detector)**

**Scintillators planes interleaved with water and lead/brass layers**  
*Optimised for  $\gamma$  detection*

**2 FGDs (Fine Grained Detectors) 3 TPCs (Time Projection Chambers):**

**Thin, wide scintillator planes**

*Provides active target mass  
Optimised for p recoil detection*

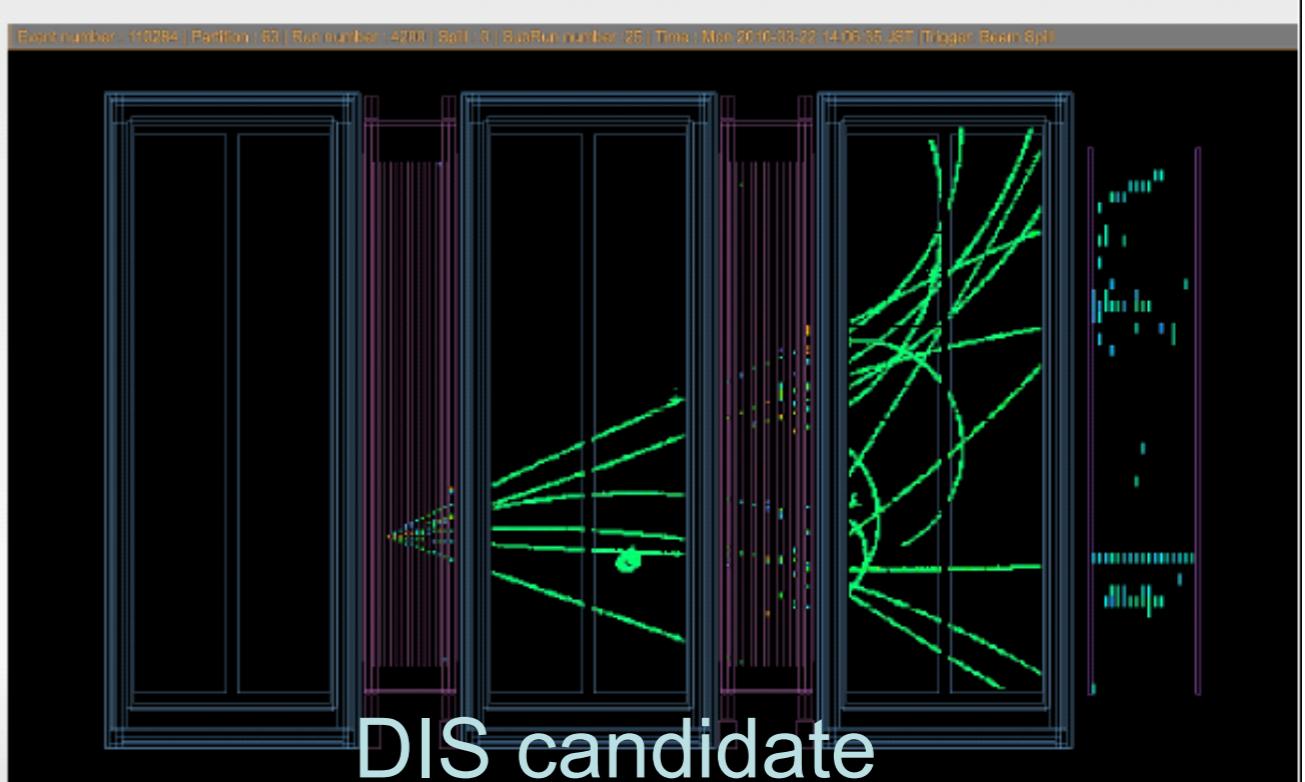
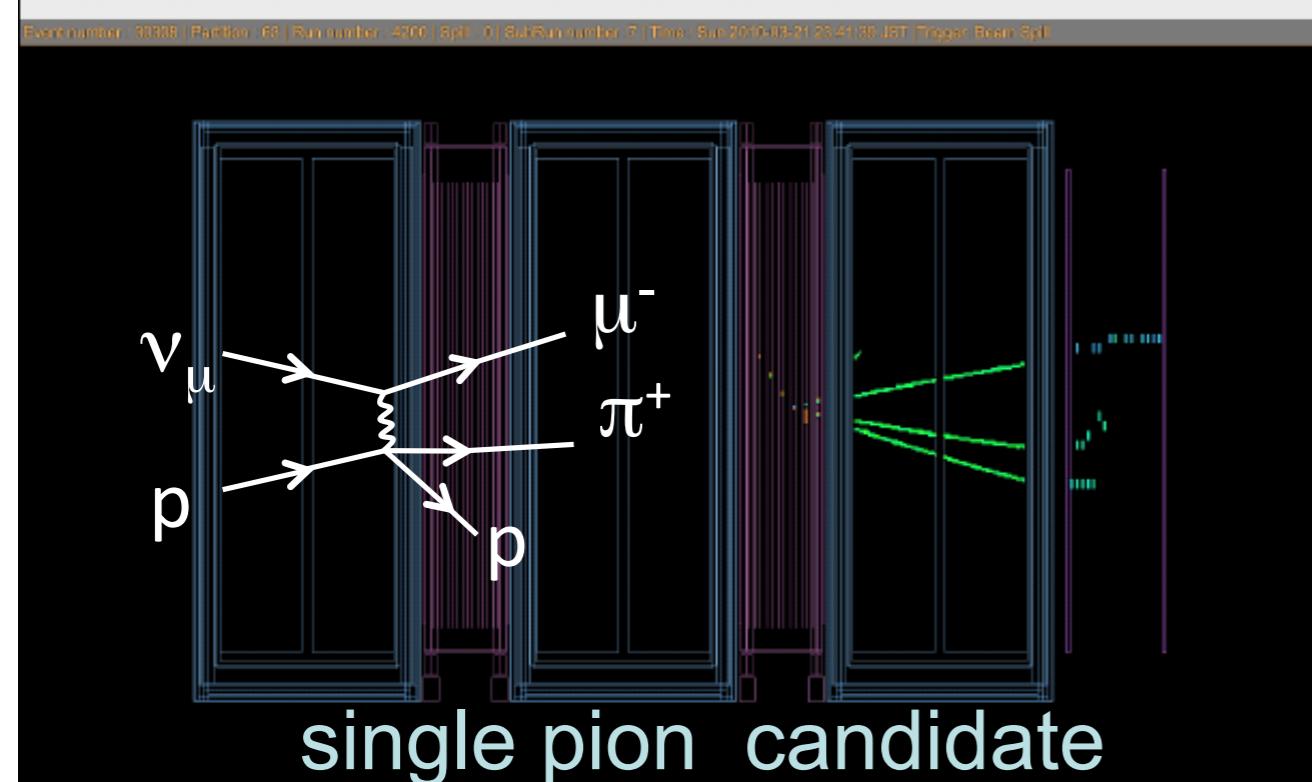
