

*Christchurch, New-Zealand*  
*May 26, 2008*

*The speaker does not take any responsibility  
for the title of his talk*

*Still he will do his best to make sense out of it*

# Where are we? Where are we going?



**A. Yu. Smirnov**

*International Centre for Theoretical Physics, Trieste, Italy  
Institute for Nuclear Research, RAS, Moscow, Russia*

# Beloved question of $\nu$ physicists

“Where are we?”

HEP SPIRES: 52 papers found

13 papers on neutrinos

J. N. Bahcall - 6, on solar neutrinos  
B. Kayser - 2

*which can be explained by elusive character of the subject of research*

Not only neutrino physicists are lost:

Where we are in

particle physics - 2  
high energy physics - 2  
heavy ion collisions - 2  
non-baryonic DM - 2

...

Can you guess?

string theory - 1

confidence and determination

The SSC: where are we now?

no comment

# in addition

Encouraging:

S. Weinberg,  
S. L. Glashow,  
L. Lederman

among those who asked

Variations on the theme:

How it started and

Where we are

and where do we stand  
and where should we go  
and we're headed and why  
coming from

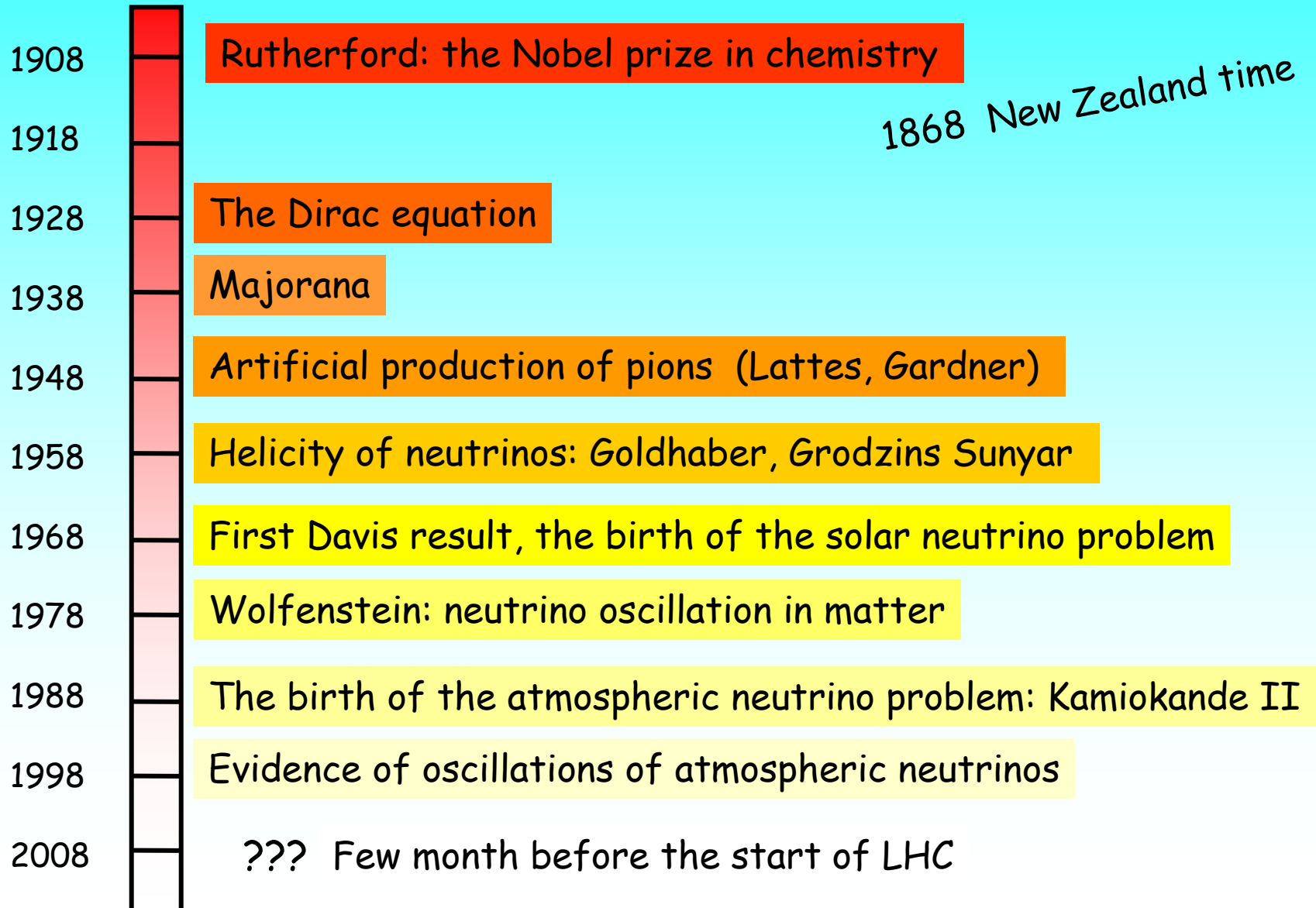
Even more profound:

What we are?

*J. Ellis (with reference to P Gauguin)*



# Where are we in time



# in space

SuperKamiokande  
Tokai

INO

CERN  
Gran Sasso

Fermilab

in 3D:

IceCube

for orientation..

just in case if one wants  
to construct underground  
laboratory or accelerator  
in Christchurch

... in extra dimensions:

somewhere on  
the EW brane



# ... in the field of neutrino physics



- Conquest territory:  
the standard  
neutrino scenario

- Where are we in  
understanding  
neutrino masses  
and mixing

- Beyond the standard  
neutrino scenario

- Future which we  
know

- Future which we  
don't know  
can imagine

# Standard neutrino scenario

1. Neutrino interactions are described by the Standard (electroweak) model.

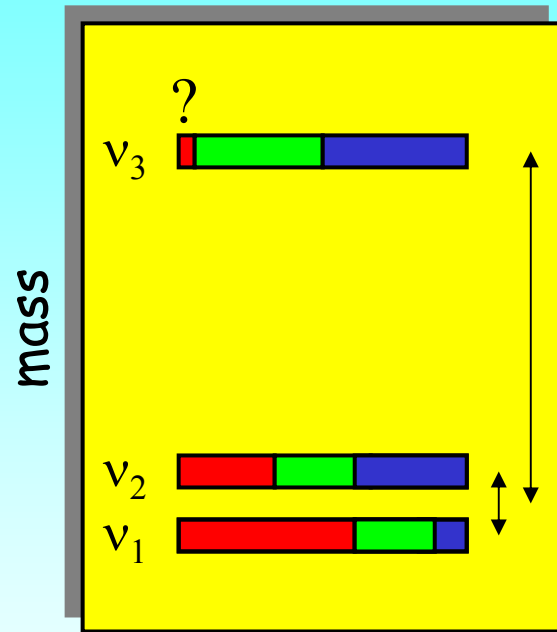
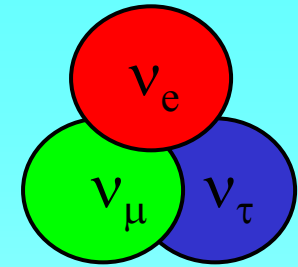
2. There are only 3 types of light neutrinos: three flavors and three mass states.

3. Neutrinos are massive. Neutrino masses are in the sub-eV range - much smaller than masses of charge leptons and quarks.

4. Neutrinos mix. There are two large mixings and one small or zero mixing. Pattern of lepton mixing strongly differs from that of quarks.

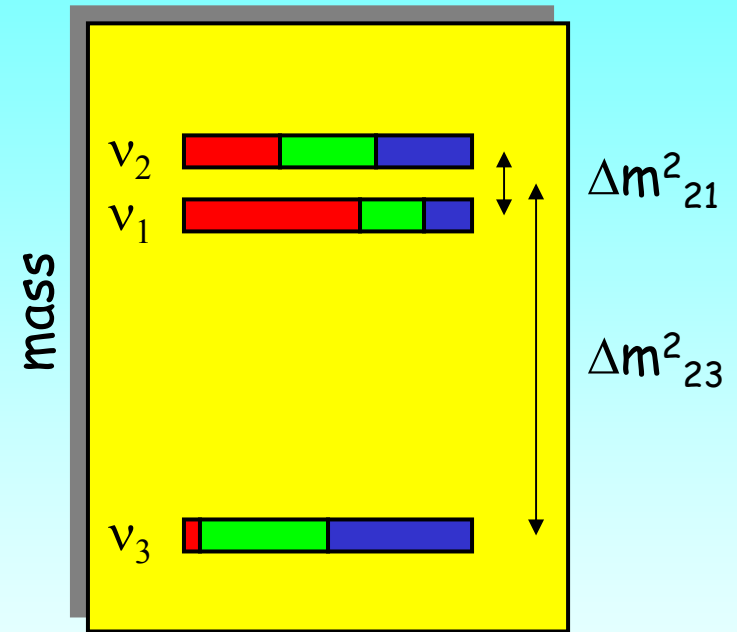
5. Masses and mixing have pure vacuum origin; they are generated at the EW and probably higher mass scales.

