

Current and Future Reactor Based Neutrino Experiments

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RCvS

Tohoku University

Physics of Massive Neutrinos
@Milos Island
May 20, 2008

Χυρρεντ ανδ Φυτυρε Ρεαχτορ Βασεδ Νευτρινο Εξπεριμεντο

Φ.Συεκανε

ΡΧνΣ

Τοηοκυ Υνιπερσιτυ

Πηψσιχο οφ Μασσιπε Νευτρινοσ
@Μιλοσ Ισλανδ
Μαψ 20, 2008

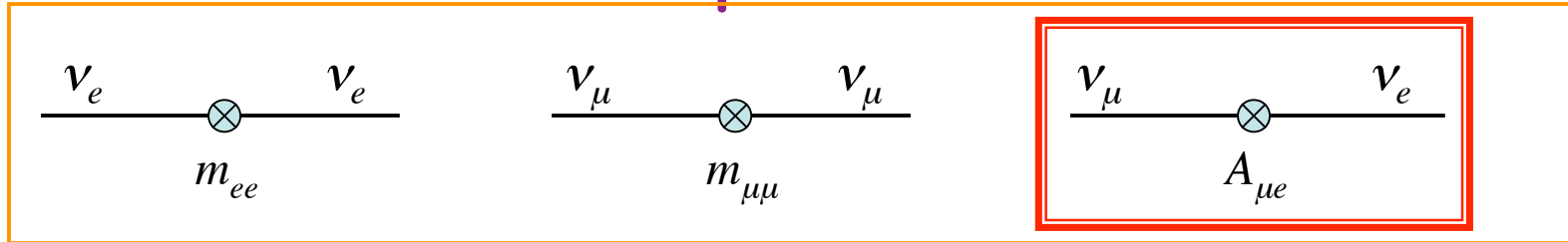
Contents of this talk

- Quick Review of ν Oscillation
- Scope of Reactor ν experiments
- $\theta_{12}, \Delta m^2_{12}$: KamLAND
- θ_{13} : DoubleChooz, Dayabay, RENO
- Future prospects;
 Precise θ_{12} , Very Precise $\theta_{13}, \Delta m^2_{13}$
- Summary

A Quick review of ν Oscillation

Flavor Transition Amplitudes

Charged lepton=mass eigenstate
Simplified view.



ν equation of motion:

$$i \frac{d}{dt} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} m_{ee} & A_{\mu e} \\ A_{\mu e} & m_{\mu\mu} \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix}$$

Definition of mass eigenstate:

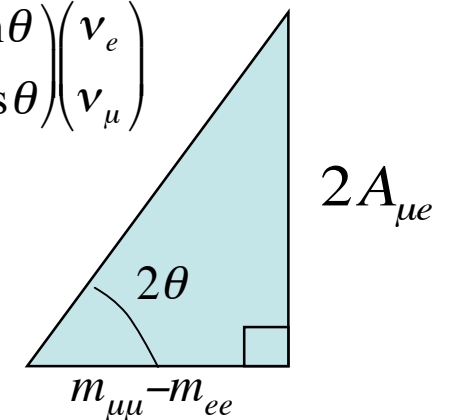
$$i \frac{d}{dt} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = \begin{pmatrix} m_1 & 0 \\ 0 & m_2 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$$\begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix}$$

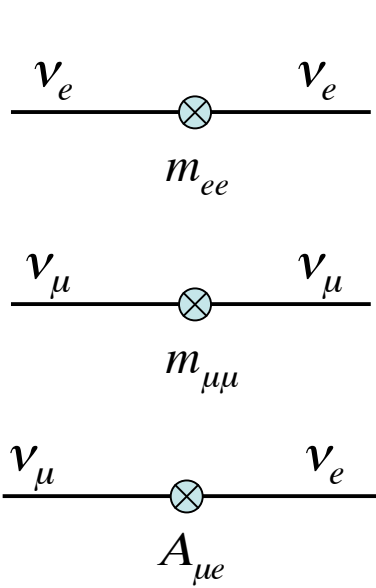
Then ν Oscillation takes place:

$$(m_1, m_2 \ll E) \quad P_{\nu_e \rightarrow \nu_\mu} = \sin^2 2\theta \sin^2 \frac{(m_2^2 - m_1^2)L}{4E}$$

$$2m_i = (m_{\mu\mu} + m_{ee}) \pm \sqrt{(m_{\mu\mu} - m_{ee})^2 + 4A_{\mu e}^2}$$



What We Measure by ν Oscillation

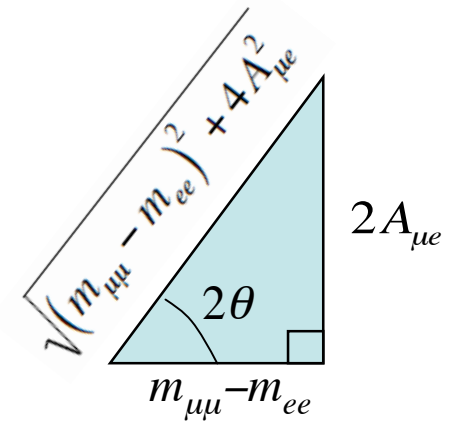


$$\left\{ \begin{aligned} \sin 2\theta &= \frac{2A_{\mu e}}{\sqrt{(m_{\mu\mu} - m_{ee})^2 + 4A_{\mu e}^2}}, \\ \Delta m_{12}^2 &= (m_{\mu\mu} + m_{ee}) \sqrt{(m_{\mu\mu} - m_{ee})^2 + 4A_{\mu e}^2} \end{aligned} \right.$$

OR

$$\left\{ \begin{aligned} A_{\mu e} &= \Delta m_{12}^2 \sin 2\theta / \langle m \rangle \\ m_\mu &= \langle m \rangle + \Delta m_{12}^2 \cos 2\theta / \langle m \rangle \\ m_e &= \langle m \rangle - \Delta m_{12}^2 \cos 2\theta / \langle m \rangle \end{aligned} \right. \quad \langle m \rangle = \frac{m_1 + m_2}{2}$$

↑
Direct mass measurement

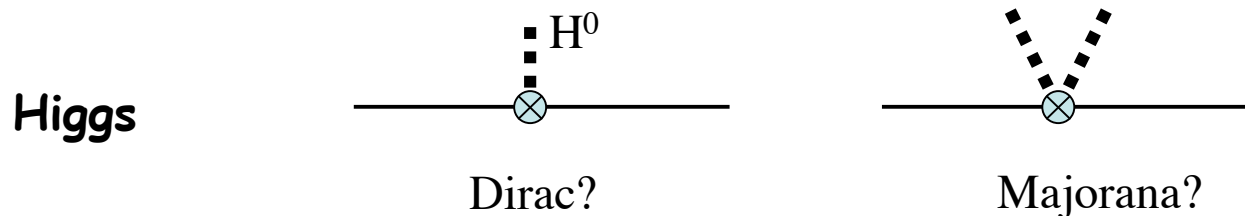


Both Mass and Mixing are a combination of flavor transition amplitude.
 ⇒ Measurement of mixing angle is as important as measurement of mass.

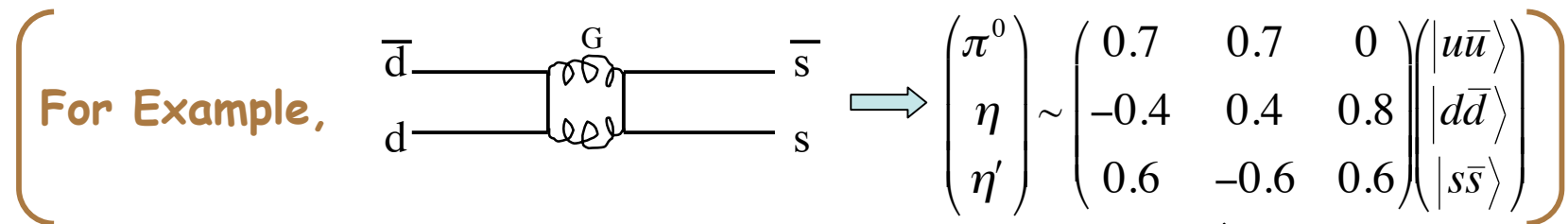
Purpose of ν Oscillation Measurements

Now we know $\nu_\alpha \xrightarrow{\otimes} \nu_\beta$ exists.

What makes the transition amplitudes?



Sub Structure??



Or something else?? A_{NP} ?

Similar to
MNS matrix!

\Rightarrow Physics of ν oscillation is to measure the flavor transition amplitudes and think of its origin.

ν Oscillations: 3 flavor case

Mixings

MNS Matrix

$$s_{ij} = \sin\theta_{ij}, \quad c_{ij} = \cos\theta_{ij}$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{-i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

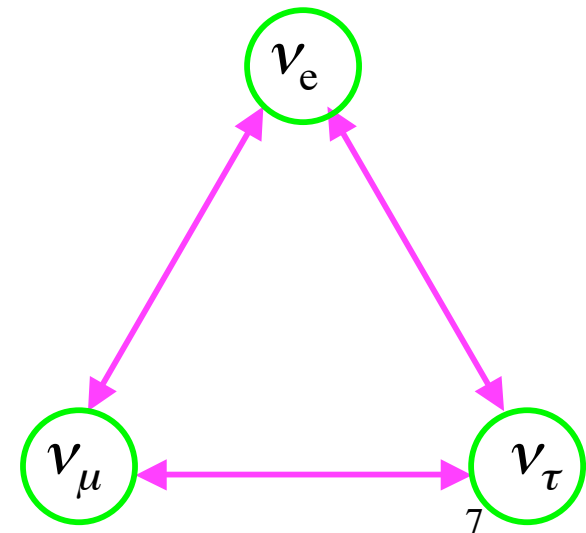
Oscillations

$$\begin{aligned} P(\nu_\alpha \rightarrow \nu_\beta) &= \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2 \Phi_{ij} \mp 2 \sum_{i>j} \text{Im}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin 2\Phi_{ij} \\ P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) & \end{aligned}$$

