

In fact:
$$U^u(t) = \begin{pmatrix} \cos \theta^u & \sin \theta^u \\ -\sin \theta^u & \cos \theta^u \end{pmatrix}$$

$$\dot{U}^u(t) = \dot{\theta}^u \begin{pmatrix} -\sin \theta^u & \cos \theta^u \\ -\cos \theta^u & -\sin \theta^u \end{pmatrix}$$

$$\begin{aligned} (U^m(t))^{\dagger} \dot{U}^u(t) &= \dot{\theta}^m \begin{pmatrix} C_m & -S_m \\ S_m & C_m \end{pmatrix} \begin{pmatrix} -S_u & C_u \\ -C_u & -S_u \end{pmatrix} \\ &= \dot{\theta}^m(t) \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \end{aligned}$$

and

$$(U^m(t))^{\dagger} H U^u(t) = \begin{pmatrix} E_A & 0 \\ 0 & E_B \end{pmatrix}$$

by definition of θ^u .

The propagator matrix is now diagonal ($\dot{\theta}^u \neq 0$) and ψ_1^m, ψ_2^u are not the energy eigenstates of the Hamiltonian. Mixing between ψ_1^m and ψ_2^u is present and a transition between one and the other is possible.

ADIABATIC APPROXIMATION

If the adiabaticity condition is satisfied:

$$|E_A - E_B| \gg |\dot{\theta}^m(t)|$$

then $\psi_1^m(t)$ and $\psi_2^m(t)$ are \sim energy eigenstates and they do not mix significantly. $P_{\text{jump}} \approx 0$ and the evolution is adiabatic.

To be quantitative:

$$\begin{aligned} \dot{\theta}^m(t) &= \frac{d\theta^m}{dt} \cdot \frac{d \tan 2\theta^m}{dt} \\ &= \frac{1}{2} \cos^2 2\theta^m \cdot \frac{\frac{\Delta m^2}{2E} \sin 2\theta_0 \dot{V}}{\left(\frac{\Delta m^2}{2E} \cos 2\theta_0 - V(t) \right)^2} \\ &= \frac{1}{2} \cos^2 2\theta^m \cdot \tan^2 2\theta^m \cdot \frac{\dot{V}}{\frac{\Delta m^2}{2E} \sin 2\theta_0} \\ &= \frac{1}{2} \sin^2 2\theta^m \cdot \frac{\dot{V}}{\frac{\Delta m^2}{2E} \sin 2\theta_0} \\ &= \frac{1}{2} \sin^2 2\theta^m \cdot \frac{\dot{V}}{V} \cdot \frac{V}{\frac{\Delta m^2}{2E} \sin 2\theta_0} \end{aligned}$$

The condition is more restrictive at resonance where $|E_A - E_B|$ is minimal and $\sin^2 2\theta^m$ is maximal ($\sin^2 2\theta^m = 1$).

At resonance, the adiabaticity condition becomes:

$$|E_A - E_B| \gg |\dot{\theta}^m(t)|$$

$$\frac{\Delta m^2}{2E} \sin 2\theta_0 \gg \frac{1}{2} \left| \frac{\dot{V}}{V} \right| \frac{\cos 2\theta_0}{\sin 2\theta_0}$$

$$\frac{\Delta m^2}{2E} \frac{\sin^2 2\theta_0}{\cos 2\theta_0} \left| \frac{V}{\dot{V}} \right| \gg 1$$

with $\frac{\dot{V}}{V} = \frac{1}{V} \frac{dV}{dr} = \frac{d \ln V}{dr}$

Notice that the adiabaticity condition depends on θ_0 .

One can also understand this as follows:

$$\text{resonance width} \equiv \delta r_R = \frac{2 \tan 2\theta_0}{\left| \frac{1}{V} \frac{dV}{dr} \right|}$$

$$\frac{\Delta m^2}{2E} \sin 2\theta_0 \cdot 2 \tan 2\theta_0 \frac{1}{\left| \frac{1}{V} \frac{dV}{dr} \right|} \delta r_R \gg 1$$