# Spectrum



Normal mass hierarchy

?

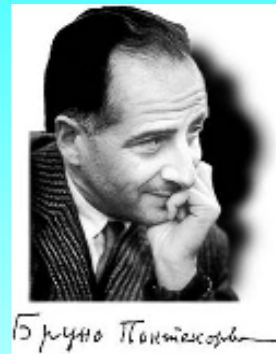


Inverted mass hierarchy

$$\nu_f = U_{PMNS} \nu_{mass}$$

$$U_{PMNS} = U_{23} I_\delta U_{13} I_{-\delta} U_{12}$$

# Standard scenario



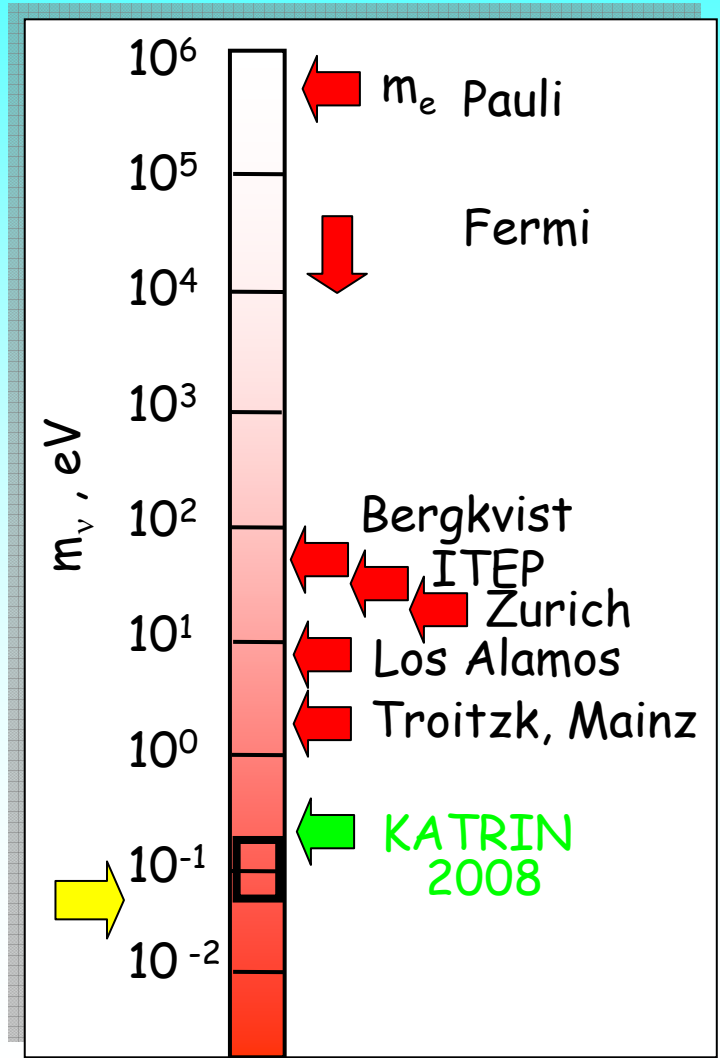
Result of work of several generations of neutrino physicists

collective efforts of experimentalists and theoreticians

Our conquest territory. Basis and Starting point for further advance. Reference scenario

Summary of results of the first phase of studies of neutrino mass and mixing

“neutrino time”



# Interactions

Gauge interactions are well known (SM) and well checked

Yukawa couplings with Higgs boson(s) - related to existence of the RH neutrinos are unknown; they are relevant for leptogenesis

Interaction with complex systems:  
nucleons and nuclei,  $\nu N$ -,  $\nu A$ : open questions

problem of strong interactions.

Single-pion production, forward production

In some cases - neutrinos are unique:  
provide axial vector current  $\rightarrow$  axial vector anomaly  $\rightarrow$   
interactions of  $Z, \gamma, \omega$  (for MiniBooNE) *J Harley, C.T. Hill, R. Hill*

Nuclear physics for  $\beta\beta$ -decay

Rare neutrino processes relevant for astrophysics,  
 $\nu\nu$ - pair production, etc.



# Propagation

Theory of  $\nu$ -oscillations in vacuum: still some discussions

``Eternal Questions''

Paradoxes of  $\nu$ -oscillations

- momentum vs. energy
- stationary source approximation
- relevance of the wave packets
- coherence

matters for

Oscillations of ``Mossbauer neutrinos''

In medium: oscillations at extreme conditions - high densities, temperatures, magnetic fields, etc..

Neutrinos in neutrino gases: effects of the  $\nu\nu$ -scattering, collective, non-linear effects

# Phenomenology

of the scenario

**Solar  
neutrinos**

**Long baseline  
experiments**

**Atmospheric  
neutrinos**

**Supernova  
neutrinos**

**Cosmic neutrinos**

**Relic neutrinos**

**Reactor  
neutrinos**

To large extend phenomenology of standard scenario has been elaborated; in some cases - in great details

Still some areas exist (cosmic, supernova neutrinos) where active research continues now

Few spots are not covered yet

# Comments

## Solar neutrinos

Comprehensive description of neutrino conversion;  
very precise analytic results

From experimental side - detection of

- the earth matter effect  
(day-night asymmetry, zenith angle dependence of signal);
- upturn of spectrum;
- N,O, pep- neutrinos
- pp-neutrinos (??)

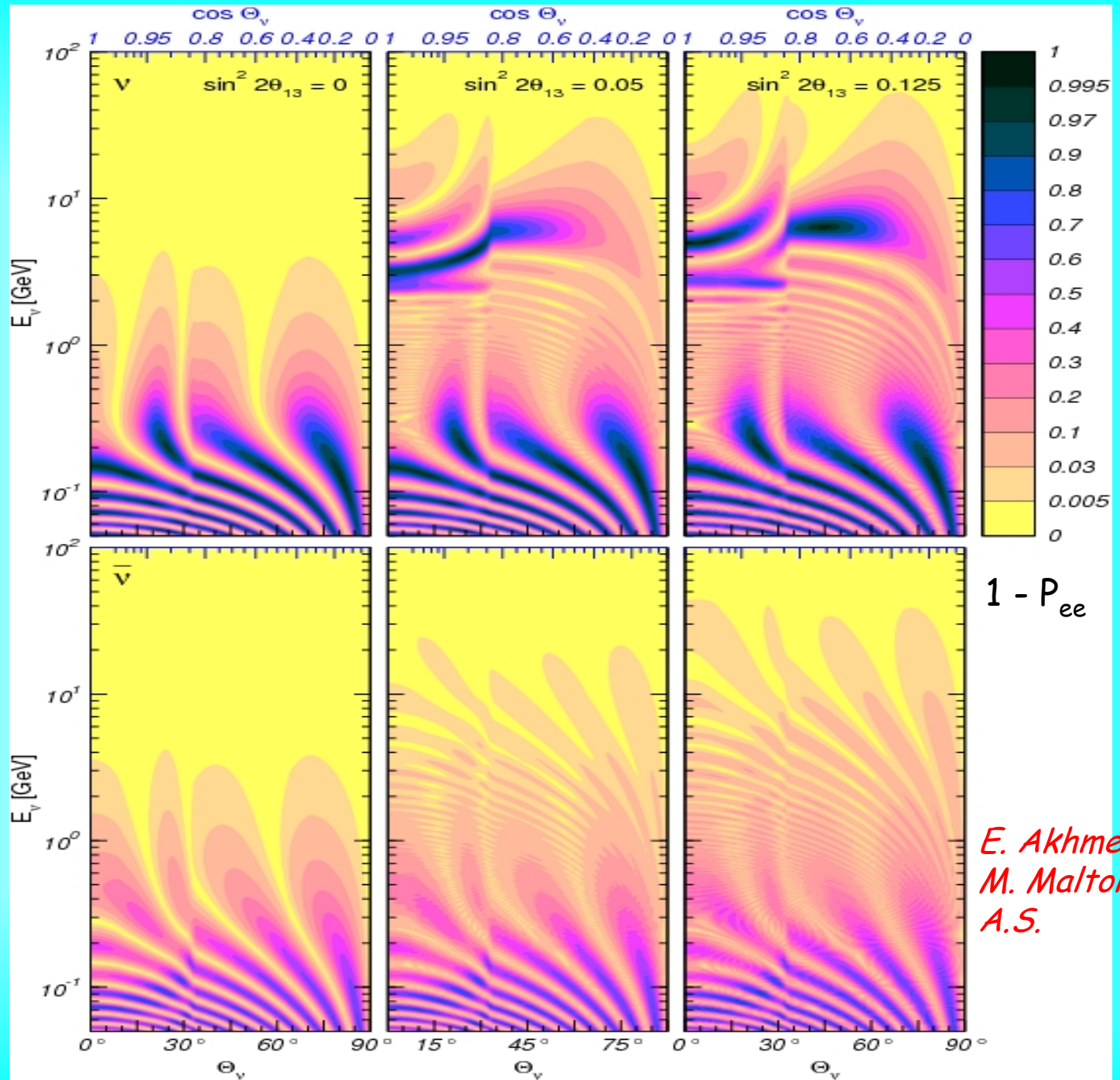
## Atmospheric neutrinos

Comprehensive description of physical processes in terms of oscillograms of the Earth