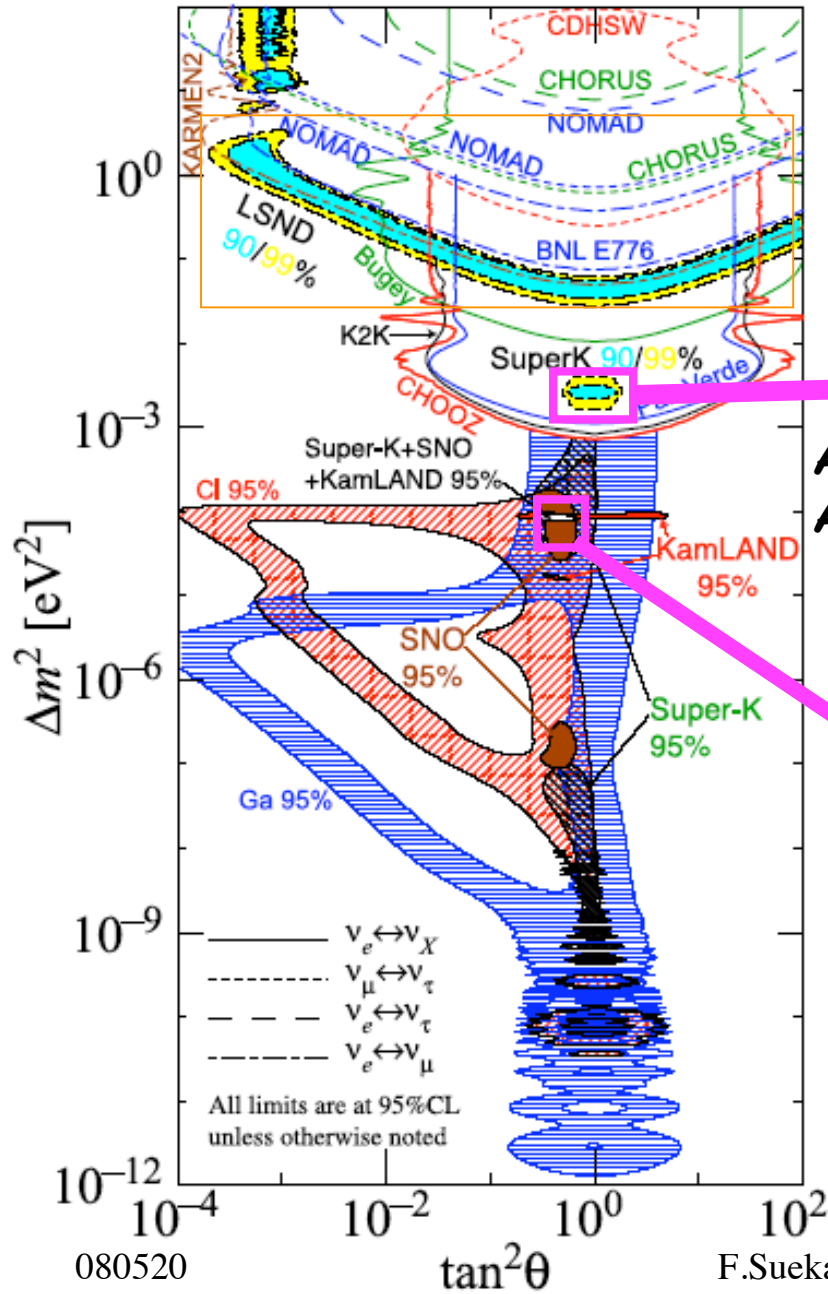
$$\left(\Phi_{ij} = \frac{\Delta m_{ij}^2 L}{4E}, \quad \Delta m_{ij}^2 = m_j^2 - m_i^2 \right)$$

$$\boxed{|\Delta m_{12}^2|, |\Delta m_{23}^2|, \theta_{12}, \theta_{23}, \theta_{31}, \delta}$$

6 parameters can be accessible from neutrino oscillation.

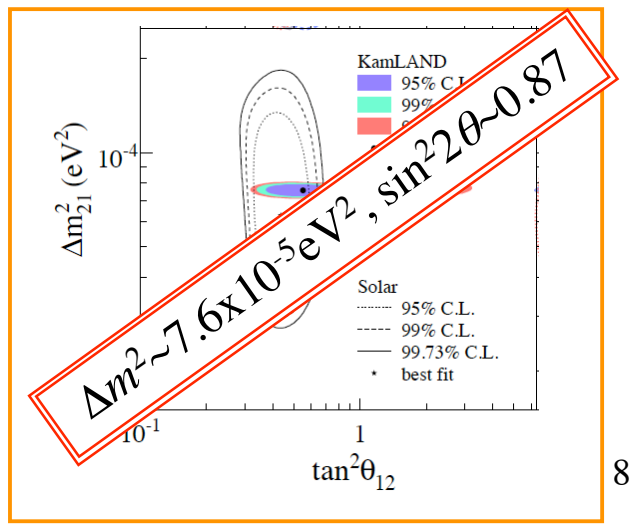
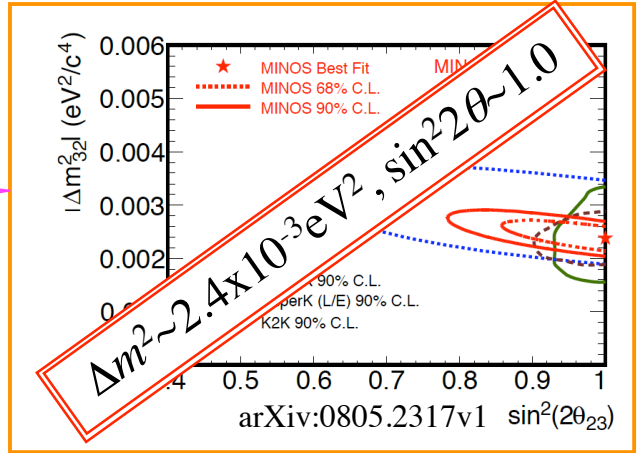


Two oscillations measured



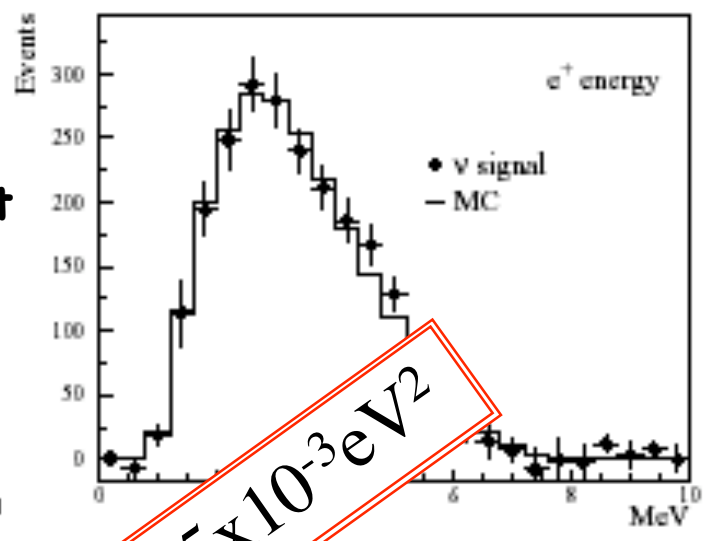
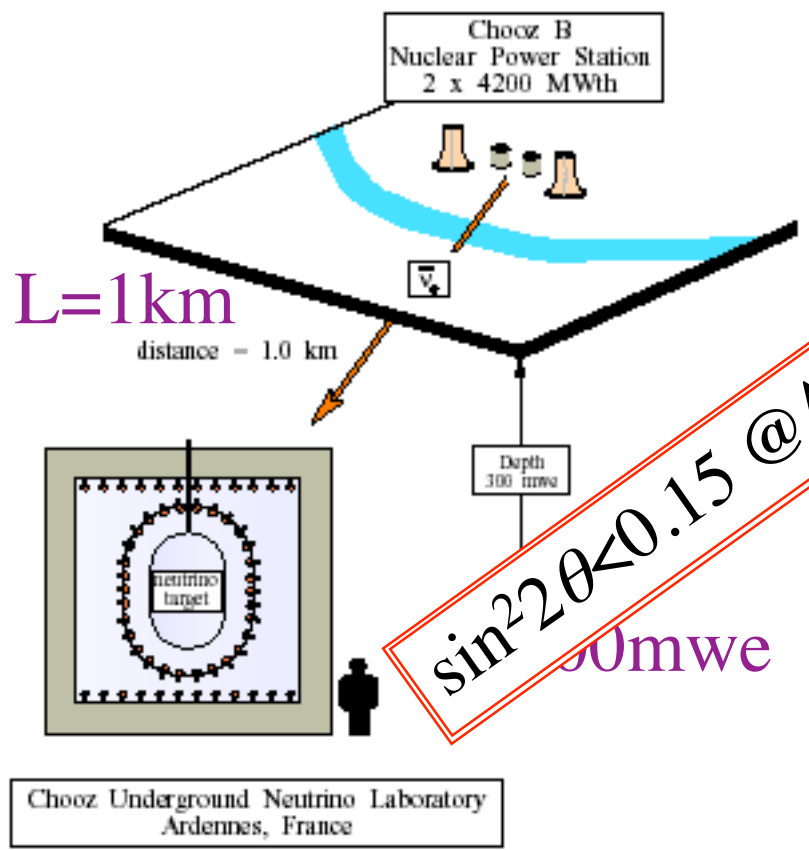
Atmospheric Accelerator

Solar Reactor

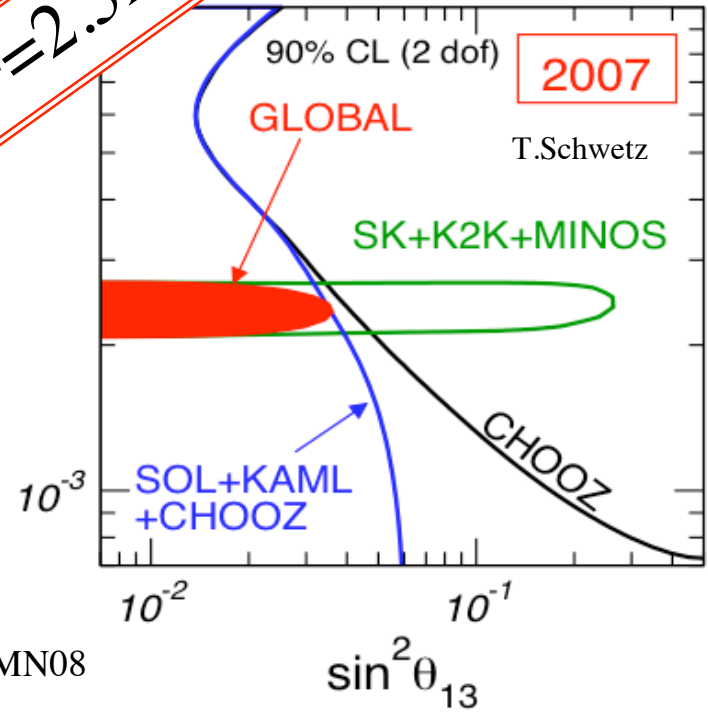


Upper limit

CHOOZ reactor ($\bar{\nu}_e \rightarrow \bar{\nu}_e$) experiment



$\sin^2 2\theta < 0.15$ @ $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$