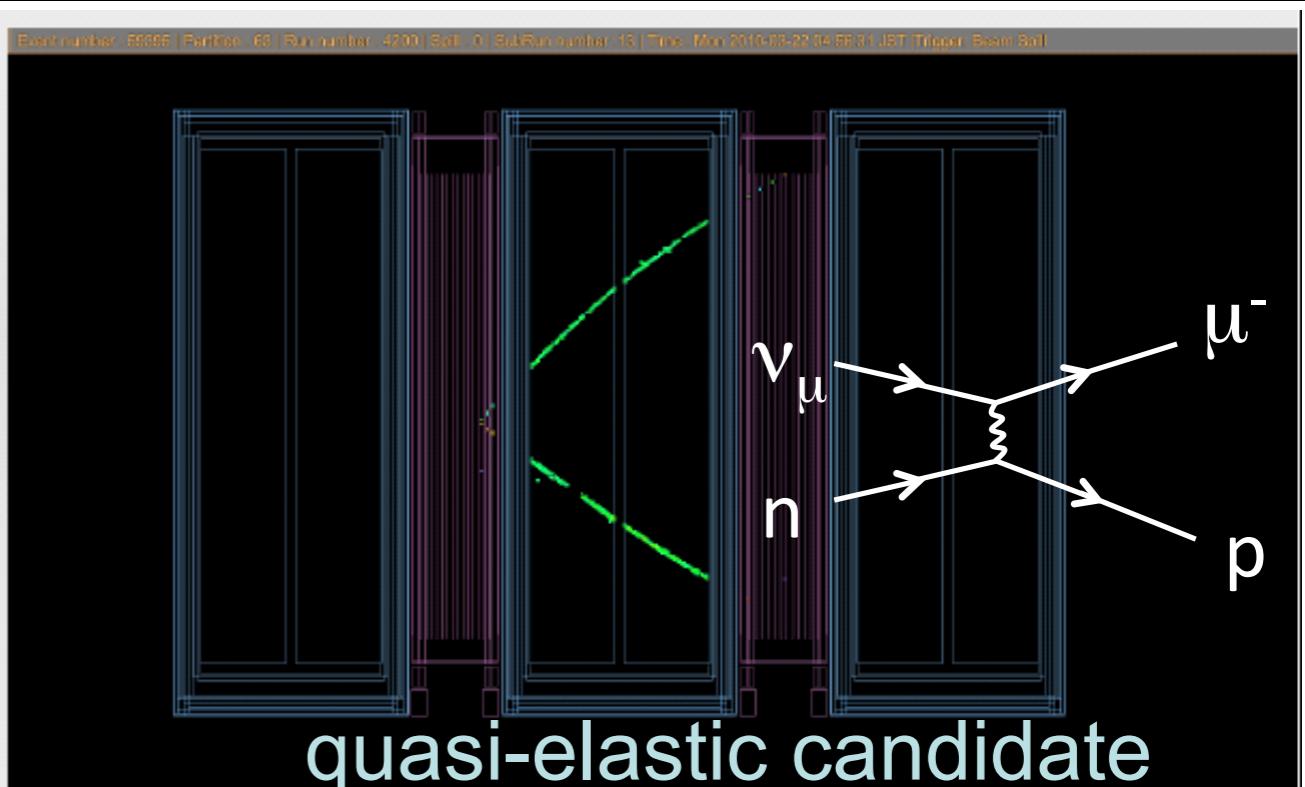
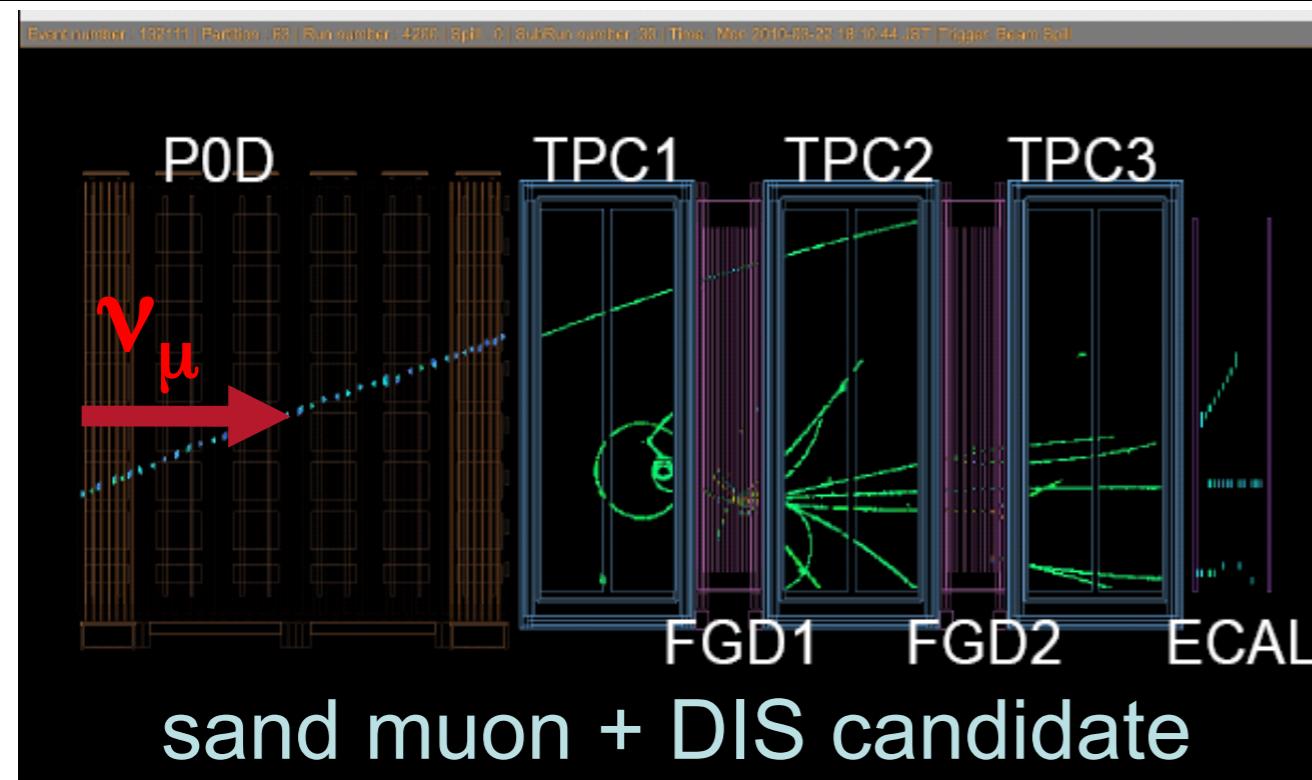
*Momentum measurement of charged particles from FGD and P0D  
PID via dE/dx measurement*