Structures of oscillograms:

- collinearity condition
  - generalized phase condition
- CP- domains given by grids on magic lines  
and lines of the interference phase condition

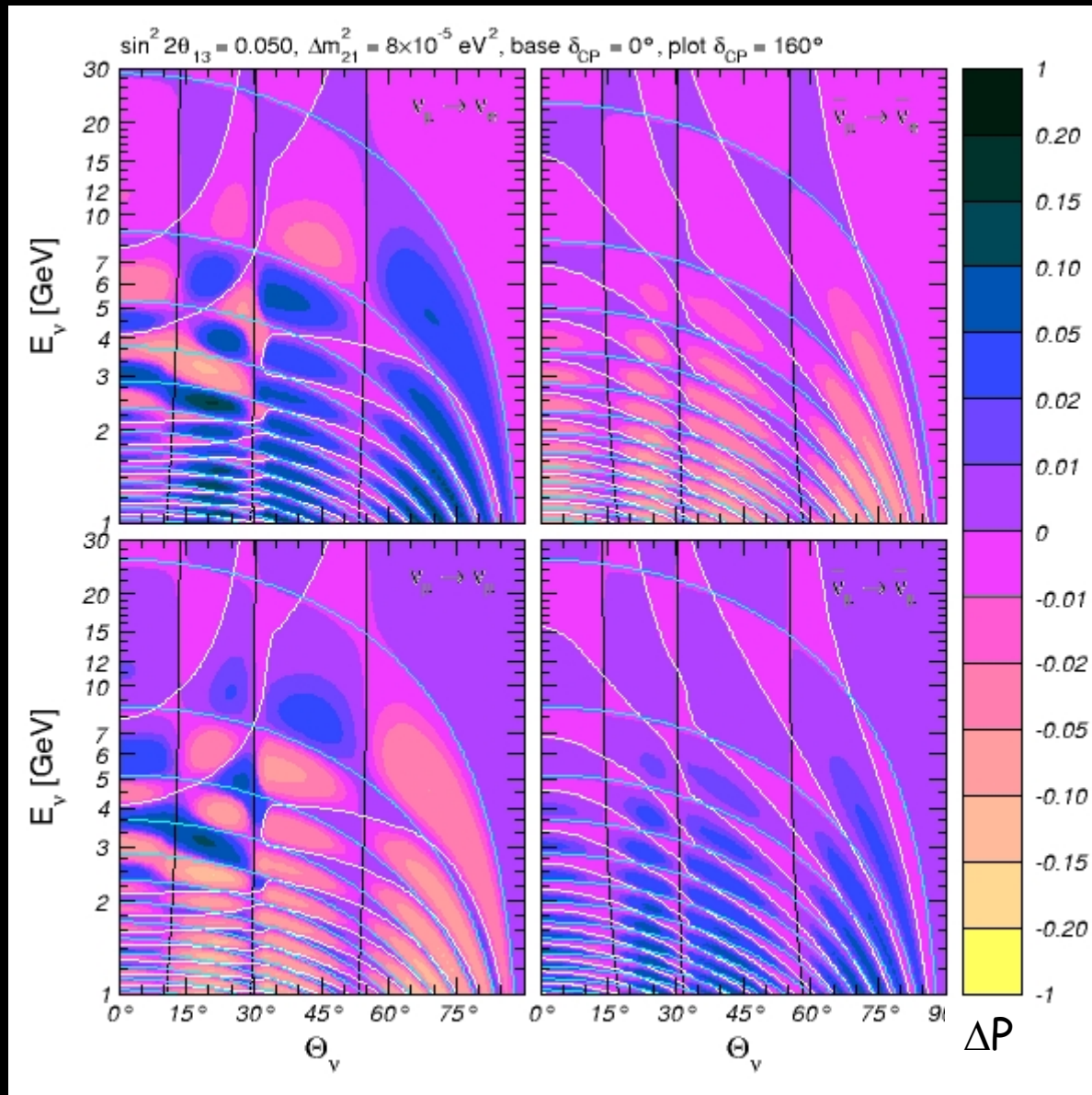
# Oscillograms of the Earth



# CP-violation domains

$$P(\delta) - P(0)$$

formed by  
grids of  
magic lines  
and lines of  
interference  
phase





# Supernova neutrinos

Further studies of collective effects:

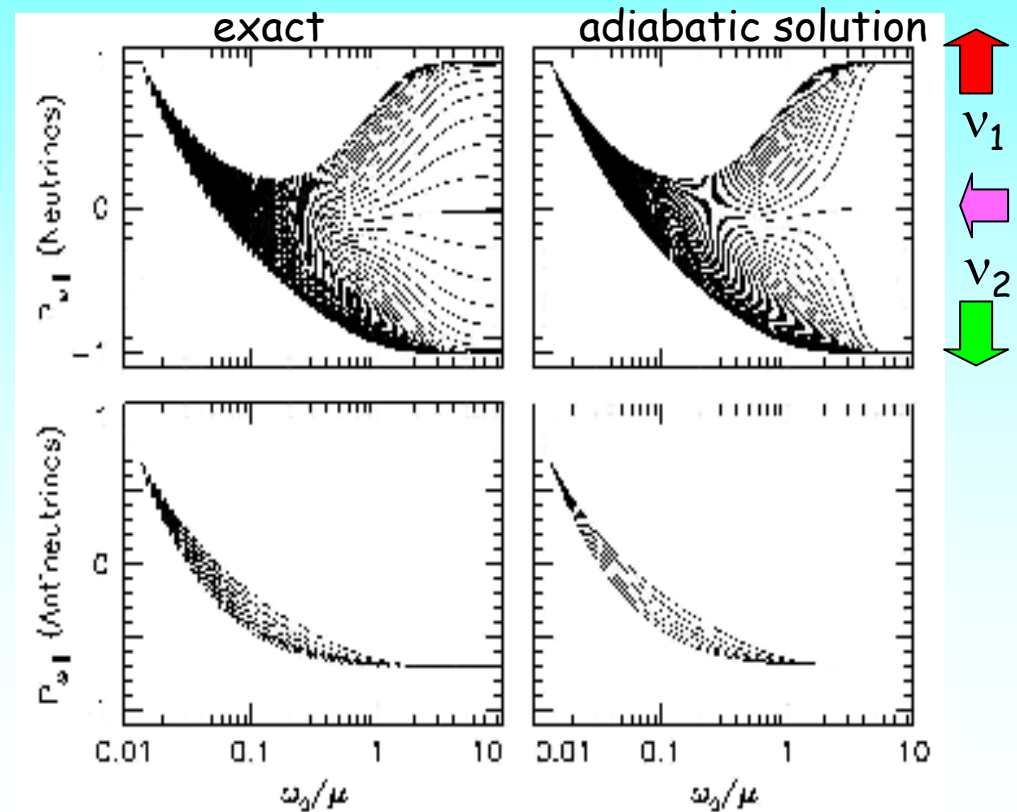
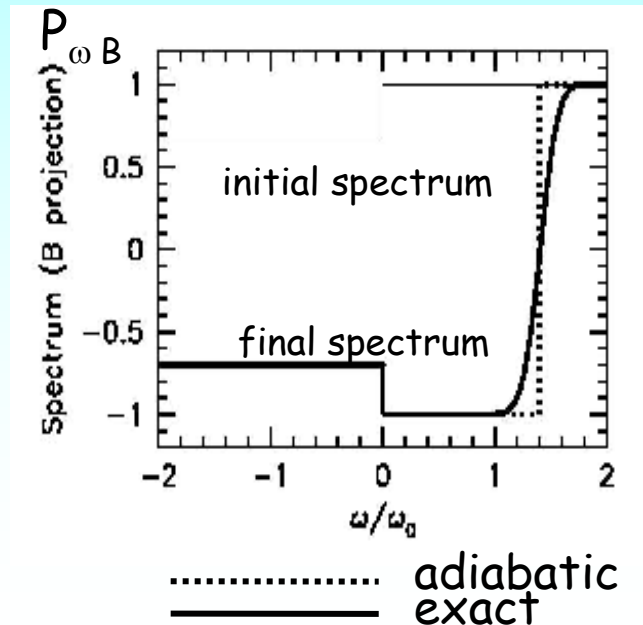
can change whole picture of conversion

$$\mu = \sqrt{2} G_F n_\nu$$

$$\omega = \Delta m^2 / 2E$$

Spectral splits (swapping)