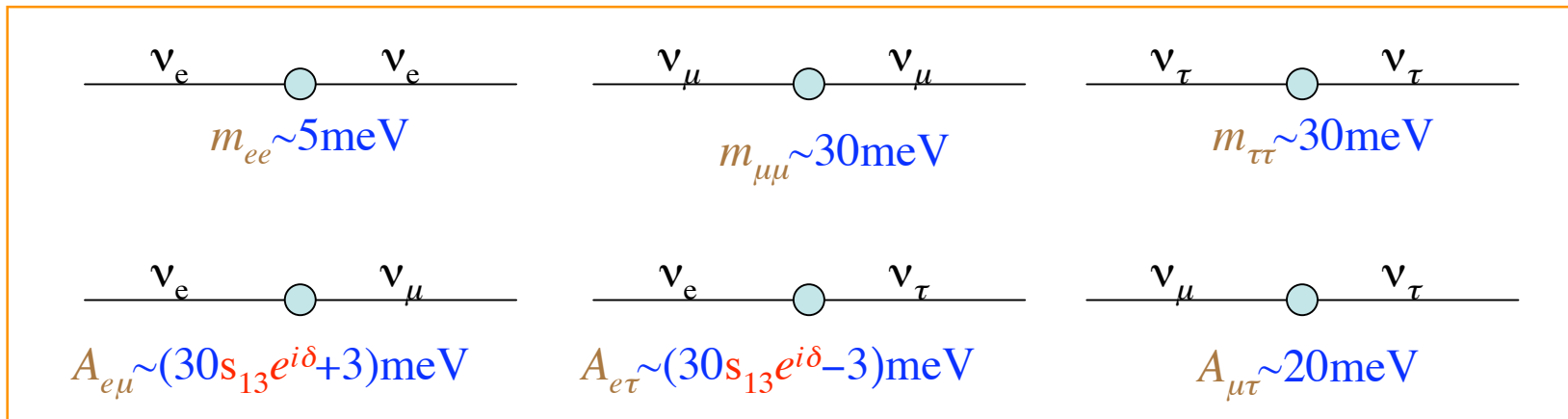


Our Current Knowledge

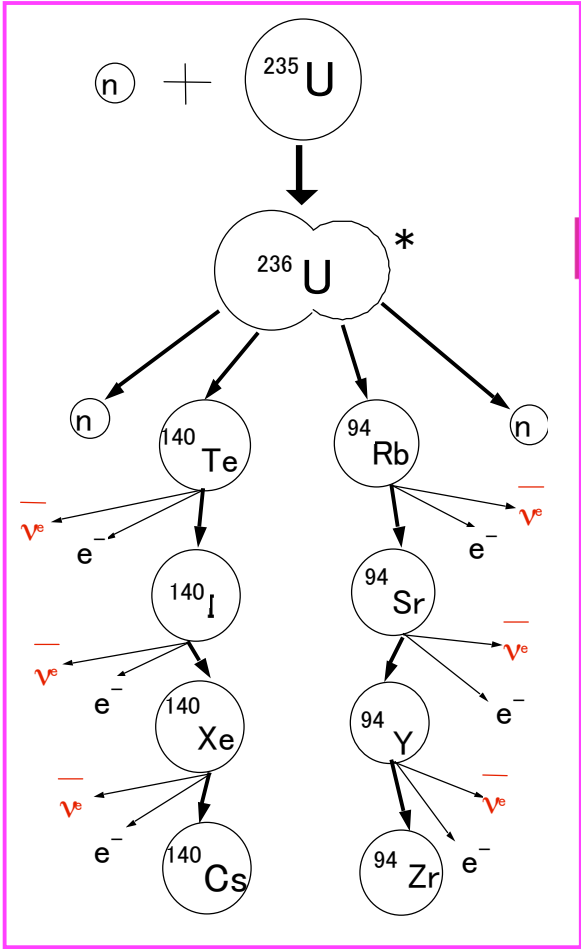
$$|m_3^2 - m_2^2| \sim 2.5 \times 10^{-3} \text{ eV}^2, \quad (m_2^2 - m_1^2) \sim 8 \times 10^{-5} \text{ eV}^2$$

$$U_{MNS} \sim \begin{pmatrix} 0.8 & 0.5 & s_{13} e^{i\delta} \\ -0.4 & 0.6 & 0.7 \\ 0.4 & -0.6 & 0.7 \end{pmatrix} \quad |s_{13}| < 0.2$$

If $m_3 > m_2 \gg m_1 \sim 0$, Flavor Transition Amplitudes become,



Reactor Neutrinos

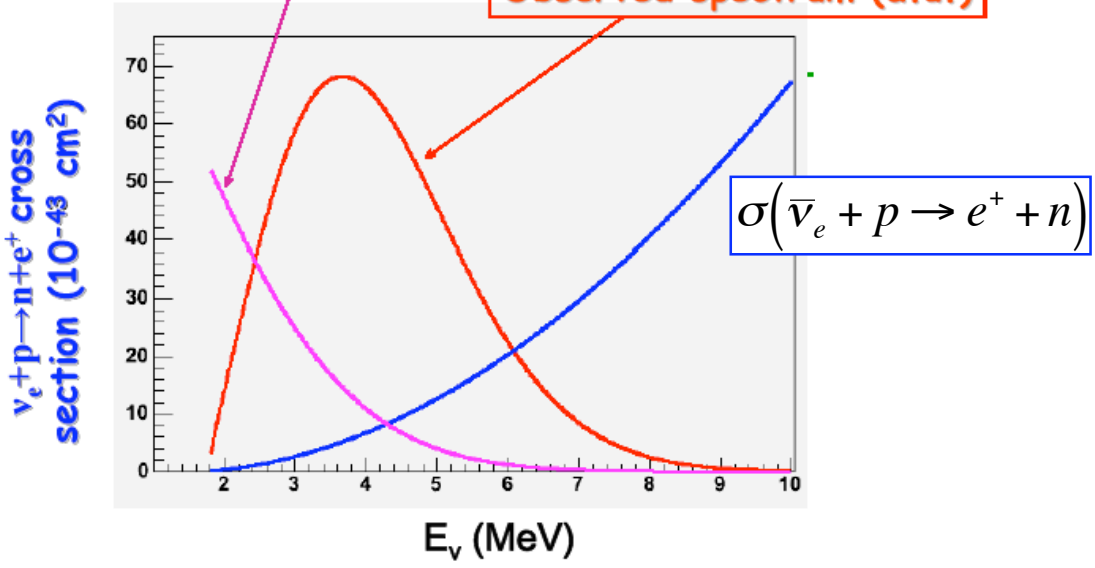


$\sim 6 \times 10^{20} \bar{\nu}_e / s / reactor$

The $\bar{\nu}_e$ energy spectrum

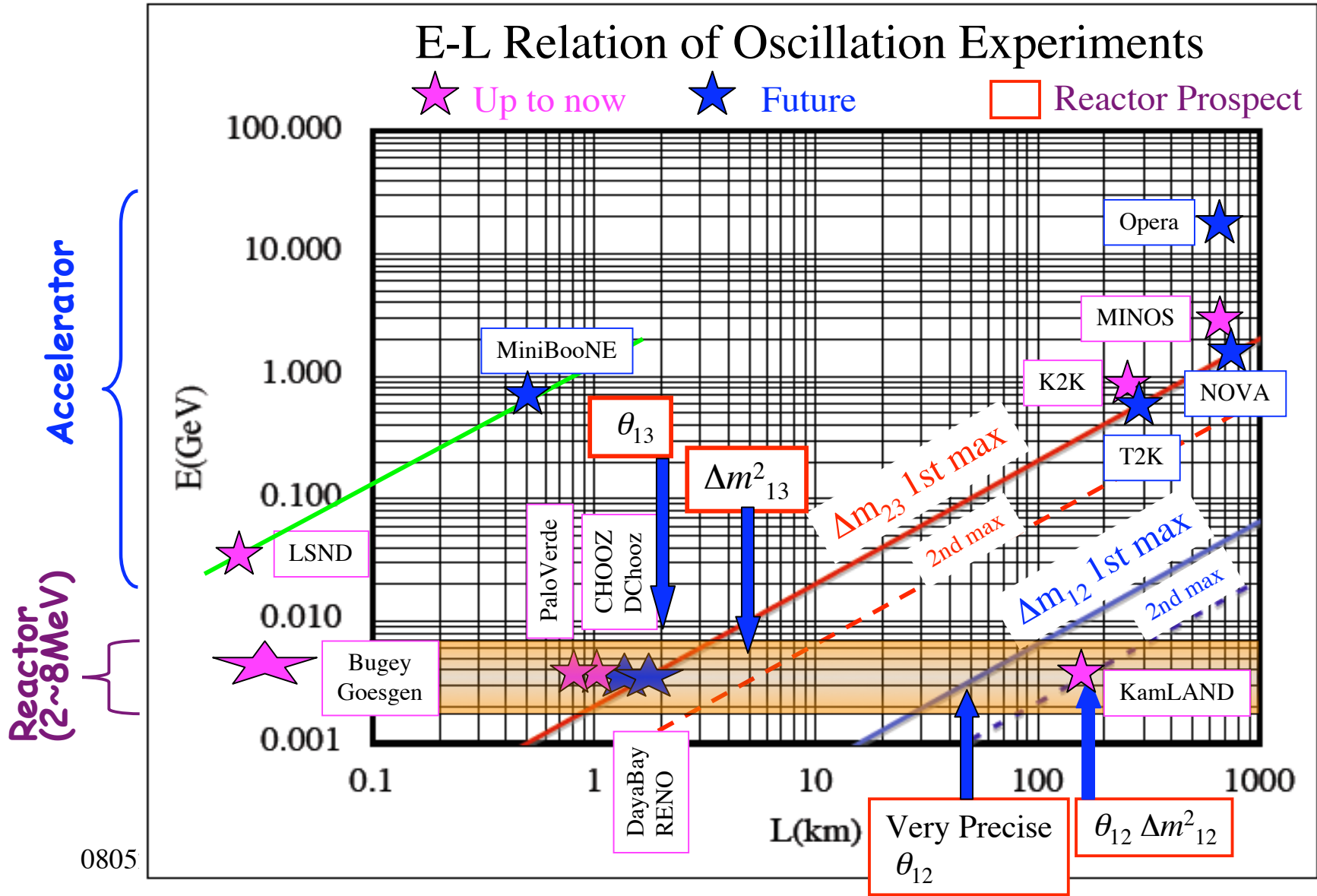
Reactor $\bar{\nu}_e$ spectrum (a.u.)

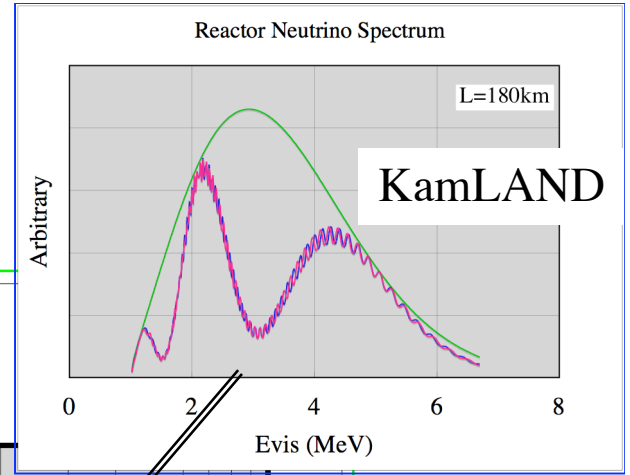
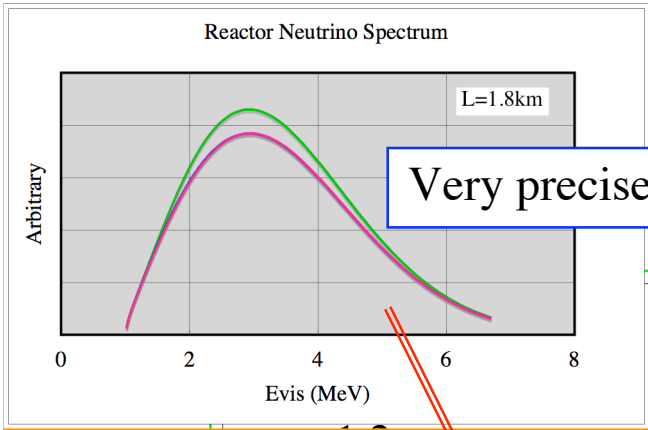
Observed spectrum (a.u.)