$(E_A - E_B)_R$

$$\frac{2\pi \delta r_R}{L_R} \gg 1$$

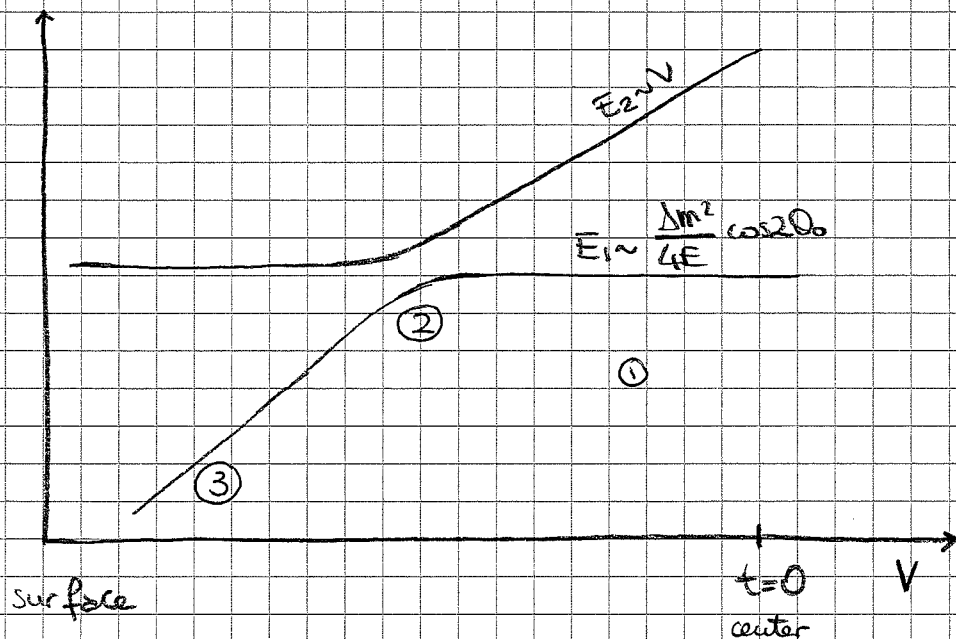
Many oscillations need to take place at resonance.

MSW effect

Mikheyev, Smirnov, Wolfenstein

Let's consider a neutrino produced in the Sun center ($V \gg \frac{\Delta m^2}{2E} \cos 2\theta_0$) which travels towards the surface.

We plot the energy eigenvalues as a function of $V(t)$:



$$E_{1,2} = \frac{V \pm \sqrt{\left(V - \frac{\Delta m^2}{2E} \cos 2\theta_0\right)^2 + \left(\frac{\Delta m^2}{2E} \sin 2\theta_0\right)^2}}{2}$$

$$\textcircled{1} \quad E_1 \sim \frac{\Delta m^2}{4E} \frac{\cos 2\theta_0}{\cancel{V}} \cdot \cancel{V} \sim \frac{\Delta m^2}{4E} \cos 2\theta_0$$

$$E_2 \sim V$$

$$\textcircled{2} \quad \text{Resonance:} \quad E_1 \sim \frac{\Delta m^2}{2E} \frac{\cos 2\theta_0 - \sin 2\theta_0}{2}$$

$$E_2 \sim \frac{\Delta m^2}{2E} \frac{\cos 2\theta_0 + \sin 2\theta_0}{2}$$

$$\textcircled{3} \quad E_{1,2} = \pm \frac{\Delta m^2}{2E}$$

In the center of the Sun, the dominant term in the H^m is $V \Rightarrow \nu_e$ is mainly aligned with the heaviest eigenstate ν_2^m . ν_2^m evolves and (if adiabatic) becomes ν_2^o at the surface. Then the probability of ν_e survival is:

~~$$P_{ee}(\text{Earth}) = |\langle \nu_e | \nu_2 \rangle|^2 = \sin^2 \theta_0$$~~

$$P_{ee}(\text{Earth}) = |\langle \nu_e | \nu_2 \rangle|^2 = \sin^2 \theta_0$$

For small values of θ_0 , one can have a large depletion of the solar flux.

We can study this further. The $|\nu_e\rangle$ state in the Sun core is produced as an incoherent superposition of $|\nu_1^m\rangle$ and $|\nu_2^m\rangle$ as large E-difference.

So

$$|\nu_e\rangle \rightarrow |\nu_1^m\rangle \quad \text{with } P_1 = \cos^2 \theta_{\text{Sun}}$$

$$|\nu_2^m\rangle \quad \text{with } P_2 = \sin^2 \theta_{\text{Sun}}$$

We also introduce a jump probability for transitions between ν_1^m and ν_2^m , P_c .