*H. Duan, G. Fuller, J. Carlson,  
Y. Z. Qian  
G. G. Raffelt, A.S.*



*G. Raffelt, A.S.*

# Cosmic neutrinos

New level of studies

## Sources:

AGN

GRBs

core-collapse  
supernovae

SN remnants

microquasars

blazars

Related to  
developments  
of  $\gamma$ -astronomy

$E \sim 1 \text{ GeV} - 10^4 \text{ TeV}$

Detailed computations of  
the neutrino yield (output)  
at different conditions

## Propagation:

Vacuum oscillations

Conversion in matter  
of source

studies of various  
non-standard effects

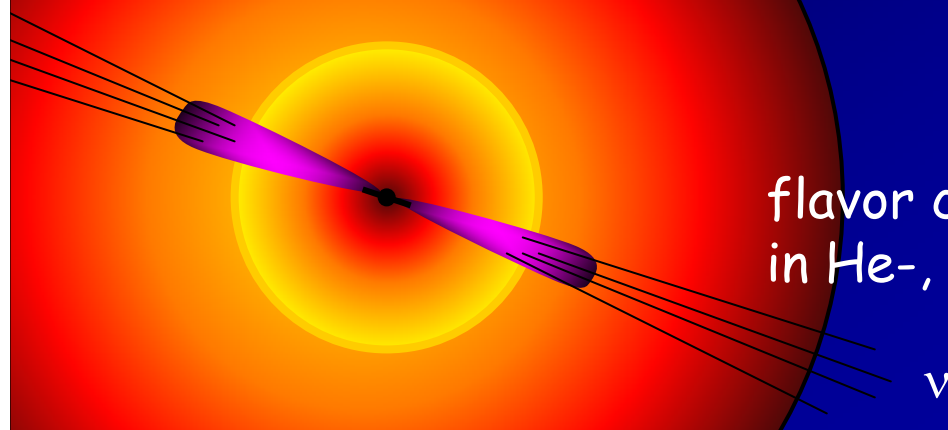
for maximal 2-3 mixing  
and 2 : 1 : 0 original ratio  
flavor equilibration:

1 : 1 : 1

deviations from  
1 : 1 : 1

- $\nu$  production mechanism
- $\theta_{23} = \pi/4$

# Neutrinos from astrophysical sources



flavor conversion  
in He-, H- envelopes

$\nu$

thick source

acceleration of protons in relativistic  
jets by the inner shocks

pp -, pv - collisions  $\rightarrow$  neutrinos

flavor conversion in outer layers  
 $\rightarrow$  breaking of 1:1:1 flavor equilibration

Measurements of deviation  
of 2-3 mixing from maximal

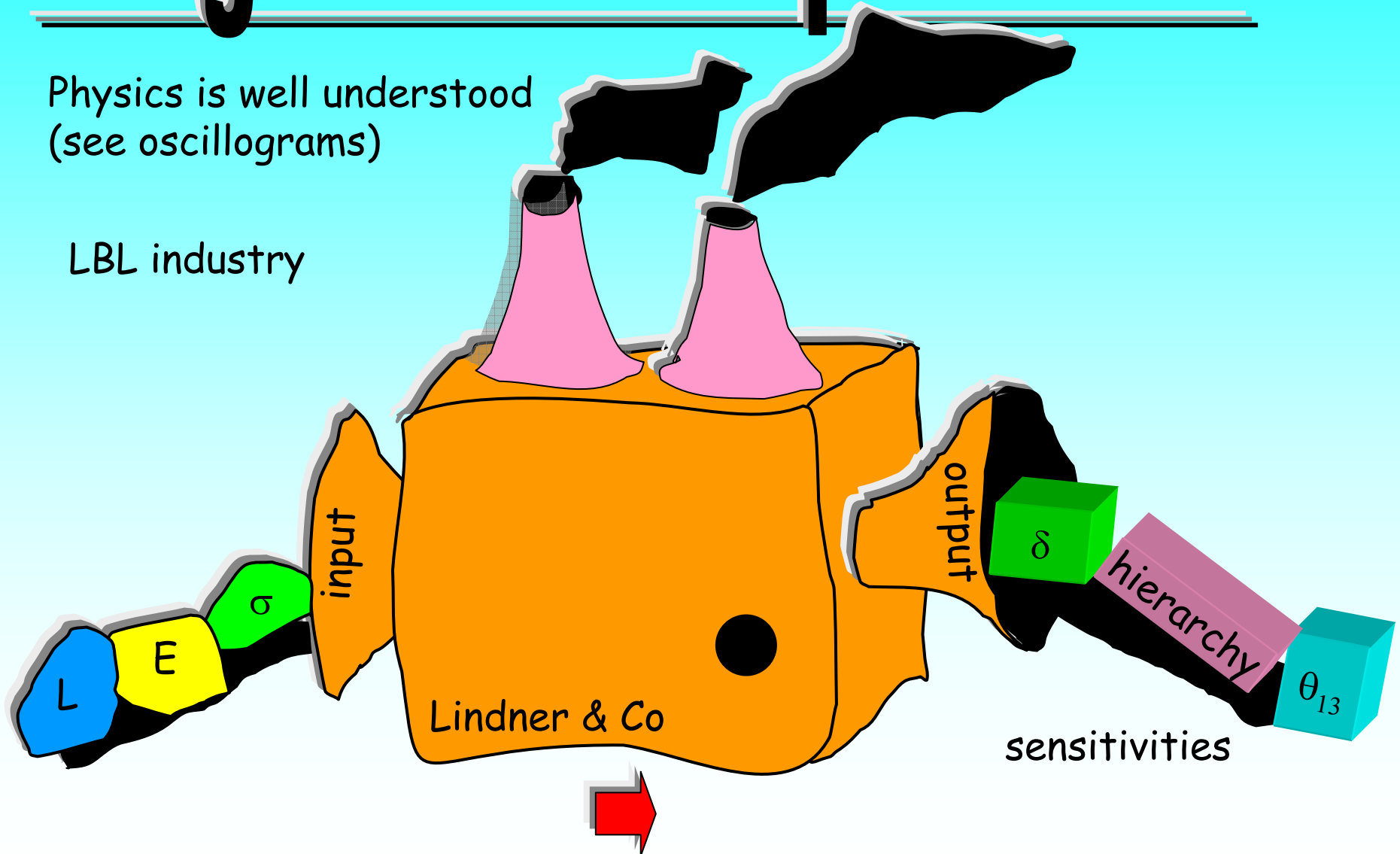
Sensitivity to 1-3 mixing,  
type of mass hierarchy,  
CP - phase



# Long baseline experiments

Physics is well understood  
(see oscillograms)

LBL industry



# Masses and Mixing

Standard scenario

Right handed components of neutrinos exist

Smallness of mass is due to some mechanism, that involves the EW scale and probably some higher scale(s) of nature

In general:

$$m_\nu = m_{\text{hard}} + m_{\text{soft}}(E, n)$$



medium-dependent soft component

Bounds on  $m_{\text{soft}}$

Neutrinos are Majorana particles?

In the context of the see-saw mechanism:

$M_R(\text{heaviest}) \sim M_{\text{GUT}}$  is an interesting possibility

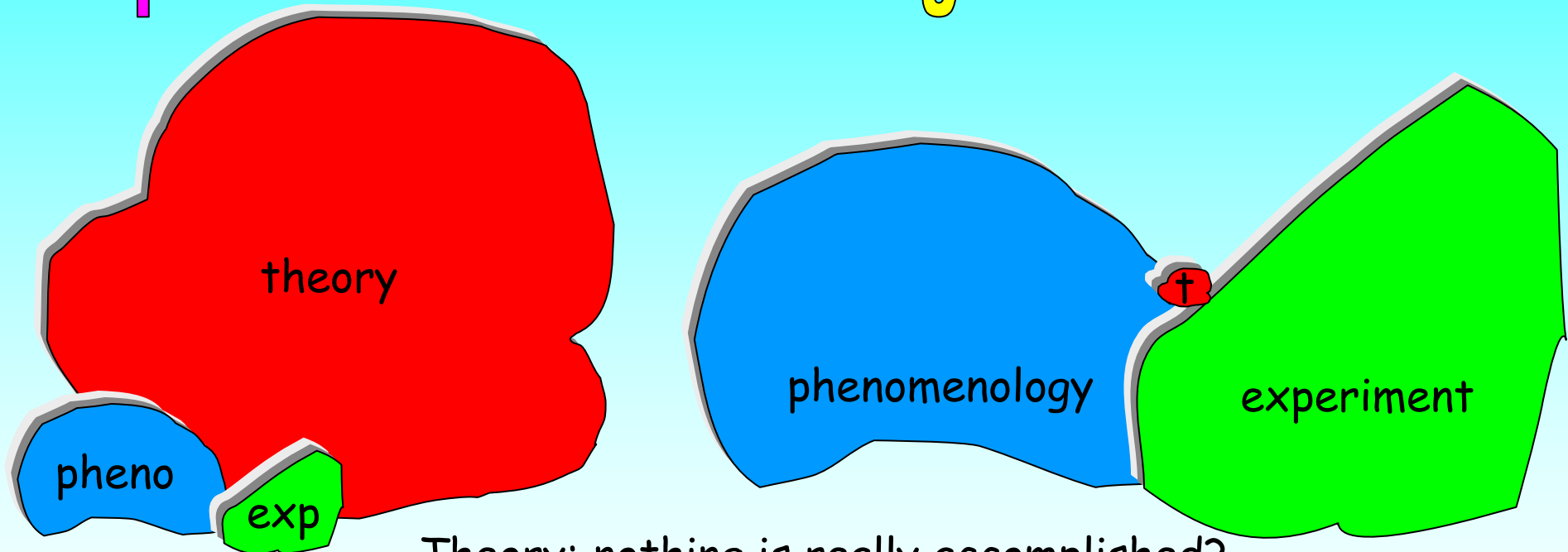
Difference of the quark and lepton mass spectra and mixing patterns is related to the smallness of neutrino mass