FGD1: Scintillator planes ~ 1 ton,

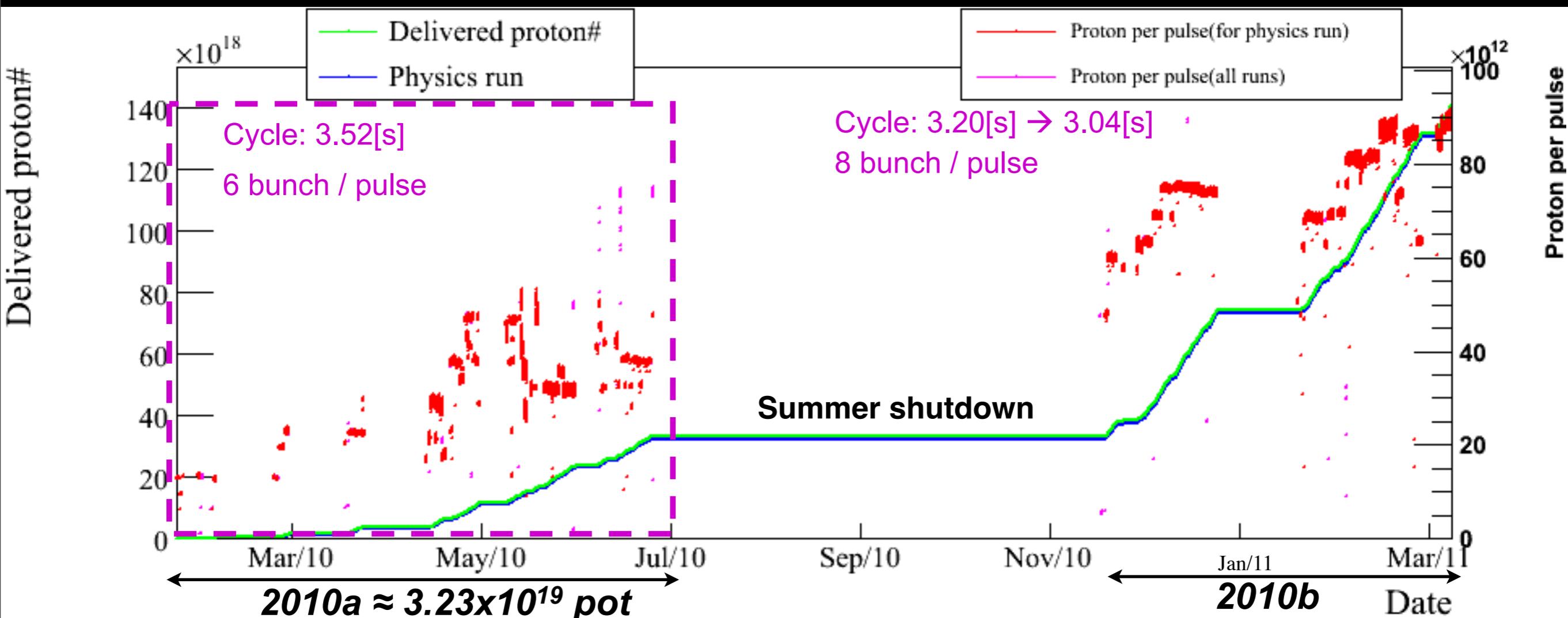
FGD2: Scinti. & H<sub>2</sub>O planes ~ 0.5 & 0.5 ton

**P0D mass:**  
16.1 tons w/ water  
13.3 tons w/o water

# ND280 off-axis event gallery



# MR protons# delivered



## Run 2010a (Jan-Jun 2010)

- 6 bunches/spill, cycle: 3.52 sec
- $3.23 \times 10^{19}$  p.o.t for T2K analysis
- 50kW stable beam operation (trials at 100 kW)
- Super-K live time >99%

## Run 2010b (Nov 2010-??? 2011)

- 8 bunches/spill,  $9 \times 10^{13}$  ppp
- cycle: 3.52 sec → 3.04 sec
- 135kW → **145 kW beam power**
- $1.45 \times 10^{20}$  p.o.t accumulated so far
- MR intensity limited by losses

