# Neutrinos: active, sterile or tachionic ?



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Turin – 6/12/2011

## Il cannocchiale di Galileo sembra infrangere le incorruttibili sfere e rimuovere l'orizzonte





## **Neutrinos:** active



## La soluzione Tolemaica





#### Matrice di mixing dei neutrini



La matrice di mixing dei neutrini  $U_{\rm MNSP}$  è molto diversa dalla matrice di mixing dei quark  $U_{\rm CKM}$ 

$$U_{MNSP} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix},$$

 $\theta_{13} \rightarrow 0 \implies$  La matrice 3x3 di mixing diventa il prodotto di due matrici 2x2. La conoscenza di  $\theta_{13}$  prelude ad ogni futura ricerca di violazione di CP leptonica.

Neutrini Atmosferici

Unknown

Neutrini Solari

## J-PARC Facility (KEK/JAEA)

Gev

Linac

FFF

## Neutrino Beam (to Kamioka)



30 Gev Main Ring

## Construction JFY2001~2008

North 1

Wednesday, March 16, 2011

## ND280 off-axis detector overview

#### Two main target regions:

- Pi-0 Detector (P0D): optimised for (NC)  $\pi^0$  events
- *Tracker: optimised for charged particle final states* **Both regions have passive water planes**



A. Rubbia

XIV International Workshop on Neutrino Telescopes (2011)

ETH

# ND280 off-axis event gallery



A. Rubbia

Wednesday, March 16, 2011



#### Reconstructed v energy cut ( $E_{rec} < 1250 \text{ MeV}$ ) : Final cut



#### June 2011 breakthrough: new <u>T2K</u> and MINOS appearance results!

→ "Evidence of θ<sub>13</sub>>0 from global neutrino data analysis" [Fogli, EL, Marrone, Palazzo, Rotunno, arXiv:1106.6028]



# **0**13 Preliminary Results

First neutrino oscillation data release of DC at LowNull @ Seoul (Korea)

Far detector data only

No-Oscillation: reactor flux prediction

Rate + Shape Analysis:  $sin^2(2\theta_{13}) = 0.085 \pm 0.029(stat) \pm 0.042(syst)$ 

Rate Only:  $sin^{2}(2\theta_{13}) = 0.093 \pm 0.029(stat) \pm 0.073(syst)$ 



# Global $\theta_{13}$ Hunt

DC results consistent with T2K results

Major impact in  $\theta$ I3 knowledge today by DC

Combined effect:  $\sin^2(2\theta_{13})$ >0 @ 3 $\sigma$ 



#### Words of caution, from someone we all miss:



#### "Half of all 3 $\sigma$ results are wrong."

J.N. Bahcall

No doubt: we need further experimental tests of  $\theta_{13}$ >0 and, in particular, at least a single result at >3 $\sigma$  level !

But, it would be very surprising if the current evidence were just due to a conspiracy of fluctuations in rather different categories of neutrino oscillation experiments.

In planning future v beams/detectors, useful to consider the likely possibility that  $sin^2\theta_{13}\sim0.02$ 

#### Vertex distribution of $v_e$ candidate events



 $\rightarrow$  Perform several checks. for example

- \* Check distribution of events outside FV  $\rightarrow$  no indication of BG contamination
- \* Check distribution of OD events  $\rightarrow$  no indication of BG contamination
- \* K.S. test on the R<sup>2</sup> distribution yields a p-value of 0.03

#### Status of 3v oscillation parameters, circa 2011



G.L.Fogli, E.L., A.Marrone, A.Rotunno (Univ. & INFN, Bari), A. Palazzo (TUM, Munich) "Evidence of  $\theta_{13}$ >0 from global neutrino data analysis", arXiv:1106.6028v2 Vacuum frequency  $\delta m^2/2E$  (unknown sign) interferes with  $v_e$  interaction energy  $\sqrt{2G_FN_e}$  (known sign) in solar matter: MSW

→ Get  $(v_1, v_2)$  hierarchy: state with largest  $v_e$  component is the lightest of the two - conventionally,  $v_1$ 

#### Main features of matter effects: established | Spectral details: need higher stat, lower E





Here at TAUP 2011: Final SNO data analysis and first Borexino pep flux data!