**Where are we  
in understanding  
neutrino mass & mixing?**

# Theory of neutrino mass

Papers 2006 - 2008

Program: nu-2008



Theory: nothing is really accomplished?  
No progress? Why? What should be done, measured?

Recall we measure neutrino parameters  
to uncover the underlying physics  
to make on this basis new predictions,  
to find applications

*Whole the excitement was that neutrino mass and mixing are manifestations of physics beyond the standard model!*

*Dramatically, after many years and many trials the underlying physics hasn't been identified. Certain problems have been realized.*

*Nevertheless, we should further pursue this search, concentrating on how the progress can be achieved*

# Bottom-up

Tri-bimaximal mixing

With different implications:

Flavor symmetry

Extension to quarks?

- subject of RGE
- no relation to masses?
- additional ambiguities (CP-phases)

Quark-Lepton complementarity

"bi-maximal - CKM"

$QLC_{\nu}$   
 $V_{CKM} + V_{bm}$

$QLC_l$   
 $V_{bm} + V_{CKM}$

Quark-lepton symmetry, GUT

structure which produces bi-maximal mixing - symmetry?

GST-approach mass-mixing

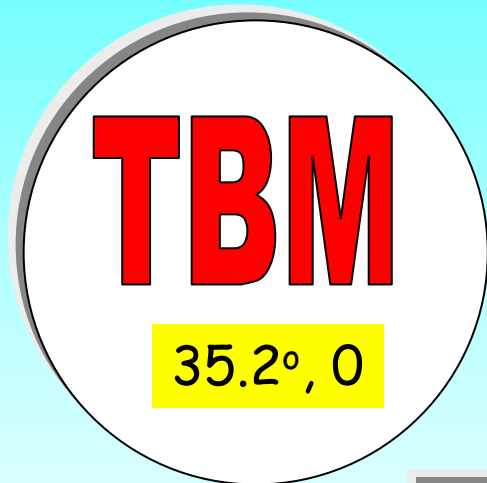
Gatto-Sartory-Tonin

The same principle as in quark sector

Large mixing is related to weak mass hierarchy of neutrinos

# Disentangling possibilities

$\theta_{12}, \theta_{13}$

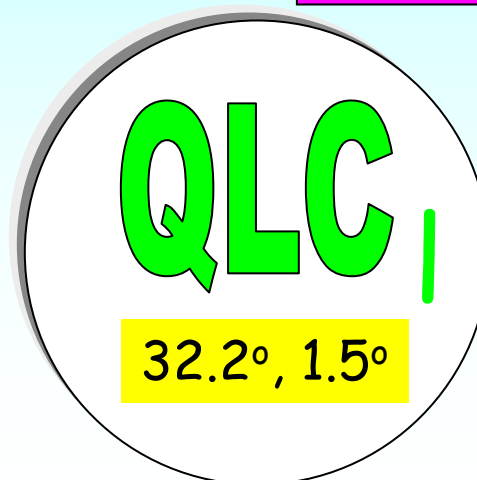


$\theta_{13}$



$\theta_{12}, \theta_{13}$

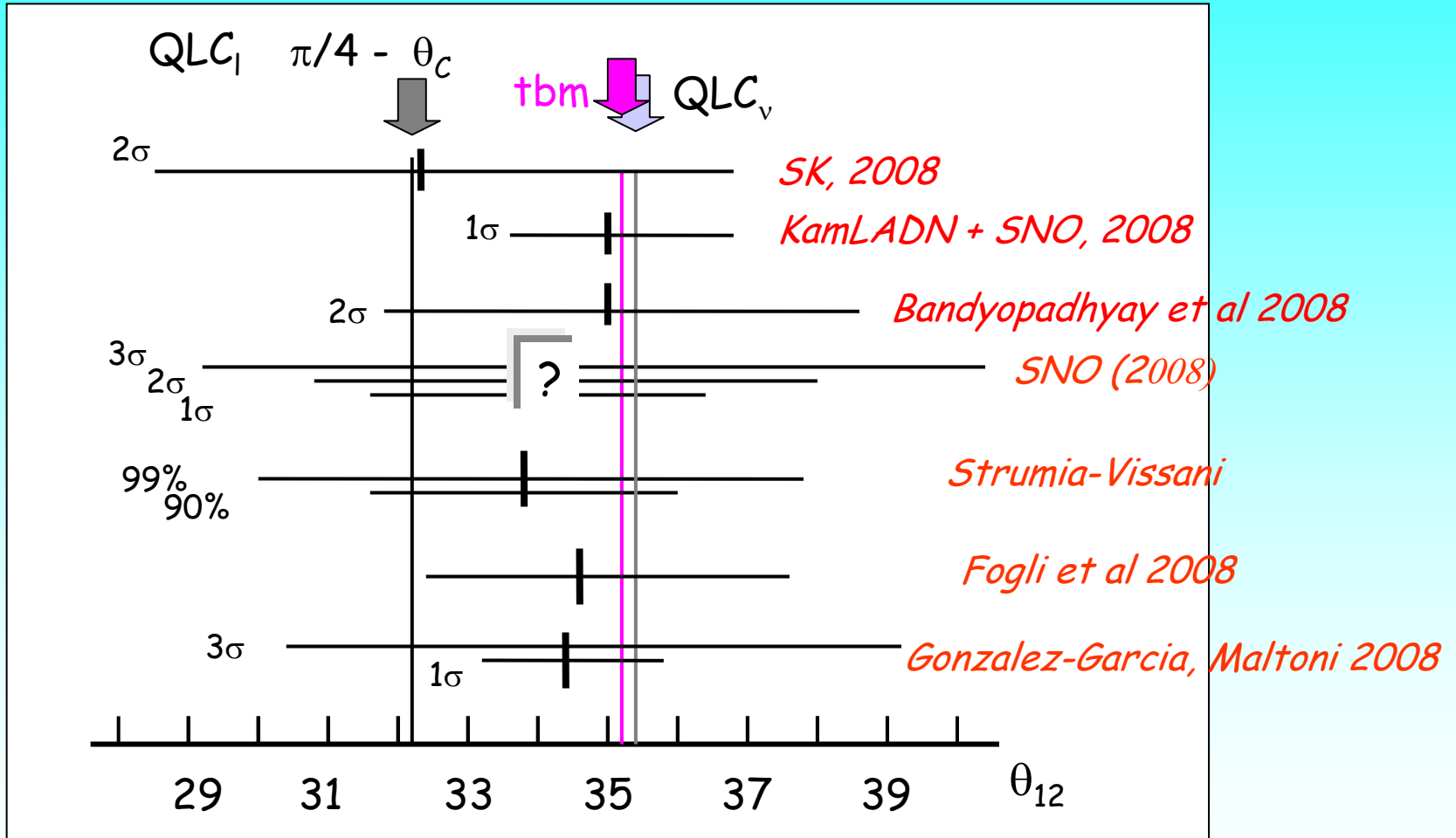
$\theta_{12}$



TBM: 23-mixing  
should be zero  
or very small

# 1-2 mixing

Precision & benchmarks



$$\theta_{12} + \theta_c \sim \pi/4$$

$$U_{QLC1} = U_C U_{bm}$$

$$U_{t_{bm}} = U_{tm} U_{13}^m$$

give almost same 12 mixing



# Koide relation

*Y. Koide, Lett. Nuov. Cim.*  
*34 (1982), 201*

$$\frac{m_e + m_\mu + m_\tau}{(\sqrt{m_e} + \sqrt{m_\mu} + \sqrt{m_\tau})^2} = 2/3$$

was obtained in attempt to explain

$$\tan\theta_C = \sqrt{3} \frac{\sqrt{m_\mu} - \sqrt{m_e}}{2\sqrt{m_\tau} - \sqrt{m_\mu} - \sqrt{m_e}}$$

with accuracy  $10^{-5}$

all three families are involved:  
no perturbation approach!