$$|\nu_1^m\rangle \rightarrow |\nu_1\rangle \quad \text{with } P = 1 - P_c$$

$$|\nu_2\rangle \quad \text{with } P = P_c$$

$$|\nu_2^m\rangle \rightarrow |\nu_1\rangle \quad \text{with } P = P_c$$

$$|\nu_2\rangle \quad \text{with } P = 1 - P_c$$

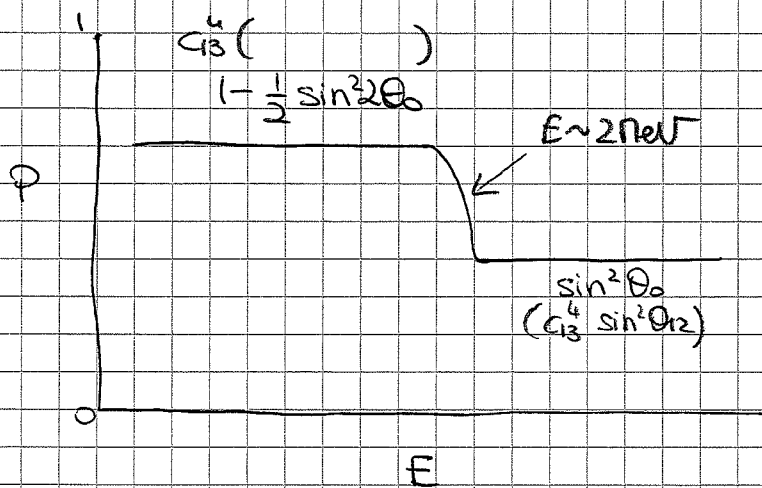
Finally, $P_{1e} = |\langle \nu_e | \nu_1 \rangle|^2 = \cos^2 \theta_0$

$P_{2e} = |\langle \nu_e | \nu_2 \rangle|^2 = \sin^2 \theta_0$

Then

$$P_{ee}^{\text{Sun}} = \cos^2 \theta_{13} \left((1 - P_c) \cos^2 \theta_0 + P_c \sin^2 \theta_0 \right) + \sin^2 \theta_{13} \left(P_c \cos^2 \theta_0 + (1 - P_c) \sin^2 \theta_0 \right)$$

For the simplified case $P_c = 0$, one has



Fits very well the spectral data with $\sin^4 \theta_0 \approx 0.3$,
 $\Delta m^2 \sim 10^{-5} - 10^{-4} \text{ eV}^2$.

Present knowledge in ν physics

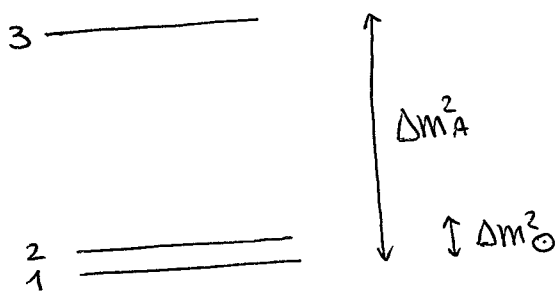
From atm, solar, reactor and accelerator ν exp,
we know

$$\Delta m_{\odot}^2 = 8.0 \times 10^{-5} \text{ eV}^2$$

$$\Delta m_A^2 = 2.5 \times 10^{-3} \text{ eV}^2$$

\Rightarrow 3- ν mixing.

ν masses can be arranged in two possible ways:

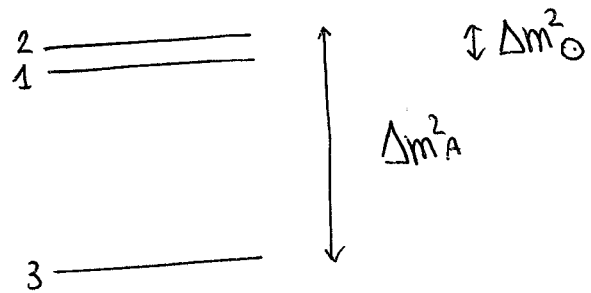


NORMAL ORDERING

$$m_1 = m_{\text{MIN}}$$

$$m_2 = \sqrt{m_{\text{MIN}}^2 + \Delta m_{\odot}^2}$$

$$m_3 = \sqrt{m_{\text{MIN}}^2 + \Delta m_A^2}$$



INVERTED ORDERING

$$m_1 = \sqrt{m_{\text{MIN}}^2 + \Delta m_A^2 - \Delta m_{\odot}^2}$$

$$m_2 = \sqrt{m_{\text{MIN}}^2 + \Delta m_A^2}$$

$$m_3 = m_{\text{MIN}}$$

In order to determine the absolute ν masses,

- absolute mass scale, m_{MIN}
- type of ordering (hierarchy).

It is conventional to consider 3 different types of ν mass spectrum:

- NH : $m_1 \ll m_2 \ll m_3$

$$m_2 = \sqrt{\Delta m_{21}^2}$$

$$m_3 = \sqrt{\Delta m_{31}^2}$$

- IH : $m_3 \ll m_1 \simeq m_2$

$$m_1 \simeq \sqrt{\Delta m_{31}^2 - \Delta m_{21}^2}$$

$$m_2 \simeq \sqrt{\Delta m_{31}^2}$$

- QD : $m_1 \simeq m_2 \simeq m_3 \gg \Delta m_{21}^2, \Delta m_{31}^2$

Summary of angles: $\sin^2 \theta_{12} = 0.30, \sin^2 \theta_{13} = 0.05,$
 $\sin^2 \theta_{23} < 0.04.$

How can we express ν mass in the diagonal and then explain their origin?