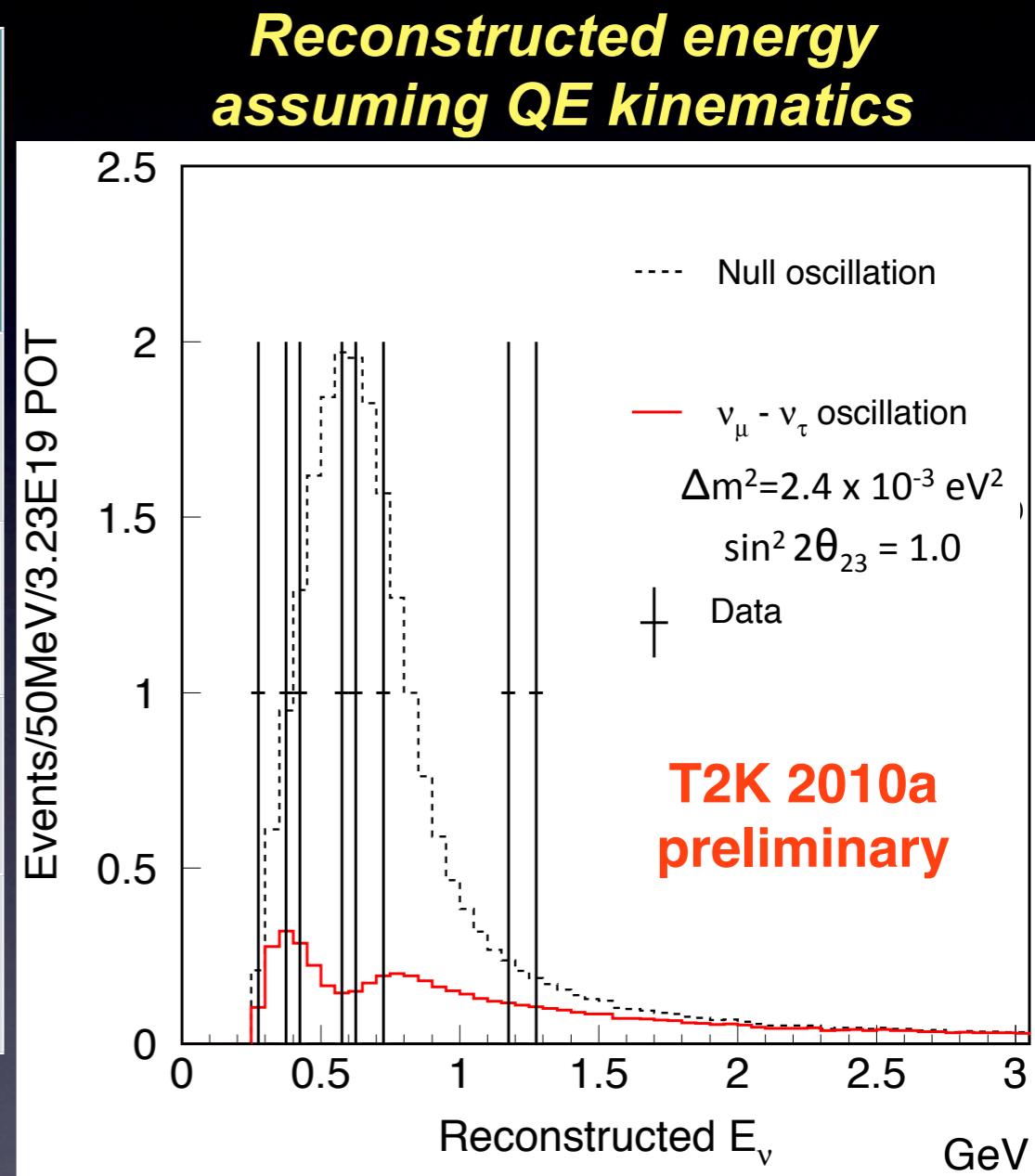
# $\nu_\mu$ disappearance analysis



## Event selection for muon disappearance measurement

T2K-SK events	Data	MC		Acc.BG (12μs window)
		No oscillation	W/ oscillation	
Fully-Contained	33	54.5	24.6	0.0094
Fiducial Volume, $E_{\text{vis}} > 30\text{MeV}$	23	36.8	16.7	0.0011
Single-ring $\mu$ -like $P_\mu > 200\text{MeV}/c$	8	$24.5 \pm 3.9$	$7.1 \pm 1.3$	-
+ number decay-e $<=1$ & $E_{\text{rec}} < 10 \text{ GeV}$	8	$22.8 \pm 3.2$	$6.3 \pm 1.0$	-

$\Delta m^2 = 2.4 \times 10^{-3} \text{ eV}^2$   
and  $\sin^2 2\theta_{23} = 1.0$



- Consistent with oscillation parameters measured by MINOS / SK / K2K
- Parameter fitting underway – T2K plans to release result in the near future

# Expected #SK events



Source	Estimated number
Beam $\nu_\mu$ (CC+NC)	0.13
Beam $\bar{\nu}_\mu$ (CC+NC)	0.01
Beam $\nu_e$ (CC)	0.16
<b>Total background</b>	<b><math>0.30 \pm 0.07 \text{ (syst.)}</math></b>
<b>Total sig.+background</b>	<b><math>1.20 \pm 0.23 \text{ (syst.)}</math></b>