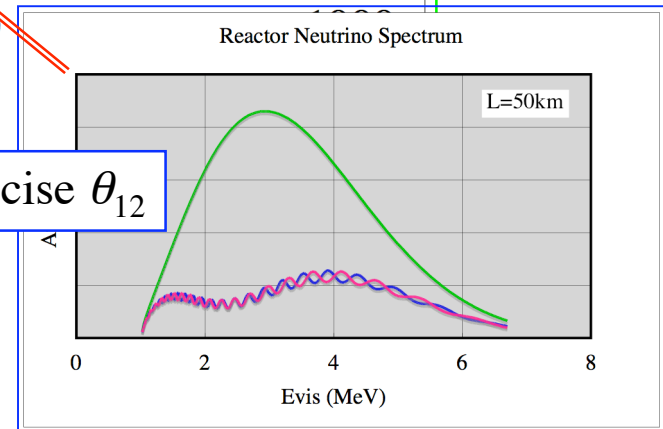
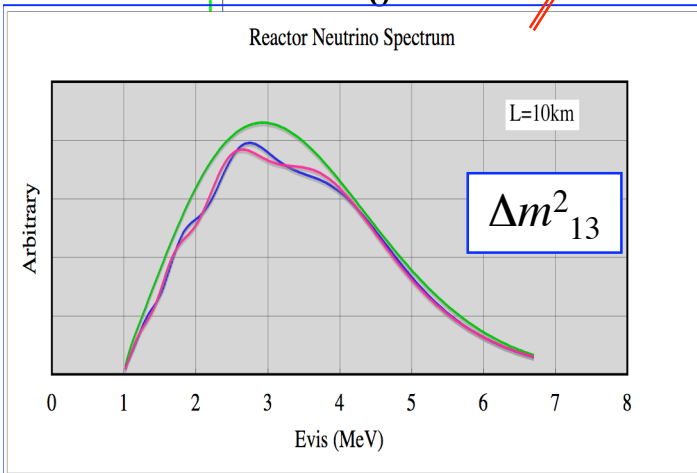
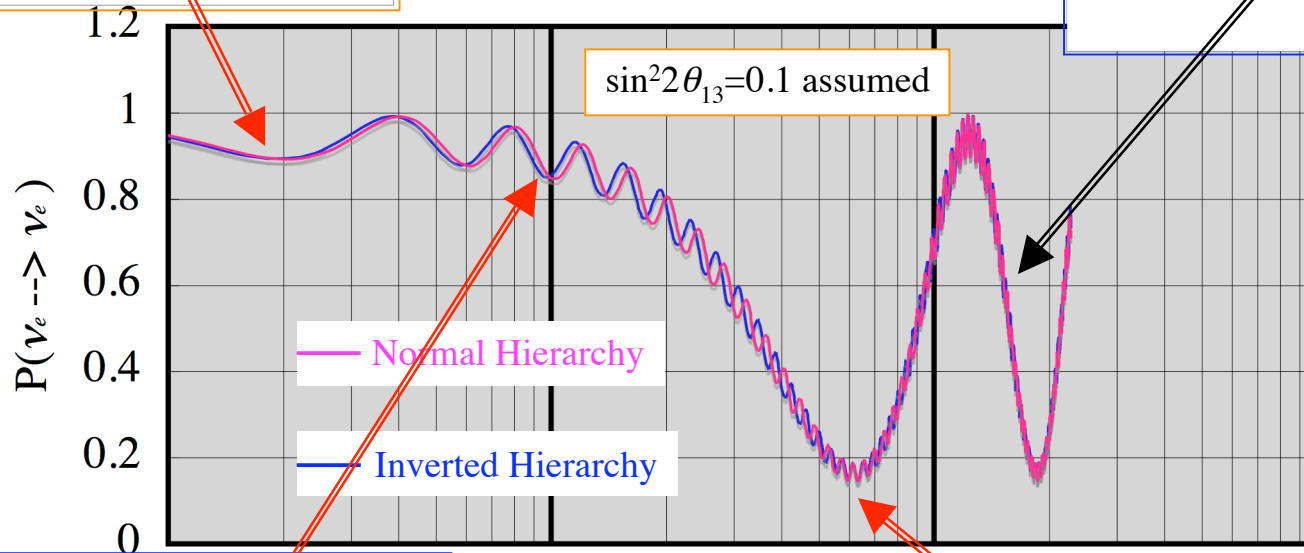
$$E_\nu \sim 4_{-2}^{+4} \text{ MeV}$$

Accessible Oscillations by Reactor ν



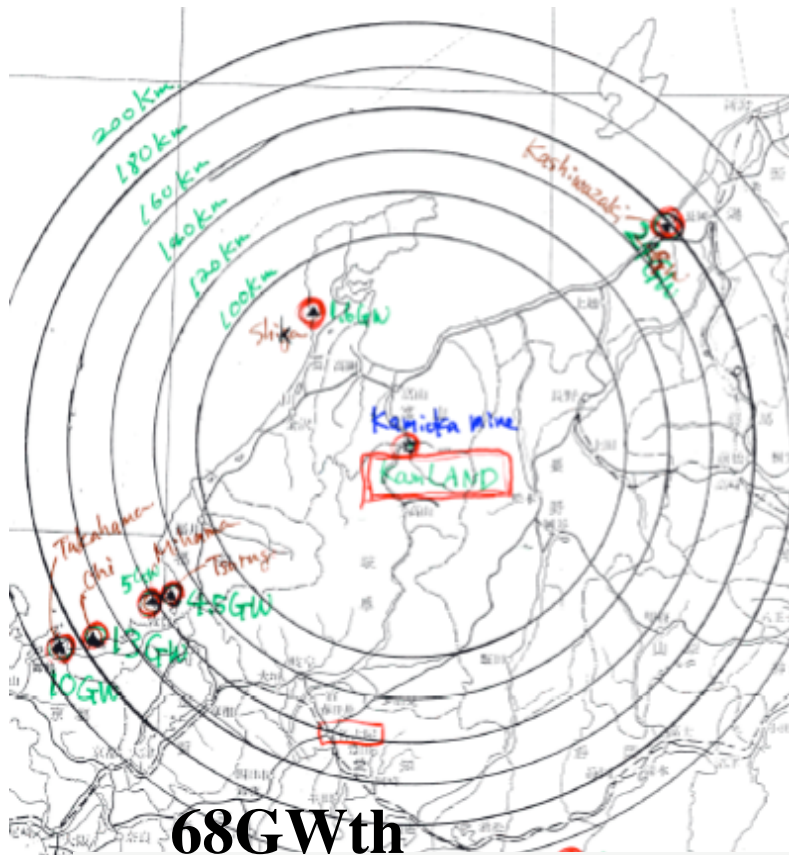


Reactor Neutrino Oscillation

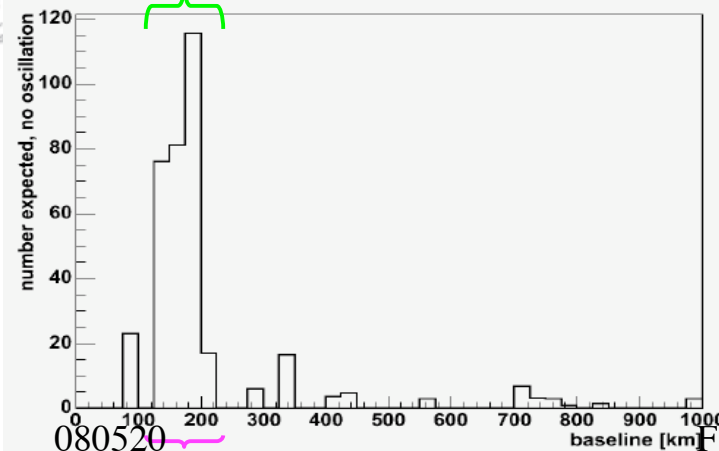


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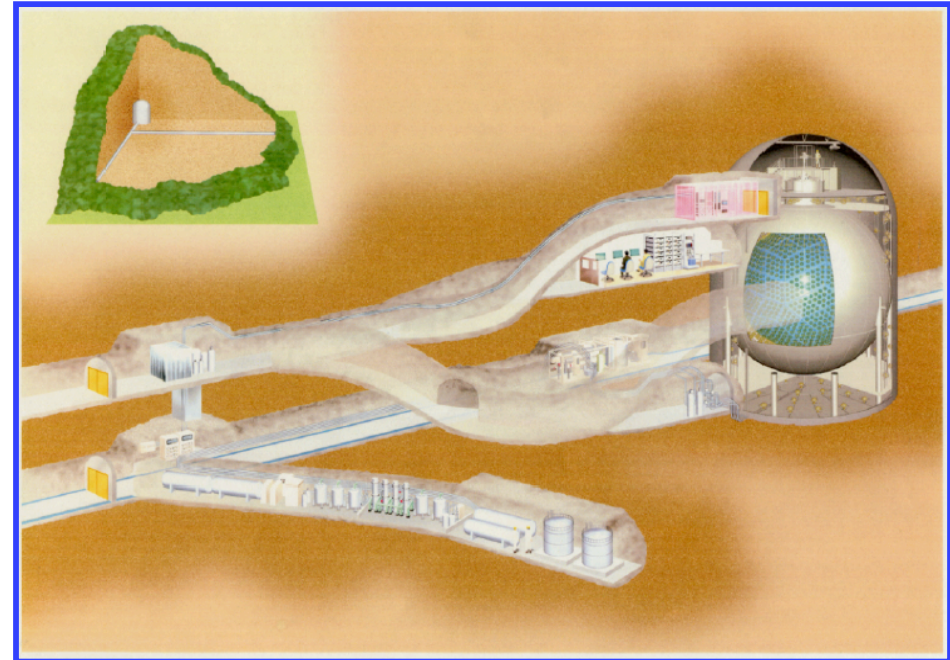
$\Delta m^2_{12}, \theta_{12}$; KamLAND