Both relations can be reproduced if

$$m_i = m_0 (z_i + z_0)^2$$
$$\sum_i z_i = 0, \quad z_0 = \sqrt{\sum_i z_i^2 / 3}$$

*C A Brannen*

Neutrinos,  
hierarchical  
spectrum

Non-abelian flavor symmetry,  
VEV alignment  
Related to TBM?

# Flavor symmetries

## Real symmetry

Flavor features of various symmetry groups have been explored

## Discrete groups:

$A_4$  (subg.  $SO_3$ ) - many studies

$T_7$  (Frobenius) (subg.  $SU_3$ )

Looks promising

Deriving the group from observations

*SC Lam:*

TBM  $\rightarrow S_4$  minimal symmetry

## Effective symmetries:

- No symmetry (or some other symmetry) at the fundamental level.
- Required symmetry appears at the effective level after decoupling of heavy degrees of freedom

Partially realized in some models

**"See-saw symmetry"**

Along with this line:

``Symmetries from mass hierarchies''

*Ferretti, S. King, A. Romanino,*

# Real or accidental?

Tri-bimaximal mixing

Q-L-complementarity

Very small 1-3 mixing

Maximal 2-3 mixing

Koide relation

Real:  
immediate ``one step''

Accidental: interplay  
of different independent  
factors contributions

Discovery of degenerate neutrino mass spectrum would be convincing evidence of existence of symmetry

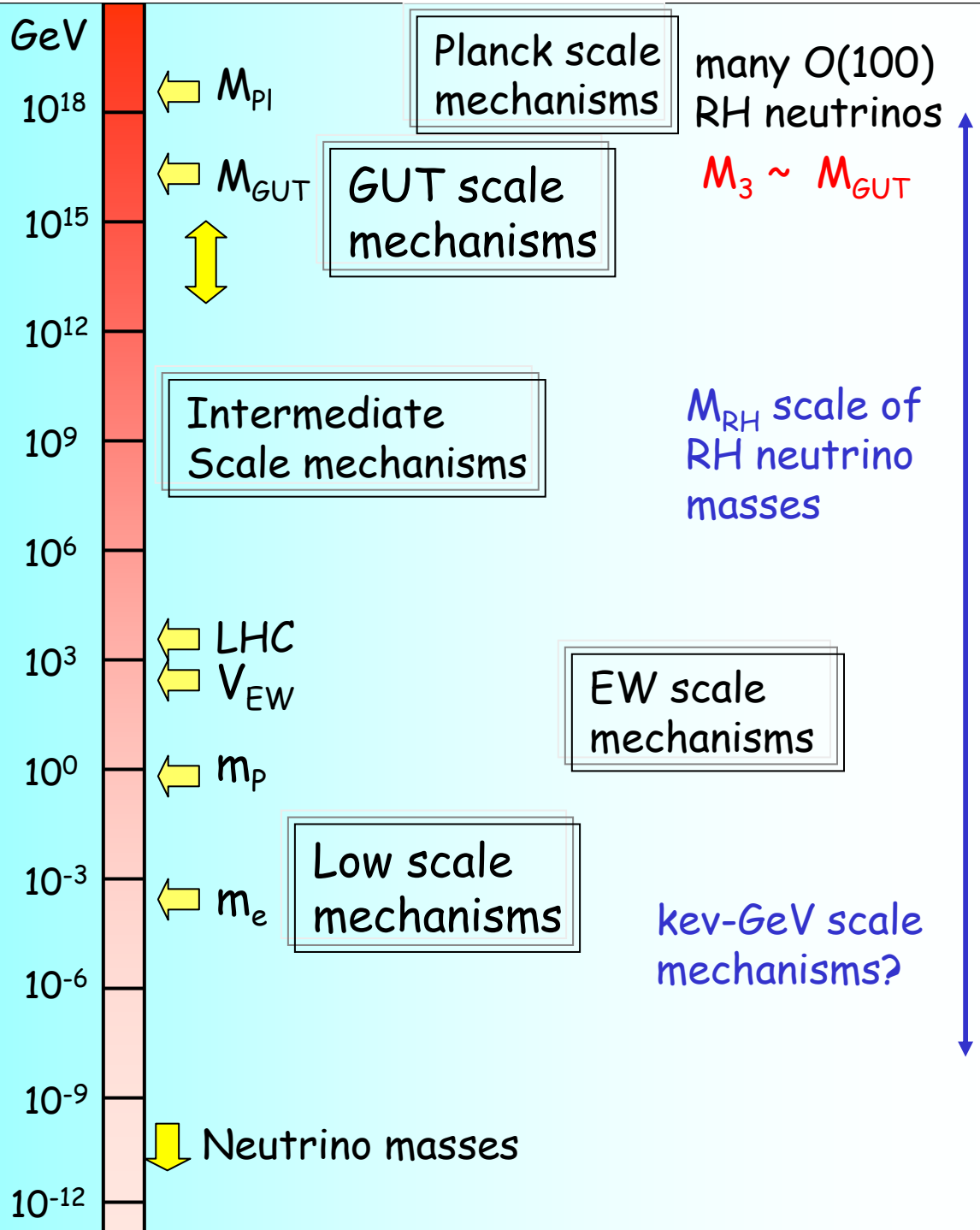
Heidelberg-Moscow

KATRIN

Degenerate spectrum and TBM are not related ?

# Energy scales of new physics

Physics behind neutrino masses is not identified



# Flavor & GUT

Generic problem:

GUT's: unification  
of quarks and leptons

difference of mass and  
mixing pattern

difference of flavor  
properties

Relate this difference to spontaneous breaking of GUT symmetry

**SO(10)**

$126_H$

Non-renormalizable operators

Singlet fermions

flavons

*C Hagedorn  
M. Schmidt A.S.*

SO(10) + singlets +  
discrete flavor symmetries



Geometrical hierarchy of  
the up quark masses

relates

Nearly maximal 2-3 mixing

# Beyond the standard scenario

Two aspects:

Tests of the standard  
scenario

Searches for new  
physics

# Classifying possibilities

## Neutrino anomalies

were driving force of the developments for more than 40 years

related to explanation of neutrino masses

New physics motivated by other fields

Various extensions of the standard model: SUSY, LR-symmetry, GUT extraD

Recall, 10 - 20 years ago  
``standard'' would be:

- zero mass,
- zero mixing...

## Unmotivated

# Anomalies: unresolved and new

seeds of new developments?

name:	feature:	possible interpretation
LSND	excess of $e^+$ -events	<i>see separate slide</i>
MiniBooNE	excess of events at low energies	
NuTeV	value of $\sin \theta_w$	mixing with very light sterile neutrino
Homestake	low signal, tension with other data	
Unnamed	time variations of solar neutrino signals?	- neutrino magnetic moment, - periodicity of energy release
SN 1987A	angular, time distributions LSD signal	Astrophysics?
Z <sup>0</sup> -width	$N_\nu < 3$	Hadron physics?
New GSI	modulation of exponential decay	Nothing to do with neutrinos?



# LSND after MiniBooNE

or MiniBooNE  
after LSND

True or fake... triggered a number of developments

Two sterile neutrinos with CP?

*M. Maltoni, T. Schwetz*

(Exotics)<sup>2</sup>

Light vector boson + 3 sterile nu

*A. Nelson, J. Walsh*

CPT violation + sterile neutrinos

*V. Barger, D. Marfatia, K. Whisnant*

Soft decoherence

*Y. Farzan, T. Schwetz, A.S.*

Something very exotic  
not connected to known  
physics processes



= Mechanism X  
or LSND-effect

Reconstruct from data  
its L and E dependence,  
checks of consistency

# New Physics

**NSI**

Non-Standard Interactions

Vector,  
 $U(1)'$

Tensor

Scalar

Light particles

Heavy

**New neutrino states**

Sterile neutrinos

**New dynamics**

Violation  
Of fundamental  
Symmetries: CPT  
Lorentz invariance  
Pauli principle

**Unparticle physics**

# Non-standard interactions

Motivation: new physics at the EW and terascale; various extensions of SM  
 $Z'$ , SUSY, KK, light particles

Rich phenomenology

- propagation
- detections

High energies where usual mixing is suppressed...