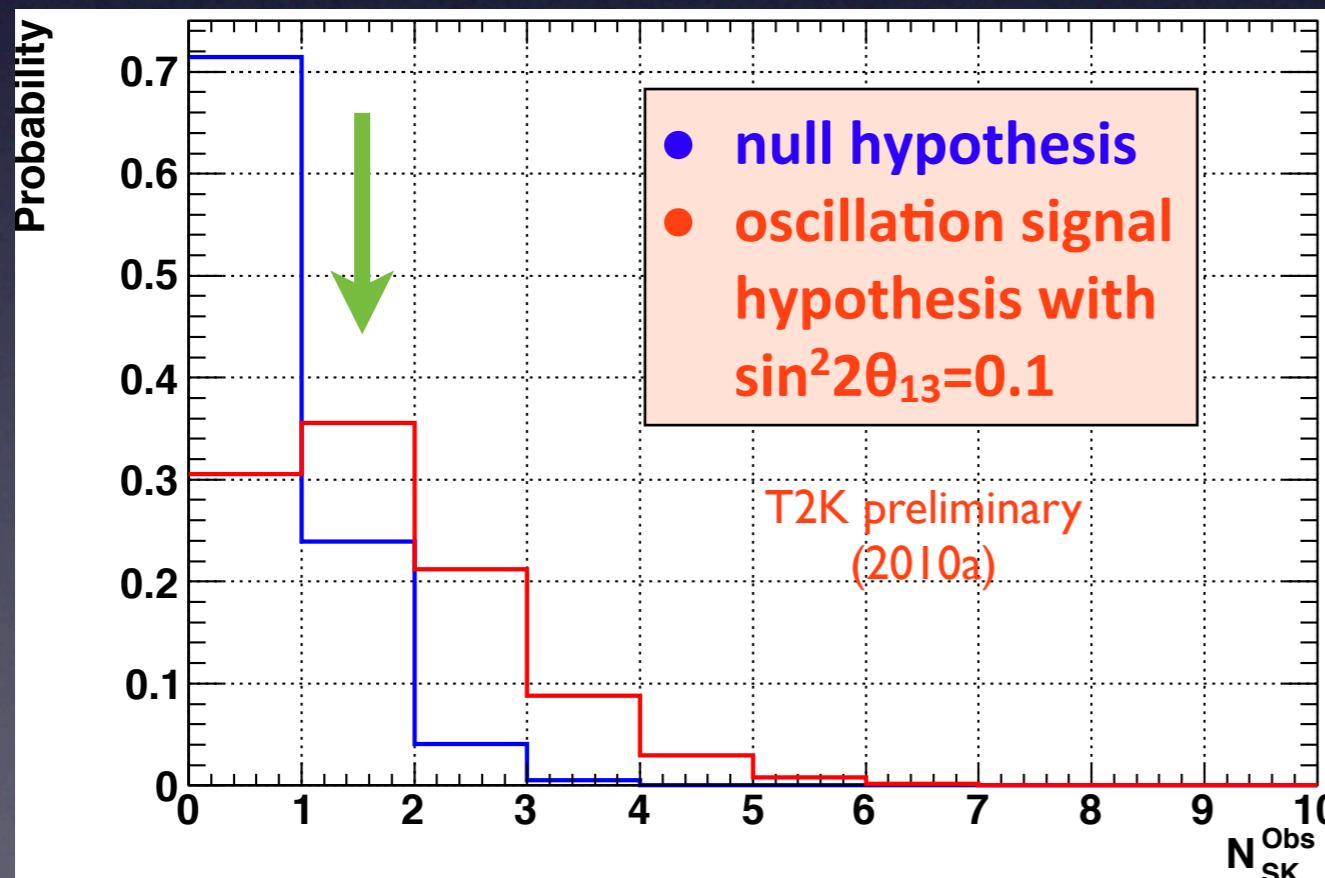
- #events normalized to p.o.t. and corrected for ND280  $\nu_\mu$  CC measured normalization
- Assumed oscillation parameters for signal:

$$\Delta m^2_{23} = 2.4 \cdot 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta_{23} = 1.0$$

$$\sin^2 2\theta_{13} = 0.1$$

$$\delta_{CP} = 0$$



T2K preliminary  
(2010a)