# Development of the SSM after 2004: impact on neutrino fluxes

Model	pp	pep	<sup>7</sup> Be	<sup>8</sup> B	<sup>13</sup> N	<sup>15</sup> O	<sup>17</sup> F	Y <sub>surf</sub>	R <sub>cz</sub> /
	×10 <sup>10</sup>	×10°	×10 <sup>9</sup>	×10°	×10°	×10°	×10°		Sun
BS04	5.94	1.40	4.88	5.87	5.62	4.90	6.01	0.0169	0.7148
BS04( <sup>14</sup> N)	5.99	1.42	4.89	5.83	3.11	2.38	5.97	0.0170	0.7153
BS05 (OP,GS98)	5.99	1.42	4.84	5.69	3.07	2.33	5.84	0.0170	0.7138
BS05(OP, AGS05)	6.06	1.45	4.34	4.51	2.01	1.45	3.25	0.0126	0.7280
BPSH2011	5.98	1.44	5.00	5.58	2.96	2.23	5.52	0.0170	0.7124
GS98	[6.03]	[1.47]	[4.56]	[4.59]	[2.17]	[1.56]	[3.40]	[0.0134]	[0.7231]
[AGSS09]									

- 1. BS04(<sup>14</sup>N) accounts for new S<sub>1,14</sub> for <sup>14</sup>N( $p,\gamma$ )<sup>15</sup>O [A. Formicola et al 2004]
- 2. BS05(OP,GS98) accounts for new opacities
- BS05(OP,AGS05) accounts for new calculations of surface abundances: Z/X ~ 0.023 -> ~ 0.018

### Solar Standard Model metallicity problem

#### Borexino and SNO solar neutrino measurement of <sup>7</sup>Be and <sup>8</sup>B



## CNO to probe metallicity problem

	<b>CNO</b> FLUX (10 <sup>8</sup> cm <sup>-2</sup> s <sup>-1</sup> )
HIGH Z SSM	5.24 ± 0.84
LOW Z SSM	3.76 ± 0.60
$\Delta \Phi$	28%

## **CNO** interaction rate



Assuming MSW-LMA

# Solar neutrinos survival probability: after Borexino including pep $v_e$ 's



#### The first $v_{\tau}$ candidate event Phys. Lett. B 691 (2010) 138



## EPILOGUE

at the

Three (v) gondolas are safe in the harbor... ...but that's not what they are made for. New gondolas might join, and all lead us towards new (physics) horizons

Thank you for your attention

## Neutrinos: sterile



#### A long standing set of anomalies



#### Summarizing

#### Reactors

0.6 0.7	0.8	0.9	1	1.1	1.2	1.3	1.4			
hilini	m		hhu	ЩШ	ПШ	ΠЦ	uluu			
ROVNO88_3S			H		0.938	+0.006	+0.058			
ROVNO88_2S		-	-		0.959	±0.009	±0.075			
ROVNO88_1S		-	<b>-</b>		0.972	±0.009	±0.076			
ROVN088_21		-			0.948	+0.009	+0.065			
BOVN088_11					0.917	×0.008	+0.063			
SRP-II 23.8 m					1.019	±0.010	+0.038			
SRP-I		-	•		0.953	±0.006	±0.035			
Krasnoyarsk-III 57.3 m			- ·		0.954	±0.010	+0.046			
Krasnoyarsk-II	H	1	-	+	0.960	+0.190	+0.053			
Krasnoyarsk-I		-			0.944	±0.034	+0.062			
ILL 8.76 m	H				0.801	+0.059	+0.048			
Goesgen-III			-		0.924	+0.043	+0.055			
Goesgen-II 46.0 m		-			0.991	±0.024	+0.059			
Goesgen-I		-			0.966	+0.023	+0.058			
Bugey3 95.0 m	+				0.873	±0.115	=0.044			
Bugey3 40.0 m					0.948	×0.009	+0.047			
Bugey-3/4		-	-		0.943	+0.004	+0.047			
ROVNO91		- 1	•		0.940	+0.023	+0.028			
Bugey-3/4					0.943	+0.000	+0.028			
τ <sub>n</sub> =885.7s Av	erage	H	1		0.943	±0.022				
ևուհում	ասհո	пЦ	հետ	uluu	ılın	du	uluu			
0.6 0.7	0.8	0.9	1	1.1	1.2	1.3	1.4			
V <sub>Measured</sub> / V <sub>Expected</sub>										