68GWth



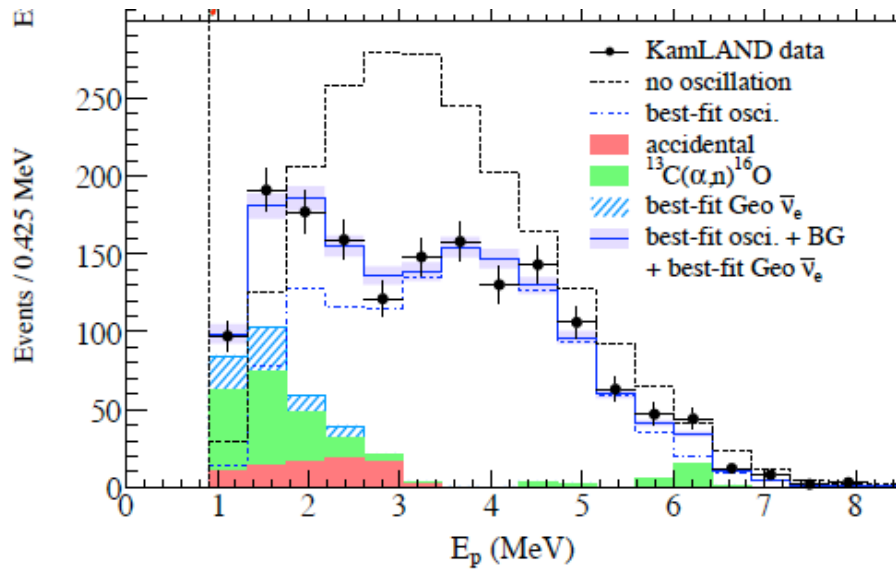
L=180 35km



2002/12: Evidence of ν_e Deficit
(PRL 90:021802, 2003)

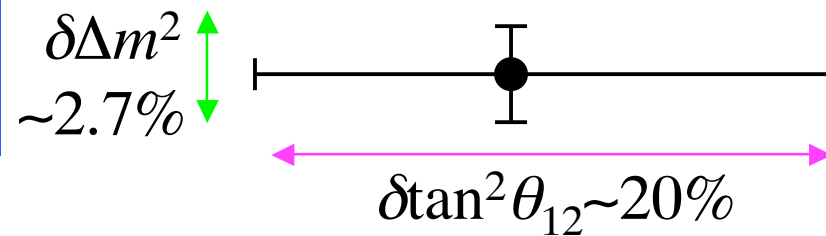
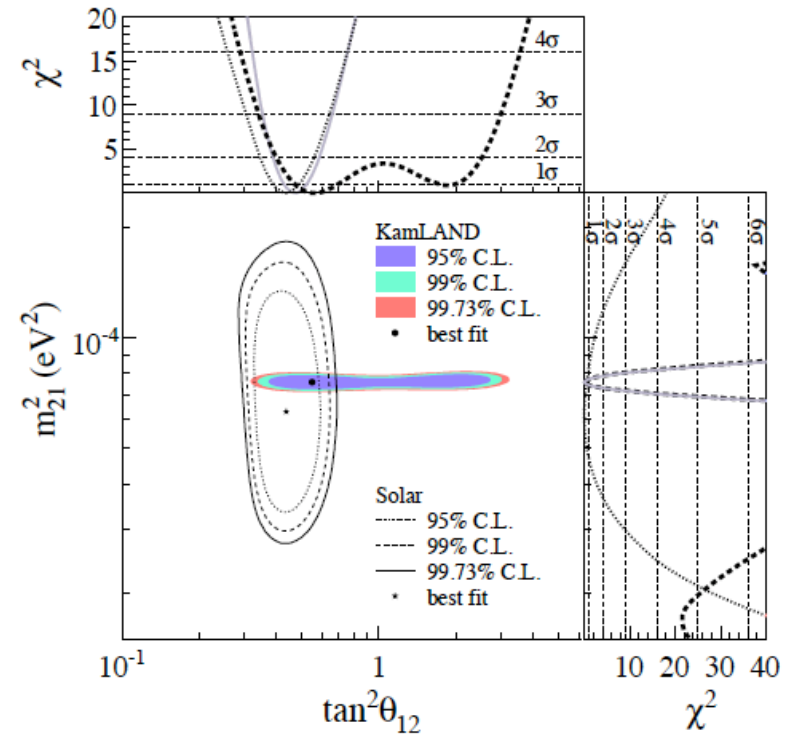
2004/06: Evidence of spectrum distortion
(PRL 94,081801,2005)

2008/01: Precise measurement of Δm^2_{12} & $\sin^2 2\theta_{12}$
(arXiv:0801.4589, submitted to PRL)



2008.1.30arXiv:0801.4589v1

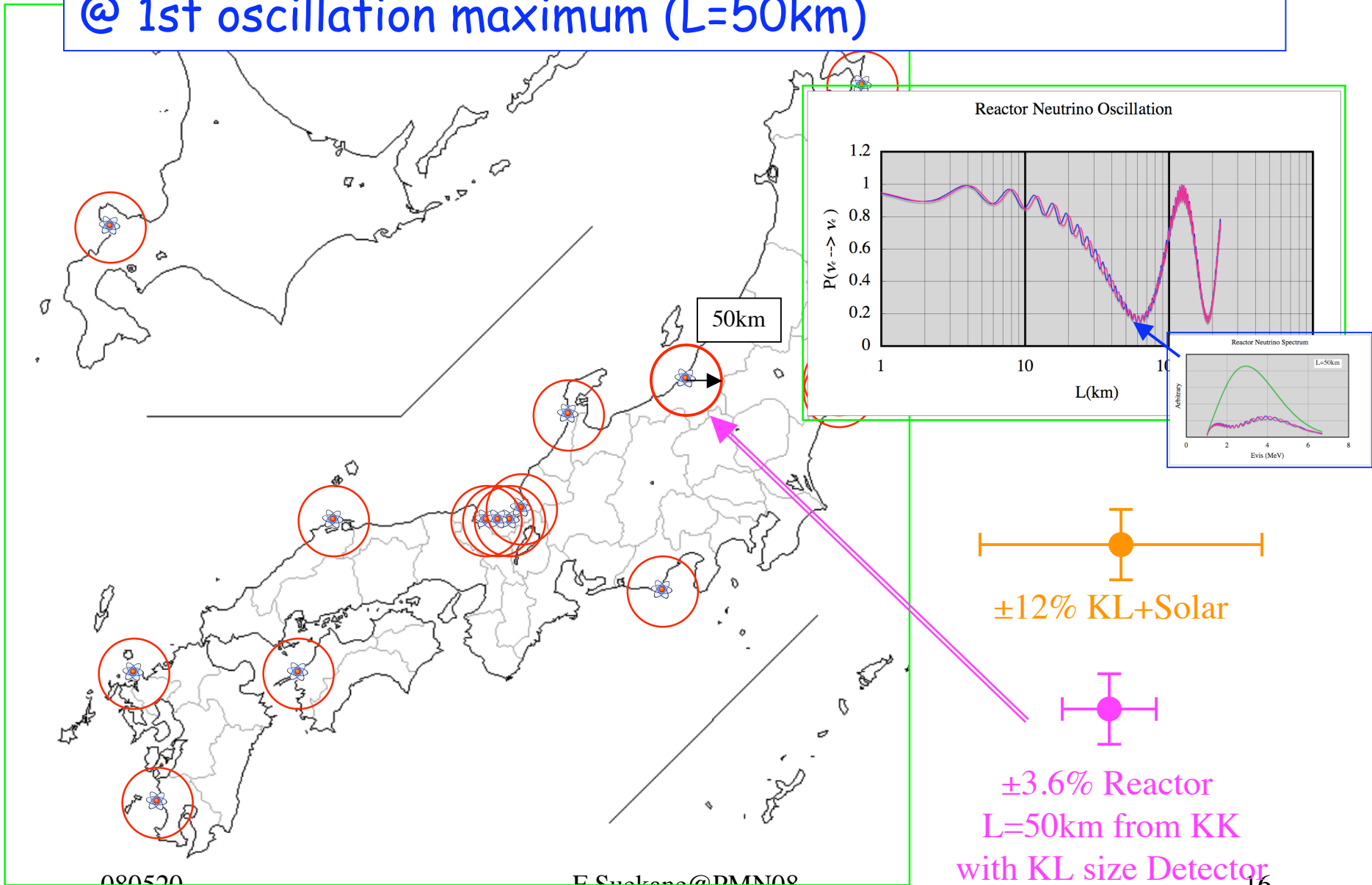
$$\begin{cases} \Delta m_{21}^2 = 7.58^{+0.14}_{-0.13} (\text{stat.})^{+0.15}_{-0.15} (\text{syst.}) \times 10^{-5} \text{ eV}^2 \\ \tan^2 \theta_{12} = 0.56^{+0.10}_{-0.07} (\text{stat.})^{+0.10}_{-0.06} (\text{syst.}) \end{cases}$$



KamLAND keeps taking reactor data while purification. However, KL has already run 5 years and significant improvement of the accuracy from here will be difficult.

On the other hand, precise θ_{12} is needed to understand the sun.

Precise $\tan^2\theta_{12}$ measurement can be performed
 @ 1st oscillation maximum ($L=50\text{km}$)



$$\theta_{13}$$

$$A_{e\mu} \sim (30 \sin \theta_{13} e^{i\delta} + -3) \text{meV}$$

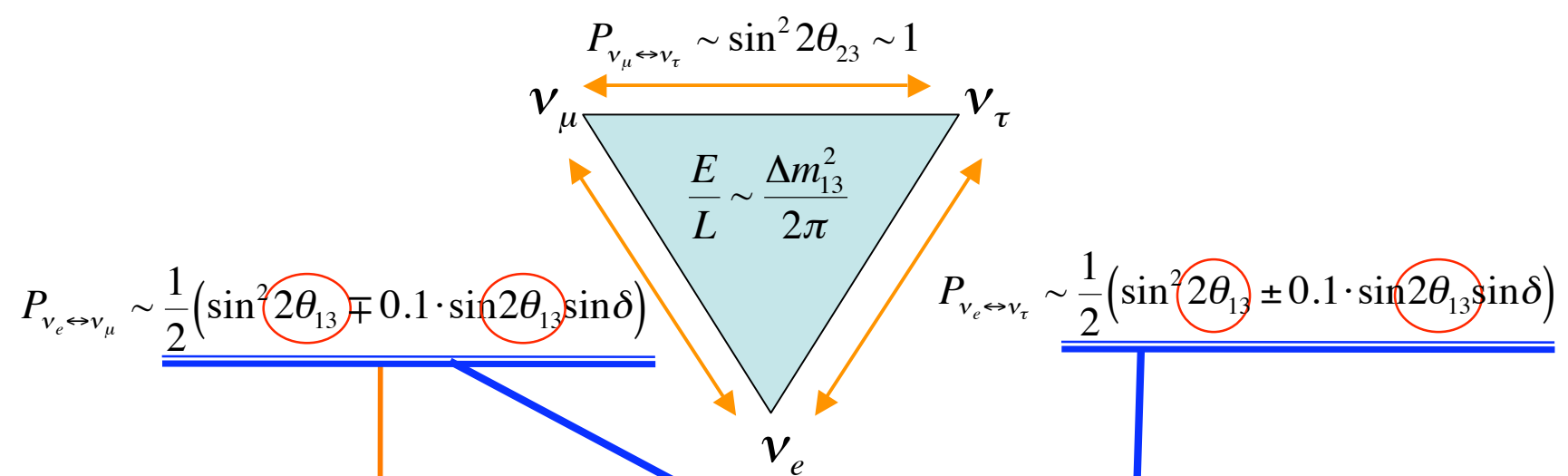
Remaining Issue	How to measure
θ_{13}	$[\bar{\nu}_e \rightarrow \bar{\nu}_e]_R = 1 - \sin^2 2\theta_{13}$ $[\nu_\mu \rightarrow \nu_e]_A \sim \sin^2 \theta_{23} \sin^2 2\theta_{13} \mp 0.05 \cdot \sin \theta_{13} \sin \delta$
δ	$[\nu_\mu \rightarrow \nu_e]_A - [\bar{\nu}_\mu \rightarrow \bar{\nu}_e]_A \sim \sin 2\theta_{13} \sin \delta$
θ_{23} degeneracy	$[\nu_\mu \rightarrow \nu_e]_A \sim \sin^2 \theta_{23} \sin^2 2\theta_{13} \mp 0.05 \cdot \sin \theta_{13} \sin \delta$
Mass hierarchy	$\text{Matter Effect} \sim 0.00017 L[\text{km}] \cdot \text{sign}(\Delta m_{23}^2) \sin^2 2\theta_{13}$
Precise θ_{12}	$[\bar{\nu}_e \rightarrow \bar{\nu}_e]_{\text{KamLAND}} = \cos^4 \theta_{13} (1 - \sin^2 2\theta_{12})$