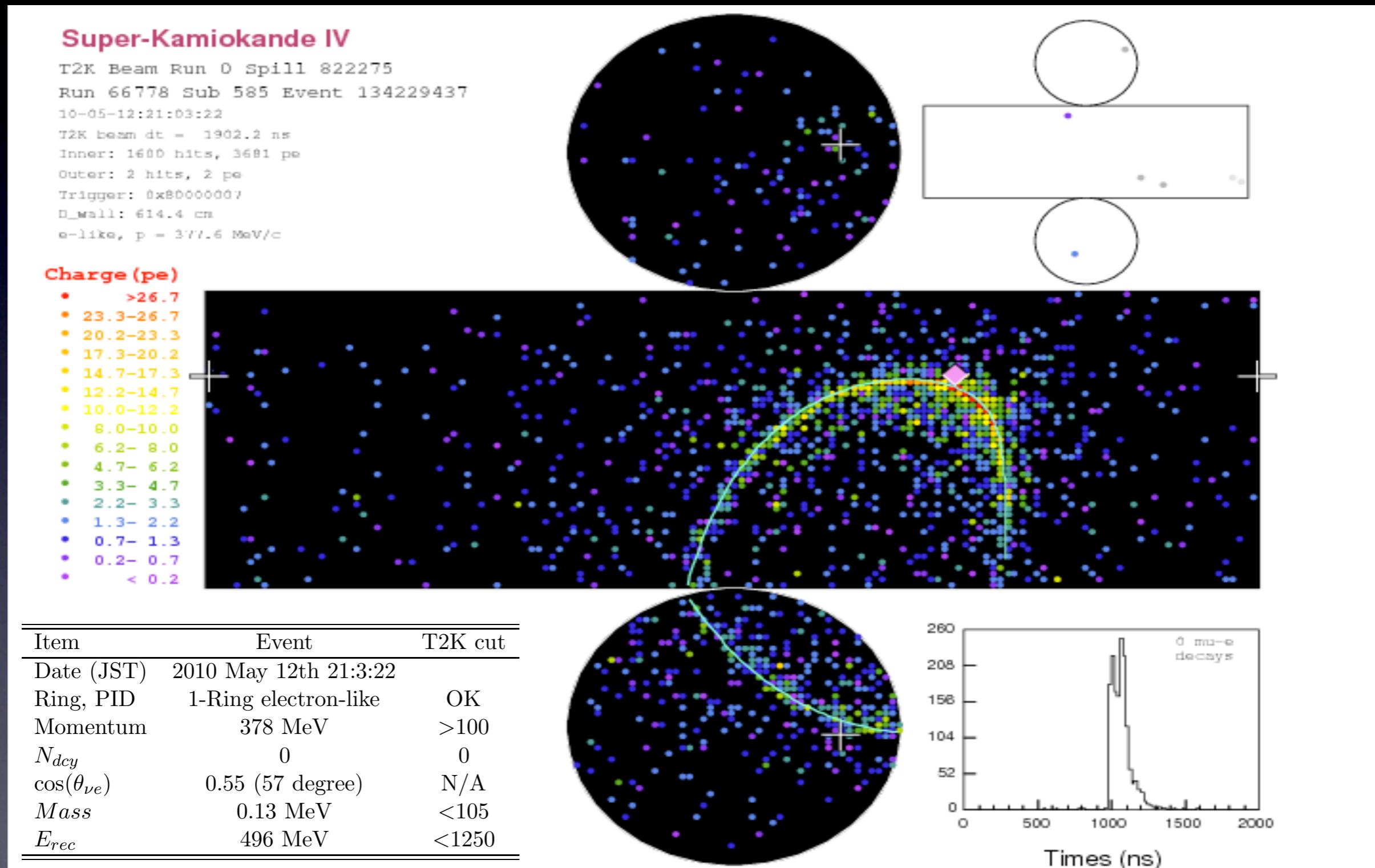
~29% probability to observe  
>=1 event when expected  
average = 0.3 event

1 data candidate!  
 $N_{\text{SK}}^{\text{obs}} = 1$

# T2K $\nu_e$ CC signal candidate (2010a)



*Signal candidate event passing all cuts*



# T2K appearance upper limit results



## *Two independent statistical procedures*

T2K preliminary  
(2010a)

Assuming  $\Delta m^2_{23} = 2.4 \cdot 10^{-3} \text{ eV}^2$  and  $\sin^2 2\theta_{23} = 1.0$ ,  $\delta_{CP} = 0$ :

### (A) Feldman-Cousins

Normal Hierarchy :  $\sin^2(2\theta_{13}) < 0.50$  (90% C.L.)  
Inverted Hierarchy :  $\sin^2(2\theta_{13}) < 0.59$  (90% C.L.)

### (B) Classical one-sided

Normal Hierarchy :  $\sin^2(2\theta_{13}) < 0.44$  (90% C.L.)  
Inverted Hierarchy :  $\sin^2(2\theta_{13}) < 0.53$  (90% C.L.)