#### Sources



#### Accelerators



#### $\nu_e$ Disappearance in Gallium radioactive source experiments



 $R \equiv$  wheighted average value of the ratio of measured and predicted  $^{71}Ge$  production rates (p) :

$$R \equiv \frac{p(\text{measured})}{p(\text{predicted})} = 0.87 \pm 0.05$$

arXiv:0901.2200[nucl-ex]

Ga radioactive source exp. results may be interpreted as an indication of the disappearance of  $\nu_e$  due to active-sterile oscillations! hep-ph/0610352 Carlo Giunti & ML





[Giunti, Laveder, PRD 82 (2010) 053005, arXiv:1005.4599]

 $\Delta m^2_{
m SBL} \gtrsim 1\,{
m eV}^2$  is OK, but  $\sin^2 2artheta_
u > \sin^2 2artheta_
u$ 

Parameter Goodness of Fit = 0.2%

CPT violation? [Giunti, Laveder, PRD 82 (2010) 113009, arXiv:1008.4750]

C. Giunti – Sterile Neutrino Fits – 17 Mar 2011 – 22/27

## **Borexino test exp**

► Borexino:

[lanni, Montanino, Scioscia, EPJC 8 (1999) 609, arXiv:hep-ex/9901012]





Interesting Tension (2.3σ difference)
Plan to have at least double current data set by this Summer.

- Data taking interrupted by target failure on Feb. 26
- Plan to have new target in April

Mar. 15, 2011



...does not show up in 2011 data!

## The Reactor Antineutrino Anomaly and implications



Th. Lasserre (CEA-Saclay, Irfu SPP & APC)

#### The 4<sup>th</sup> neutrino hypothesis

Combine all rate measurements, no spectral-shape information
 Fit to anti-v<sub>e</sub> disappearance hypothesis



Absence of oscillations disfavored at 98.6% C.L.



34

#### The neutrino run

- (A) electron-like neutrino data. Comparison between the data (black dots) and the calculated distributions due to misidentified  $v_{\mu}$  events (red)—and genuine  $v_e$  events (blue). The sum is indicated in black. One notices an anomaly at low energies, which is incompatible with LNSD predictions.
- (B) according to Giunti & Laveder scaling of the events is applied with a factor 1.26, within the permitted uncertainty of F =  $1.24 \pm 0.21$  and gives an acceptable fit to the data. The  $v_e$  with and without scaling and disappearance are also shown.





## Intermezzo sulla fisica di Majorana



## Natura del neutrino



#### *Nuovo Cimento* **14** (1937) 171-184

#### TEORIA SIMMETRICA DELL'ELETTRONE E DEL POSITRONE

Nota di Ettore Majorana

Sunto. - Si dimostra la possibilità di pervenire a una piena simmetrizzazione formale della teoria quantistica dell'elettrone e del positrone facendo uso di un nuovo processo di quantizzazione. Il significato delle equazioni di DIRAC ne risulta alquanto modificato e non vi è più luogo a parlare di stati di energia negativa; nè a presumere per ogni altro tipo di particelle, particolarmente neutre, l'esistenza di « antiparticelle » corrispondenti ai « vuoti » di energia negativa.

We show that it is possible to achieve complete formal symmetrization in the electron and positron quantum theory by means of a new quantization process. The meaning of Dirac equations is somewhat modified and it is no more necessary to speak of negative-energy states; nor to assume, for any other type of particles, especially neutral ones, the existence of antiparticles, corresponding to the "holes" of negative energy.

## **Najorana** Neutrino





## laboratori Gran Sasso



Cu cryostat

Ge diodes

ultrapure  $H_2O$ 

#### **Implications of Gallium and Reactor Anomalies**

 $\beta$  Decay

 $(\beta\beta)_{0\nu}$  Decay

₽ 9 99.73% 99.73% Gallium Gallium Reactors œ Reactors œ Tritium Tritium Ga + Re + Tr Ga + Re + Tr 99% 99% ú 9 <sup>2</sup>۲2  ${}^\Delta\!\chi^2$ 4 95.45% 90% 90% 2 2 68.27% 68.27% 0  $10^{-3}$  $10^{-2}$ 10<sup>-1</sup> 10<sup>-2</sup> 10<sup>-1</sup>  $|U_{e4}|^2 \sqrt{\Delta m^2}$  [eV]  $|U_{o4}|\sqrt{\Delta m^2}$ [eV]  $m_{eta}^2 = \sum_k |U_{ek}|^2 m_k^2$  $m_{\beta\beta} = \left| \sum_{L} U_{ek}^2 m_k \right|$ 