All the measurements are related to θ_{13}
=> Determination of θ_{13} is urgent

How to Measure θ_{13}

$2.4 \times 10^{-3} \text{eV}^2$

$$\left. @ \frac{\Delta m_{13}^2 L}{4E} \sim \frac{\pi}{2} \right\} \begin{array}{l} E \sim \text{MeV}, L \sim 1 \text{km} \\ E \sim \text{GeV}, L = 100 \sim 1000 \text{km}; \end{array} \begin{array}{l} \text{Reactor Experiments} \\ \text{Accelerator experiments} \end{array}$$



Accelerator Measurement

$$P_A(\nu_\mu \rightarrow \nu_e) \sim \sin^2 2\theta_{13} \left(\sin^2 \theta_{23} - \frac{0.05}{\sin 2\theta_{23}} \sin \delta \right)$$

$$1 - P_R(\bar{\nu}_e \rightarrow \bar{\nu}_e) = P_R(\bar{\nu}_e \rightarrow \bar{\nu}_\mu) + P_R(\bar{\nu}_e \rightarrow \bar{\nu}_\tau) \mp \sin^2 2\theta_{13}$$

Reactor Measurement

Complementarity of Reactor-accelerator θ_{13} measurement

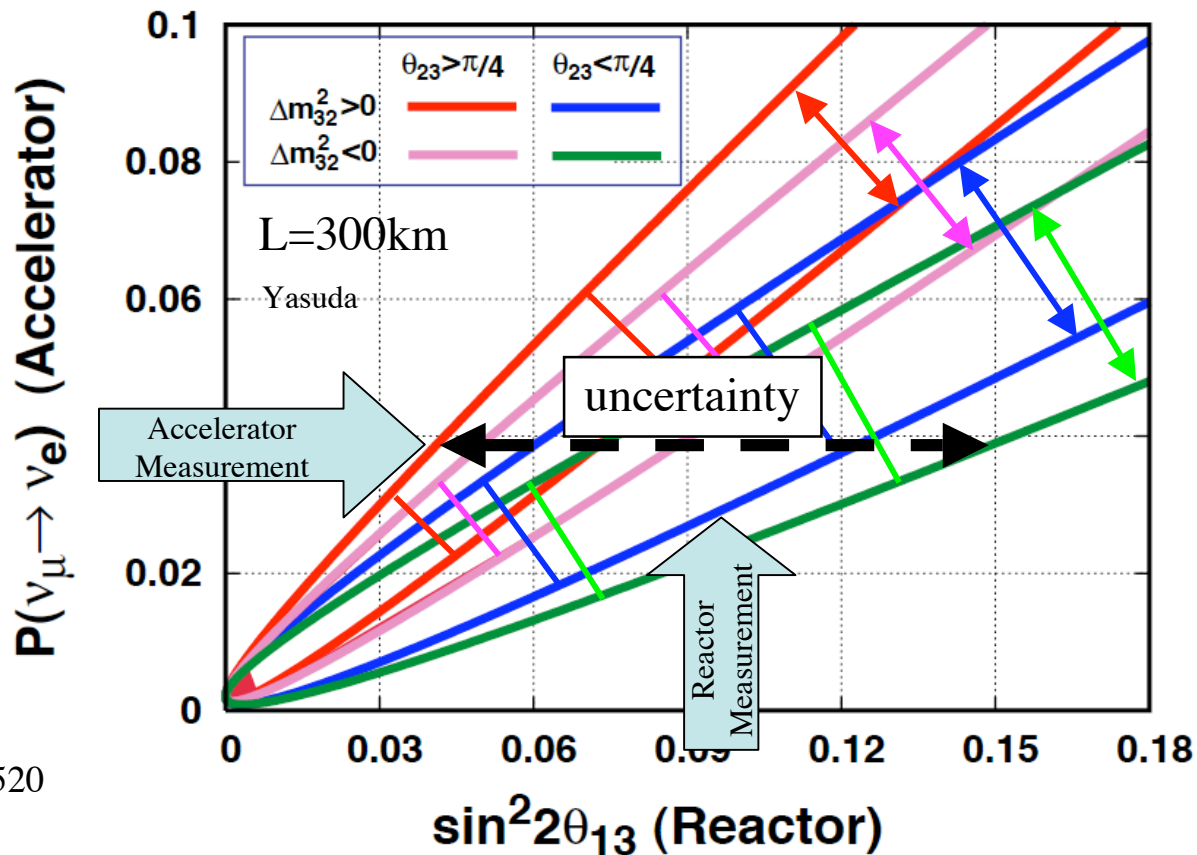
θ_{23} degeneracy ↘

$$P_{AC}(\nu_{\mu} \rightarrow \nu_e) = \frac{0.50 \pm 0.11}{(1 \mp 0.00017L[km])^2} \sin^2 2\theta_{13} \pm 0.045 \sin 2\theta_{13} \sin \delta$$

Matter effect ↗

$\sin^2 2\theta_{23} = 0.95$

δ dependence



How to improve the Chooz limit

Systematic Error:

	Chooz
Reactor cross section	1.9 %
Number of protons	0.8 %
Detector efficiency	1.5 %
Reactor power	0.7 %
Energy per fission	0.6 %

Cancel → Near/Far detector comparison
& Cut-insensitive measurement

Chooz case:

$$\left(\frac{\delta N_\nu}{N_\nu} \right)_{\text{CHOOZ}} = \pm 2.8\%(\text{stat.}) \pm 2.7\%(\text{sys.}) \Rightarrow \sin^2 2\theta_{13} < 0.15$$

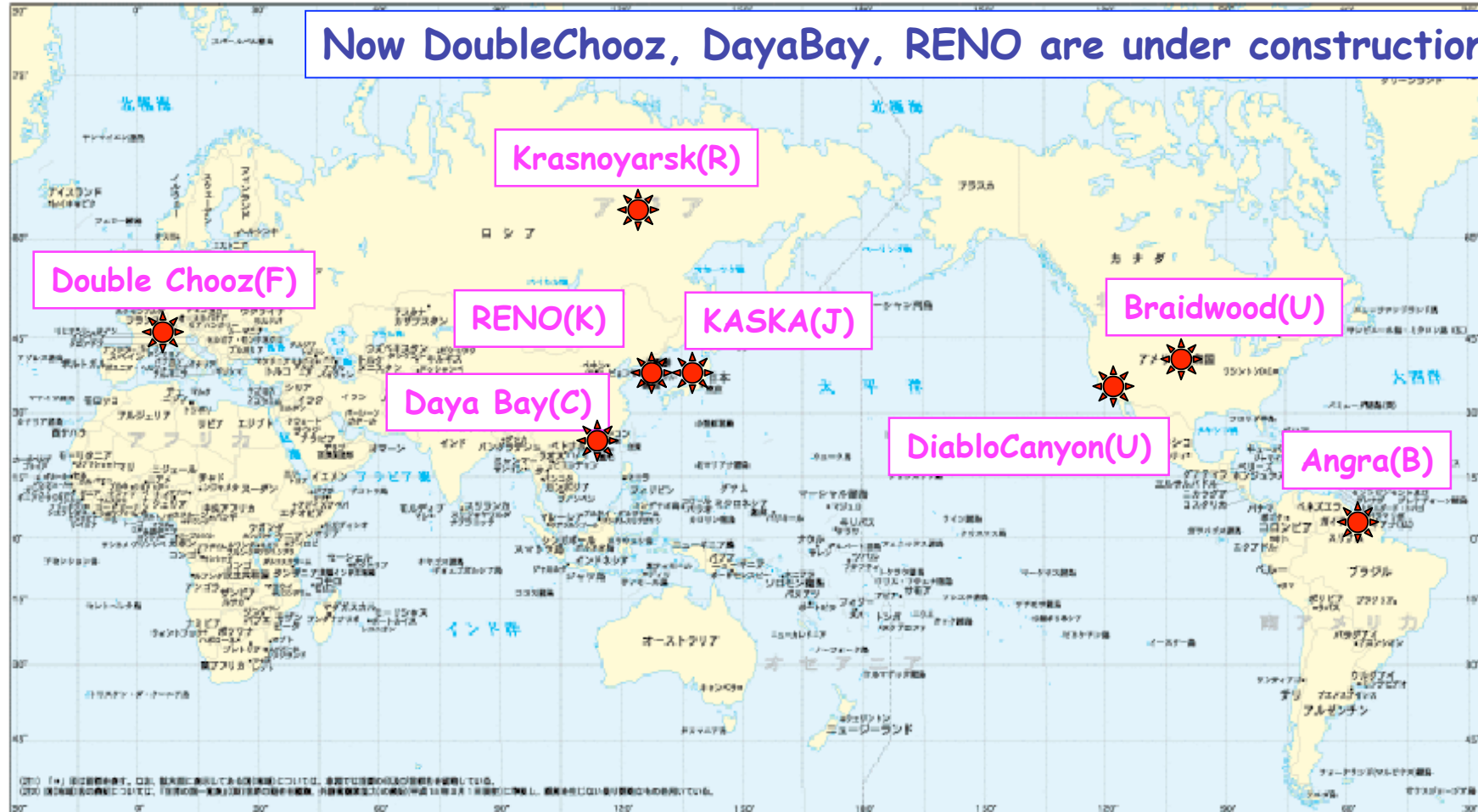
Statistic error:

Chooz ran only a few months

=> Longer run, Larger Detector & Higher Reactor Power

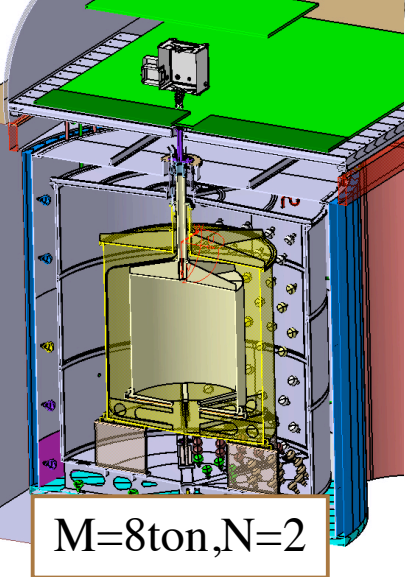
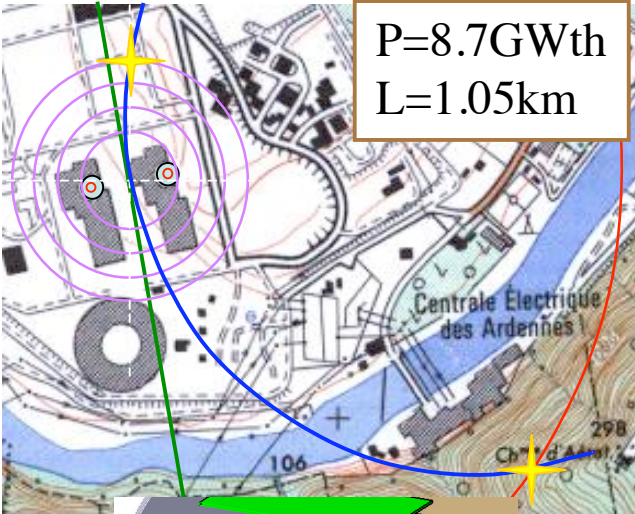
Reactor- θ_{13} Site Historical Map

Now DoubleChooz, DayaBay, RENO are under construction

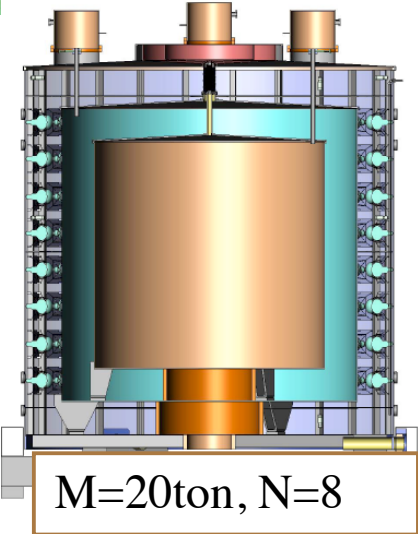
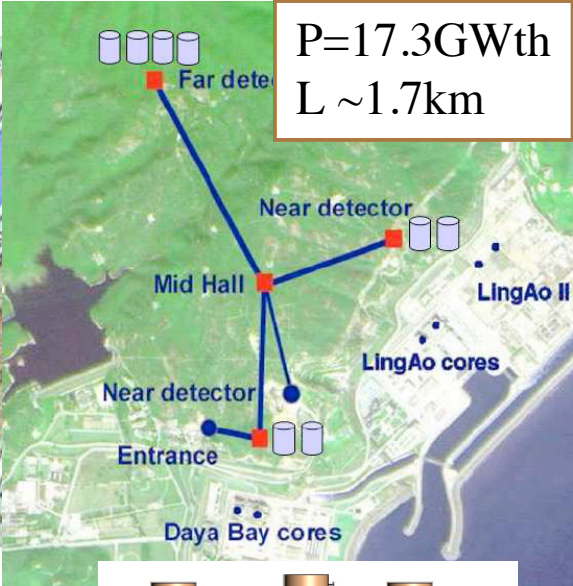


Reactors and Detectors

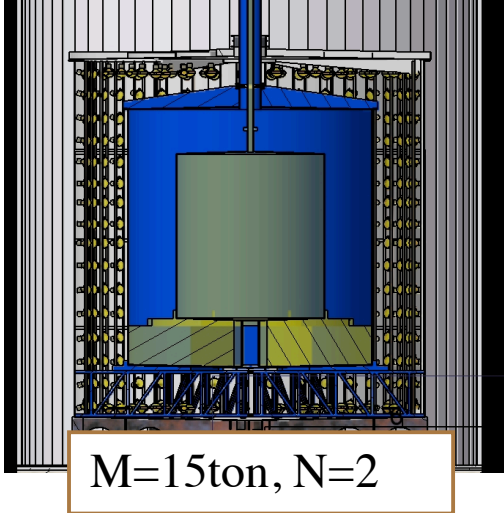
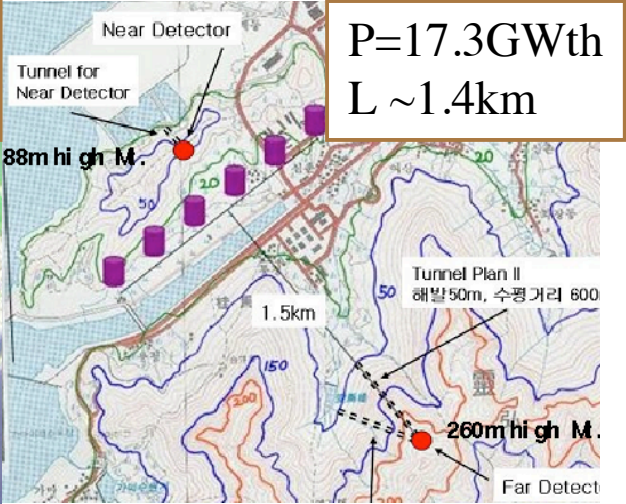
Double Chooz



Daya Bay



RENO

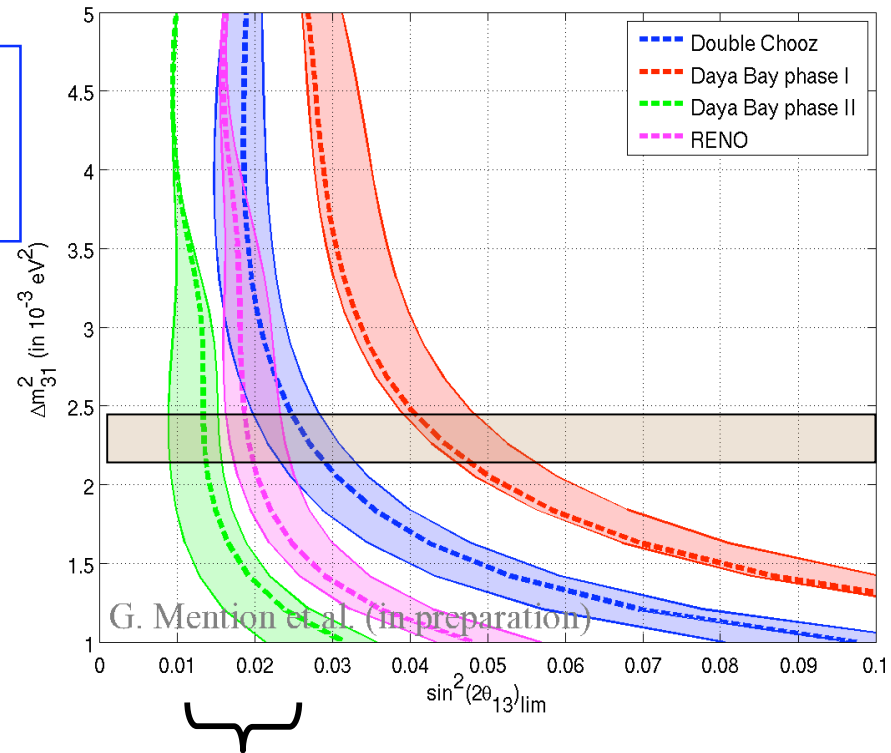


Sensitivity & Start years

DCHOOZ:
 2009 Starts with Far Detector
 2010 run with Far+Near Detector

DayaBay:
 2010 Starts with $P=11.6\text{GW}$
 2011 $P \rightarrow 17.4\text{GW}$

RENO:
 2010 Starts



Sensitivities=0.01~0.03

→ We will know the results within a few years!

Status of the experiments

DoubleChooz
RENO } → { Please hear
the **T. Lachenmaier** & **S.-B. Kim's** talks
this afternoon

DayaBay → K.B.-Luk kindly provided me the following slides

Civil Construction

- Total tunnel length is about 3100 m
- SAB ready by July 2008
- Daya Bay Hall: Nov 2008
- Ling Ao Hall: Jun 2009
- Far Hall: Sept 2009



Complete entrance portal of access tunnel in March 08



Inside the construction tunnel



Constructing surface assembly building

Getting Ready To Build The Detectors



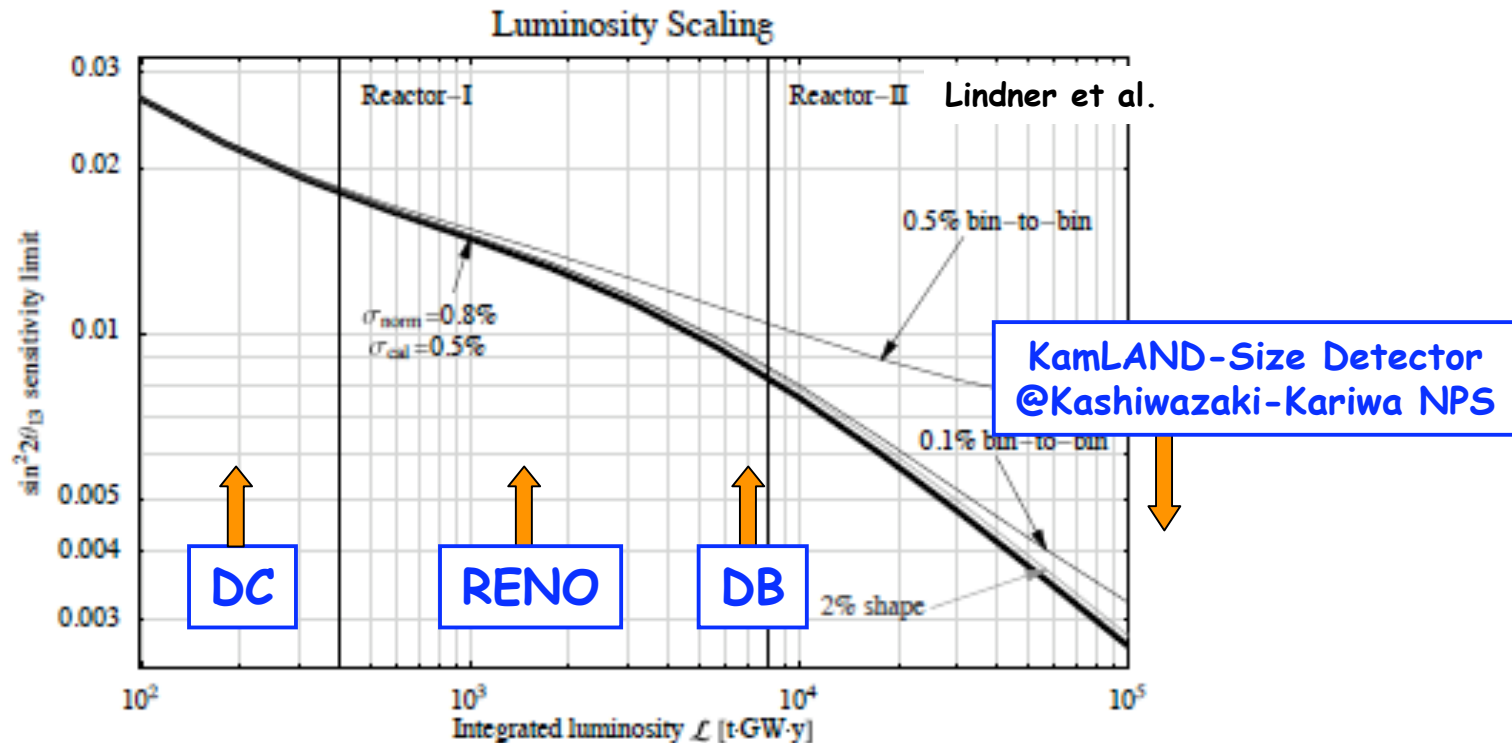
4-m prototype in the U.S.