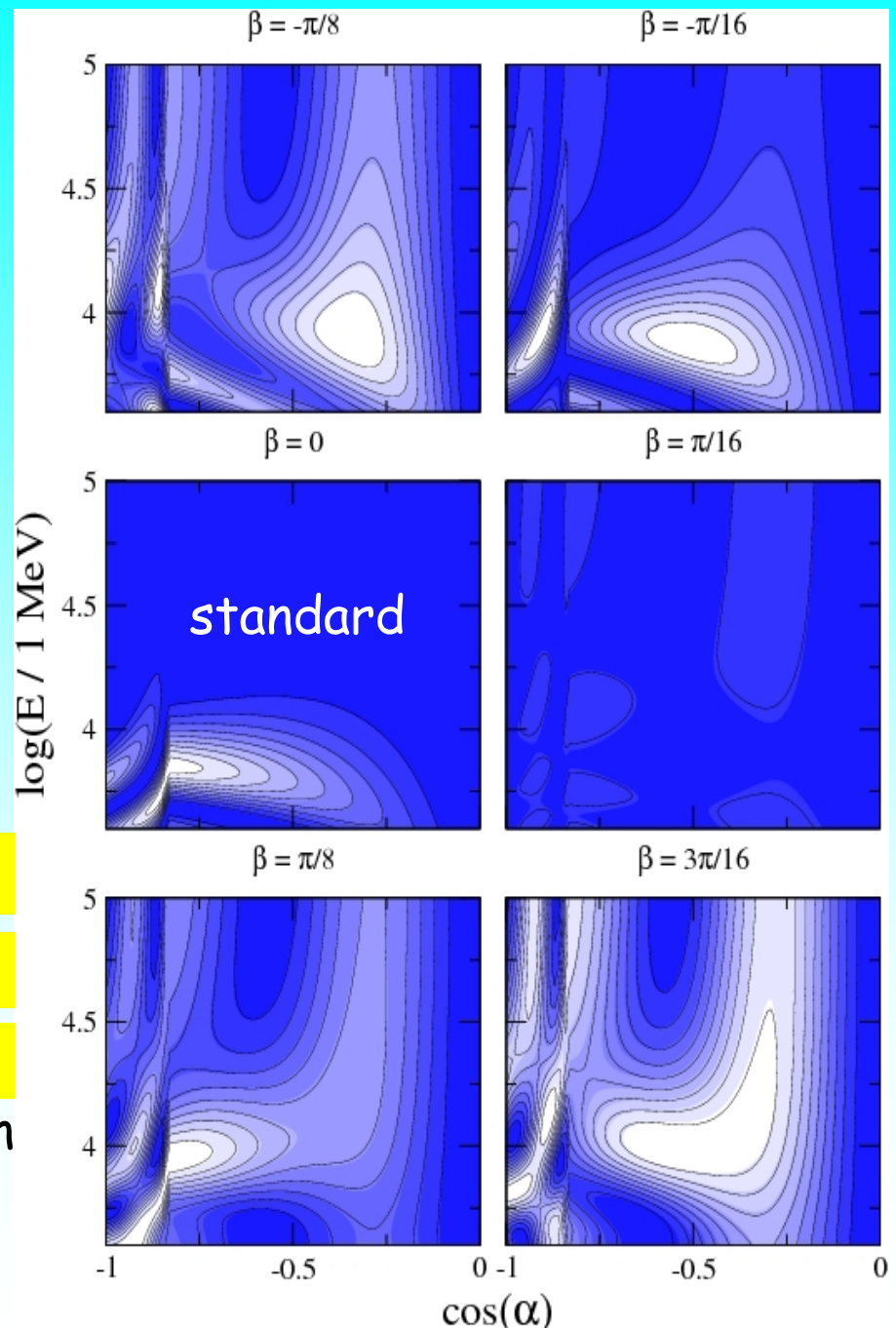
$$P_{ee}$$

$$\theta_{13} = 8^\circ$$

$$\varepsilon_{e\tau} \sim -0.5 \sin 2\beta$$

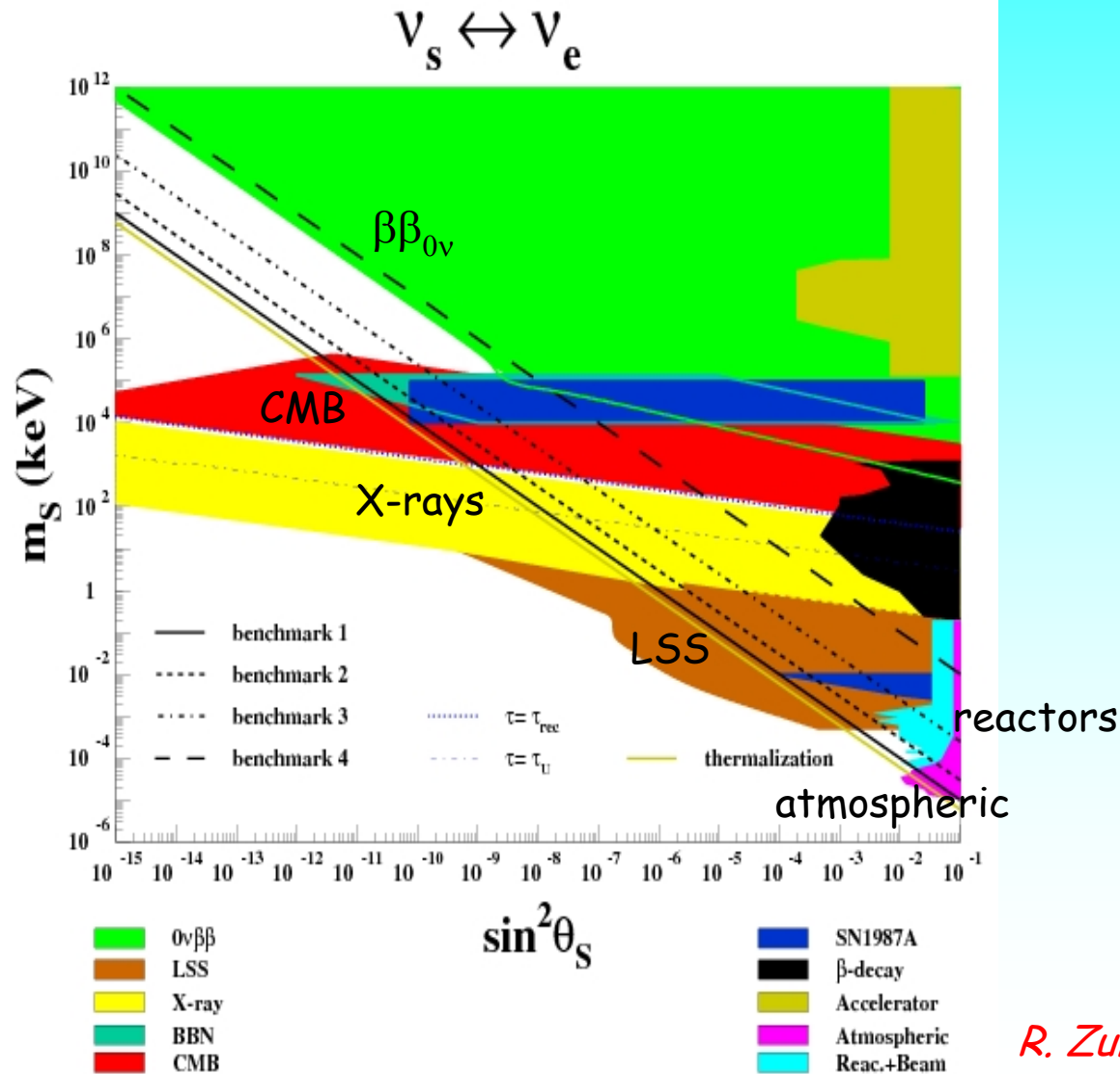
white: strong transition

Checks at LHC,  
 Rare processes with L-violation



*M. Blennow, T. Ohlsson, 0805.2301*

# Sterile Neutrinos



Contours of constant induced mass and bounds

thermalization line:  
above if S are in equilibrium

For benchmark parameters S were thermalized

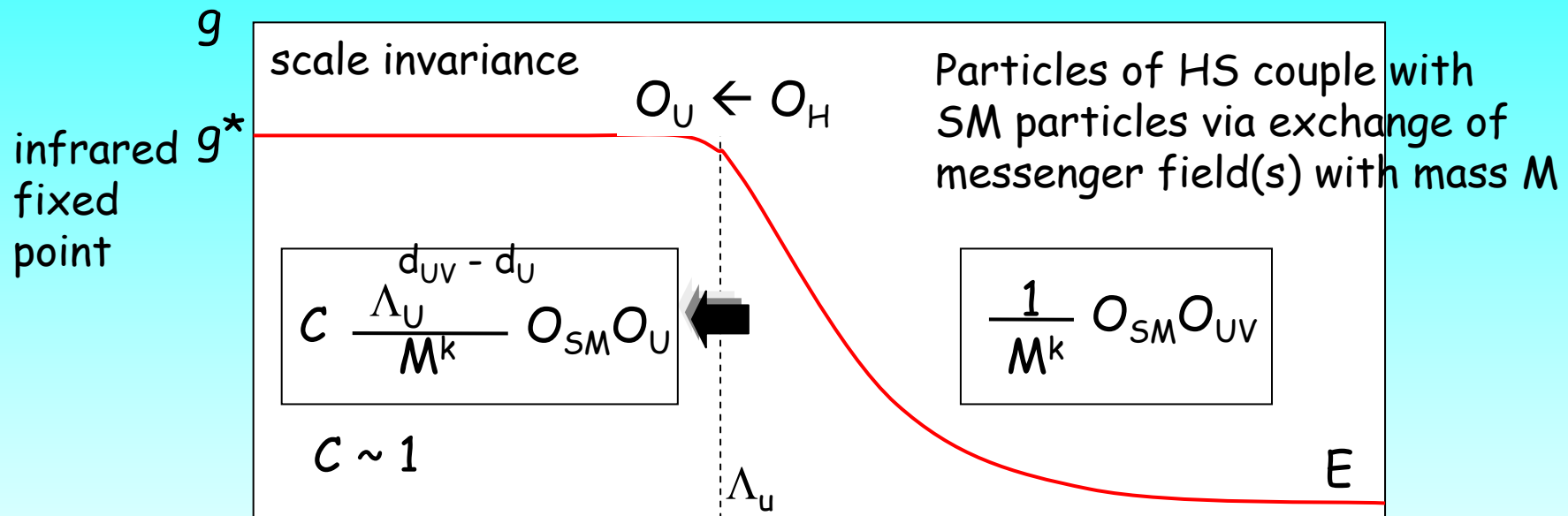
Bounds - in non-equilibrium region

*R. Zukanovic-Funchal, A.S.*

# Unparticles

*H. Georgi*

Hidden sector (HS) e.g., gauge theory with fermions and coupling  $g$



If  $g^* \gg 1 \rightarrow$  appearance of composite (confined) states of the HS particles (described by operators  $O_U$ )

hadrons  $\leftarrow$  quarks

*N. Krasnikov*  
*M. A. Stepanov*

Key difference:

Scale invariance  $\rightarrow$  continuous mass spectrum of confined states, Each has infinitesimal coupling with SM particles. Integral is finite