[Giunti, Laveder, In Preparation]

C. Giunti – Sterile Neutrino Fits – 17 Mar 2011 – 25/27



#### COMPLEMENTARITY relation :

 $\theta_{12} \sim 32^o \qquad \theta_{es} \sim 13^o \qquad \theta_{12} + \theta_{es} = 45^o$ 

#### A.G. Riess et al. astro-ph 1103.2976



... if they are roses they'll flower...



... A BRIGHT FUTURE for Majorana  $\nu$  physics !!!

#### Two detectors at the CERN-PS neutrino beam



Two positions are foreseen for the detection of the neutrinos The far (T600) location at 850 m from the target: L/E ~ 1 km/GeV; The new location at a distance of 127 m from the target: L/E 0.15 km/GeV

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#### Sensitivity to disappearance anomaly



Sensitivities (90% CL) in the  $\sin^2(2\theta_{new})$  vs.  $\Delta m_{new}^2$  for an integrated intensity of (a) at the 30 kWatt beam intensity of the previous CERN/PS experiments, (b) the newly planned 90 kWatt neutrino beam and (c) a 270 kWatt curve. They are compared (in red) with the "anomalies" of the reactor + Gallex and Sage experiments. A 1% overall and 3% bin-to-bin systematic uncertainty is included (for 100 MeV bins).

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#### Comparing sensitivities (arXiv:0909.0355)



Expected sensitivity for the proposed experiment exposed at the CERN-PS neutrino beam (left) for 2.5 10<sup>20</sup> pot and twice as much for anti-neutrino (right). The LSND allowed region is fully explored both for neutrinos. The expectations from one year of at LNGS are also shown.

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#### The detector layout at FAR/NEAR sites

Liquid Argon is the best option for v measurements in the PS beam energy range (electronic bubble chamber): o full measurement of v interactions  $\circ$  electron  $\pi^0$  discrimination at 10<sup>-3</sup> level  $\circ$  NC rejection at 10<sup>-3</sup> with 90%  $v_e$ identification efficiency Memorandum to SPSC, 09/03/2011

Spectrometer is essential to allow o full Charge ID o systematics control o clean separation nu and anti-nu NESSiE proposal (t.b.s.)

Identical LAr-TPC + spectrometer combination at near site.



**Our Goal** 



(full simulation AND reconstruction)

## Neutrinos: tachionic ?



#### Summary of the principle for the TOF measurement



#### Results

For CNGS  $v_{\mu}$  beam,  $\langle E \rangle = 17$  GeV:

 $\delta t = TOF_c - TOF_v =$ 

 $(1043.4 \pm 7.8 \text{ (stat.)}) \text{ ns} - 985.6 \text{ ns} = (57.8 \pm 7.8 \text{ (stat.)}^{+8.3}\text{-}_{5.9}(\text{sys.})) \text{ ns}$ 

#### relative difference of neutrino velocity w.r.t. c:

 $(v-c)/c = \delta t / (TOF_c - \delta t) = (2.37 \pm 0.32 \text{ (stat.)}^{+0.34} \text{ (sys.)}) \times 10^{-5}$ 

(730085 m used as neutrino baseline from parent mesons average decay point)

#### $6.2 \sigma$ significance

#### Short-bunch wide-spacing neutrino beam





4x10<sup>16</sup> pot accumulated
Proton bunch-length 3ns
35 beam-related events

20 events selected

### Results with the bunched-beam



 $\delta t = 62.1 \pm 3.7 \text{ (stat.)}$ 

The systematic uncertainties are equal or smaller than those affecting the result with the nominal CNGS beam

These result excludes biases affecting the PDF based analysis

## Imaginary mass in matter ?

#### F.Tamburini & ML 1109.5445



## Excellent data vs MC agreement !

F.Tamburini & ML 1111.4441



## **GRAZIE** Milla !!!