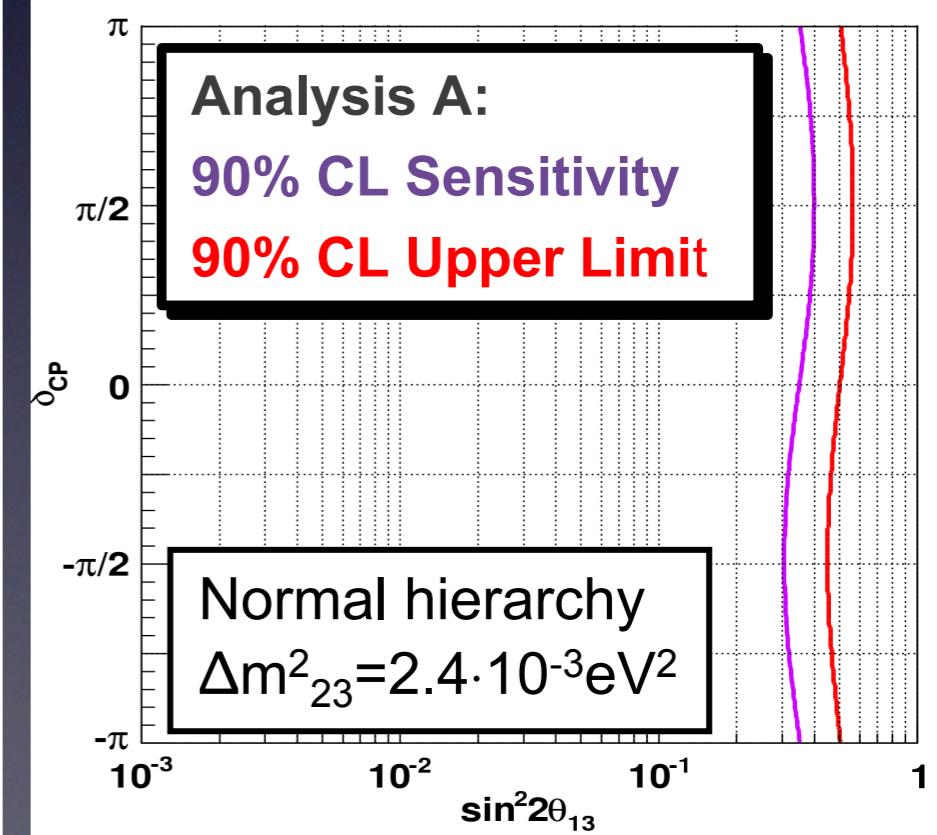
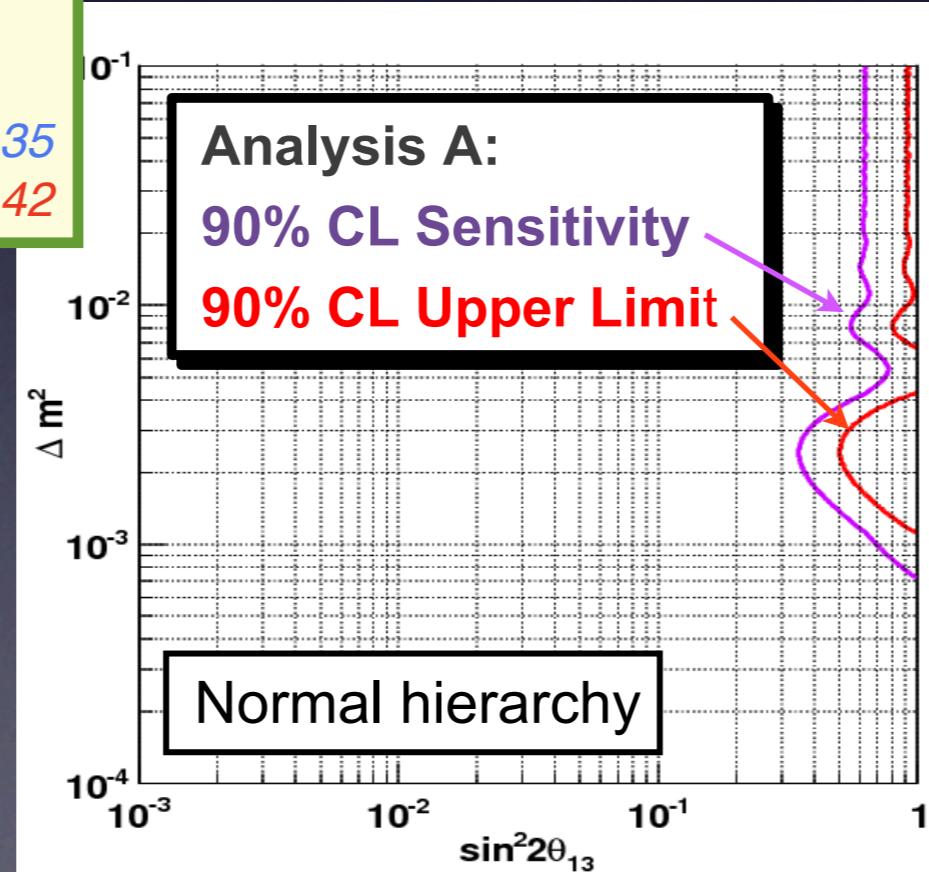
#### T2K 2010a 90% C.L. sensitivity

##### (Analysis A):

Normal Hierarchy :  $\sin^2(2\theta_{13}) < 0.35$

Inverted Hierarchy :  $\sin^2(2\theta_{13}) < 0.42$

- More collected data on tape
- Analyses underway



As a last speaker, on behalf of all participants I would like to most warmly thank the Organisers of this Conference

I am sorry that this is my first NeuTel without the charming presence of Milla Baldo Ceolin who has always been the Queen of the place

So I extend our best greetings to Milla and our warmest wishes of a prompt recovery



# EPILOGUE

Three ( $\nu$ ) gondolas are safe in the harbor...  
...but that's not what they are made for.  
New gondolas might join, and all lead us  
towards new (physics) horizons



Thank you for your attention