3-m prototype in Taiwan

Very Precise θ_{13}

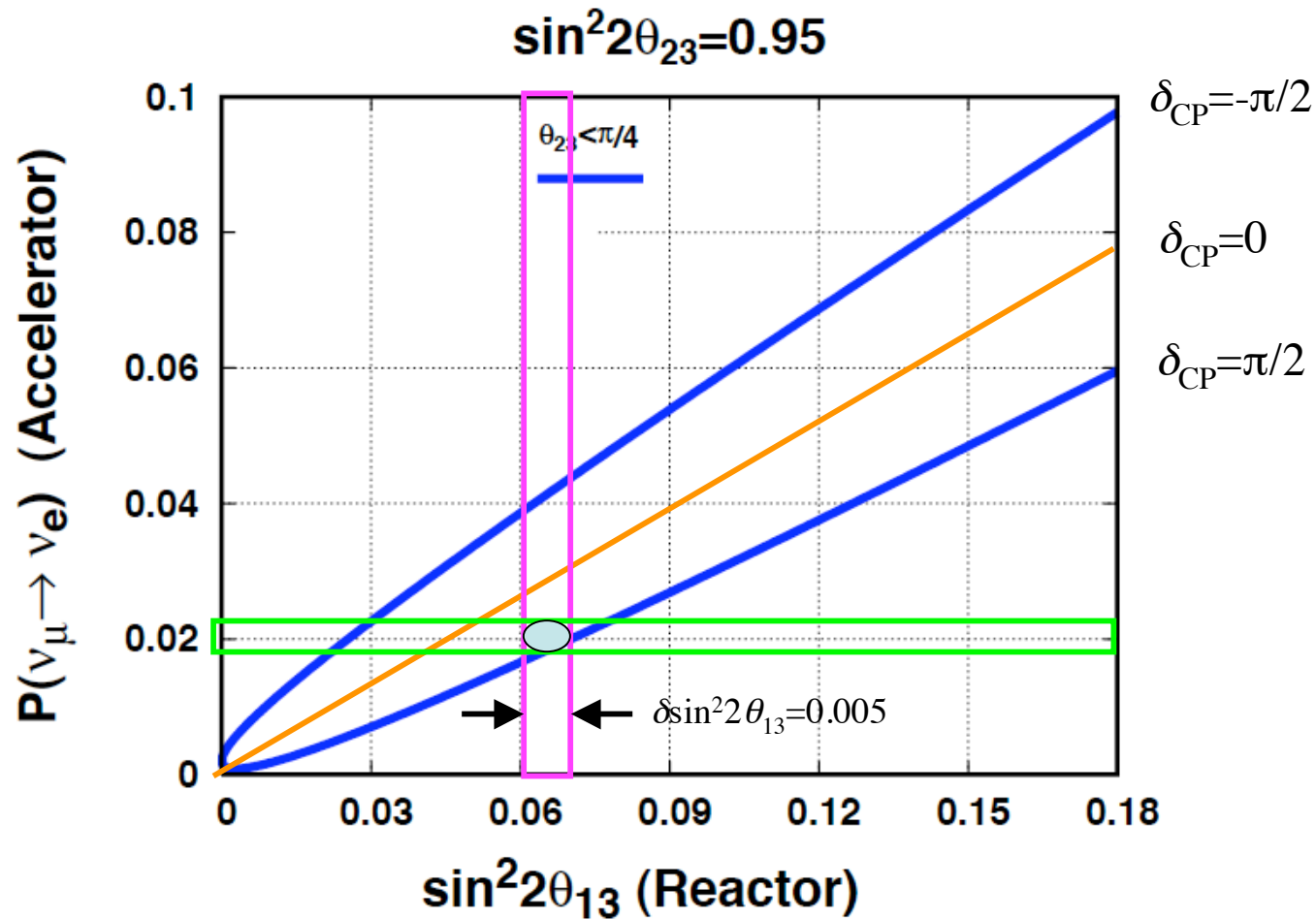
$\delta \sin^2 2\theta_{13} < 0.01$ will be possible by large statistics & shape analysis.



The current Reactor θ_{13} experiments will be in stable condition within a few years

=> It is time to think of the next generation experiments seriously

δ_{CP} detection in future



If $\delta \sin^2 2\theta_{13} < 0.01$ is achieved, there is chance to detect finite δ_{CP} combined with Accelerator data.

Reactor-Accelerator cooperation

J-Parc group has already started to discuss about post T2K CP experiments.













<http://j-parc.jp/NP08/>

Thursday 06 March 2008

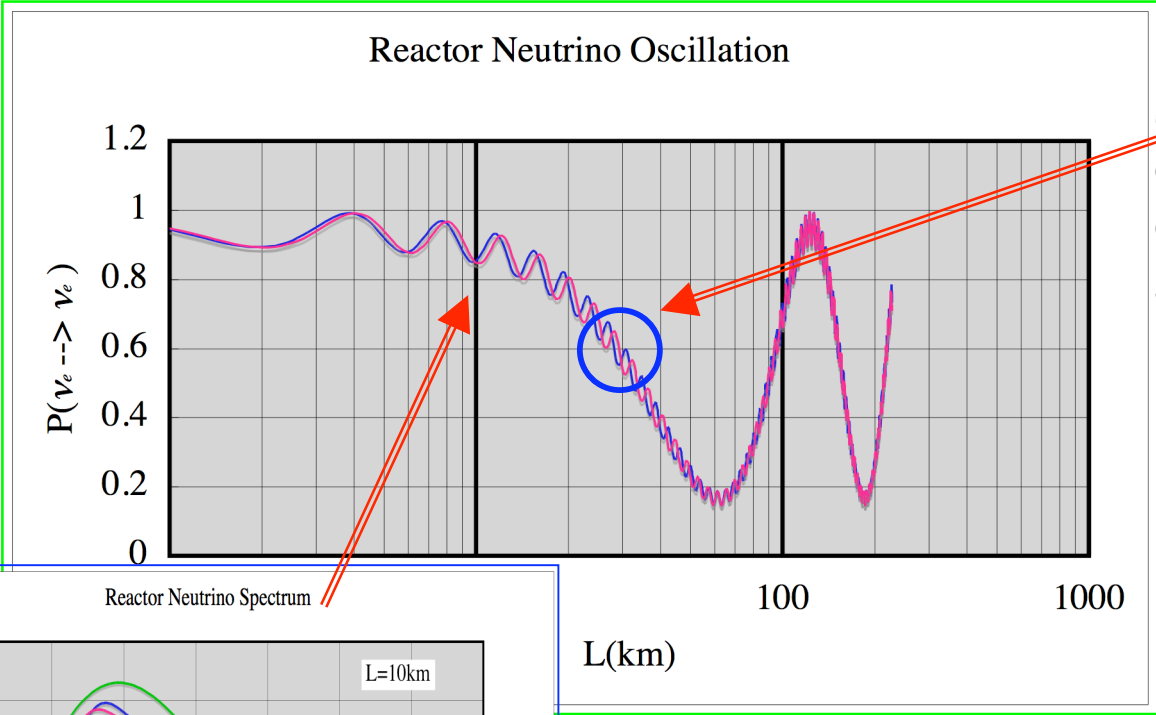
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09:30->13:00 Parallel sessions 1

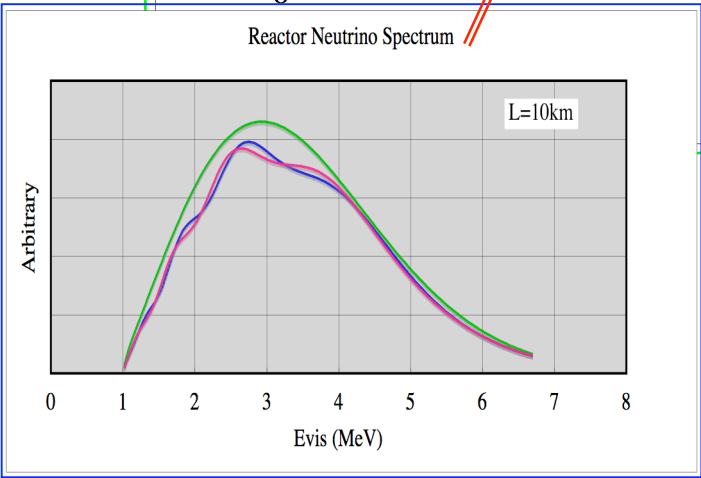
- 09:30 Future Beam Option for Long Baseline Neutrino Experiment (30') ( [Slides](#) ) Yves Declais (*Lyon Univ. IPNL/IN2P3/CNRS*)
- 10:00 Introduction to the Work on the 2nd Phase Experiment with J-PARC Neutrino Facility (15') ( [Slides](#) ) Takuya Hasegawa (*KEK*)
- 10:15 CP Violation Physics at a J-PARC Beam -Liquid Ar TPC Case- (45') ( [Slides](#) ) Takasumi Maruyama (*Tsukuba Univ.*)
- 11:00 CP Violation Physics at a J-PARC Beam -Water Cherenkov Detector Case1- (25') ( [Slides](#) ) Kenji Kaneyuki (*ICRR Univ. of Tokyo*)
- 11:25 CP Violation Physics at a J-PARC Beam -Water Cherenkov Detector Case2- (20') ( [Slides](#) ) Fanny Dufour (*Boston Univ.*)

Future reactor and accelerator programs should be interweaved from the first, in order to make efficient strategy to attack δ_{CP} .

Δm^2_{13} & Mass Hierarchy



Large phase shift due to mass hierarchy. Can't this be used to solve it??



Several Oscillations take place in the Energy Spectrum
 \Rightarrow Precise Δm^2_{12} measurement will be possible @L~10km

Summary

- Typical energy of reactor ν is $\sim 4\text{MeV}$.
Both Δm^2_{12} & Δm^2_{13} oscillations are accessible.
- $\tan^2\theta_{12}$, Δm^2_{12} are being measured by KamLAND with $L\sim 180\text{km}$.
- For precise $\tan^2\theta_{12}$ measurement, $L=50\text{km}$ with KamLAND size detector $\Rightarrow 3.6\%$
- 1st phase $\sin^2 2\theta_{13}$ experiments (DoubleChooz, Dayabay, RENO) will start within 1~2 years. Their targetting sensitivities are $\delta\sin^2 2\theta_{13}=0.01\sim 0.03$
- 2nd phase $\sin^2 2\theta_{13}$ experiment ($\delta\sin^2 2\theta_{13}<0.01$) should be useful to detect δ_{CP} together with accelerator measurement of precise $P(\nu_{\mu}\rightarrow\nu_e)$.
- There is a possibility to measure Δm^2_{13} @ $L\sim 10\text{km}$.
- Reactor ν experiments /have had/are going to have/will have/ rich programs to do.