Individual (mass) modes: negligible effect

# Unparticle effects

Effects in solar neutrinos

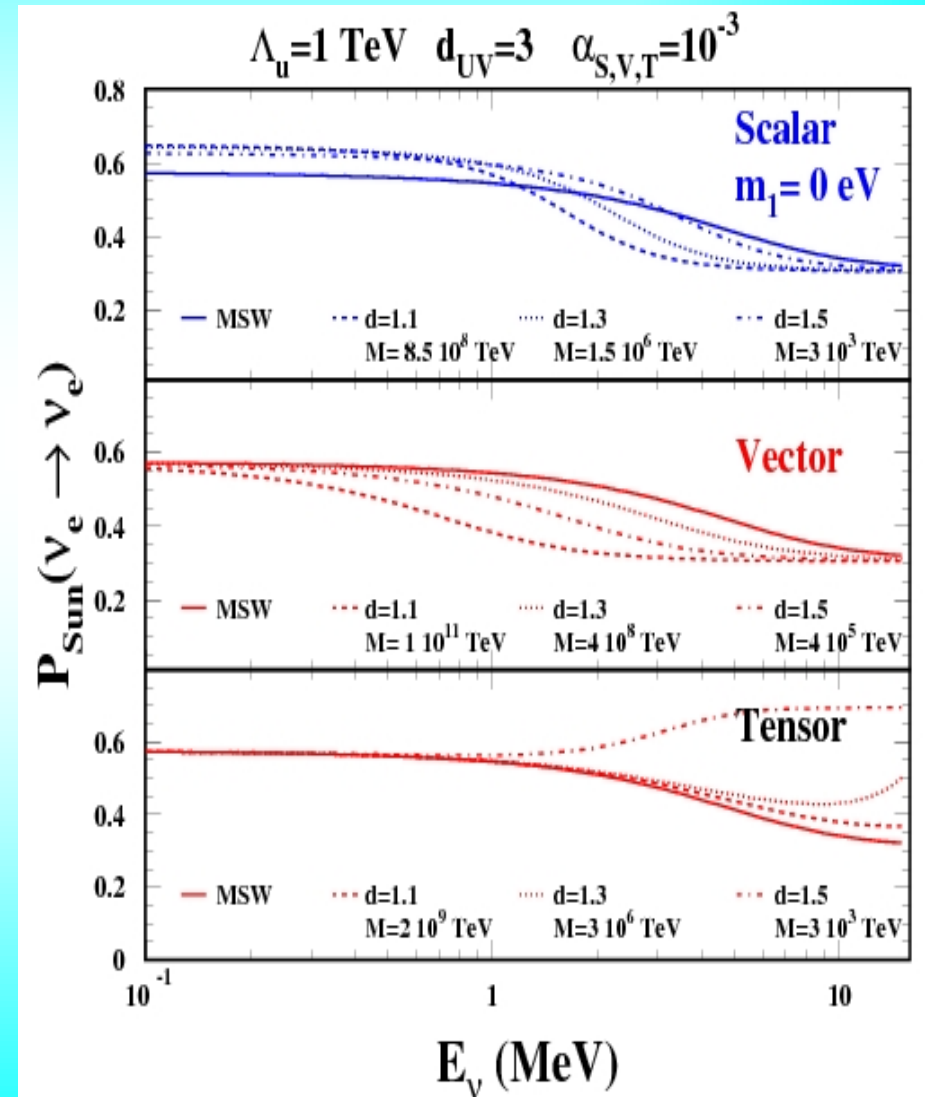
Neutrino decay:  $\nu_i \rightarrow \nu_j U$

*L. Anchordoqui, H. Goldberg*

Unparticle exchange:  
modify mater potential  
and effective neutrino mass  
→ modify survival probability

*M.C. Gonzalez-Garcia,  
P.C. de Holanda,  
R. Zukanovich-Funchal*

- M - mass of messenger
- $d_H$  - dimension of operator in hidden sector,
- d - dimension of unparticle operator
- $\Lambda_U$  - infrared fixed point



**Future which we  
know**

# Accomplish reconstruction of the neutrino mass and mixing spectrum

The program emerged more than 10 years ago

well motivated and elaborated

Reconstruction of neutrino mass matrix

May not reconstruct completely

- 1-3 mixing
- deviation of 2-3 mixing from maximal
- CP-phase
- $m_{ee}$ , nature of neutrinos
- absolute scale
- Majorana phases

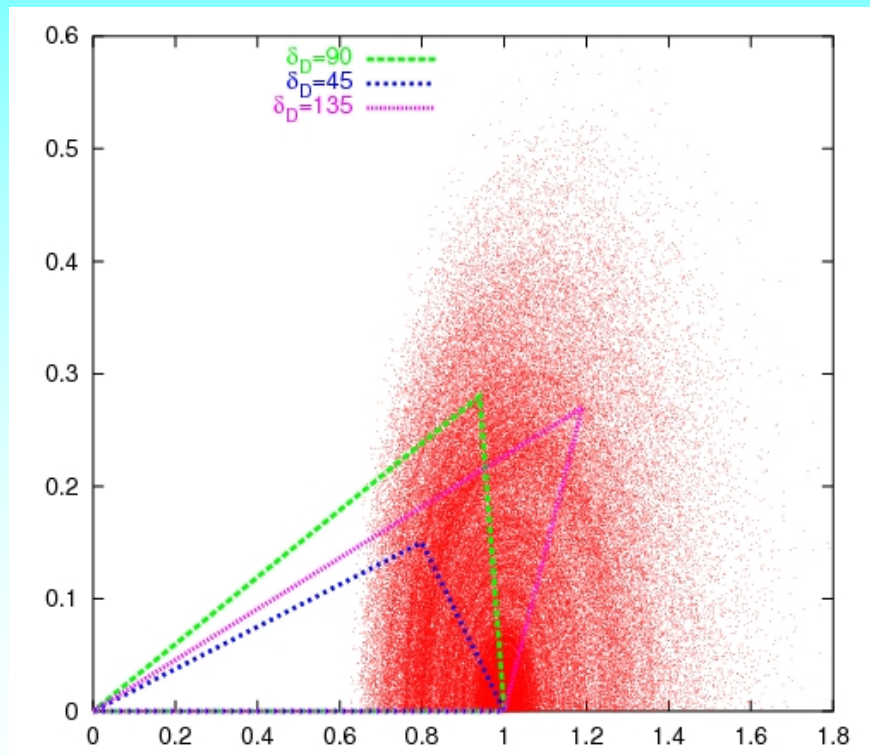
20 - 30 years?



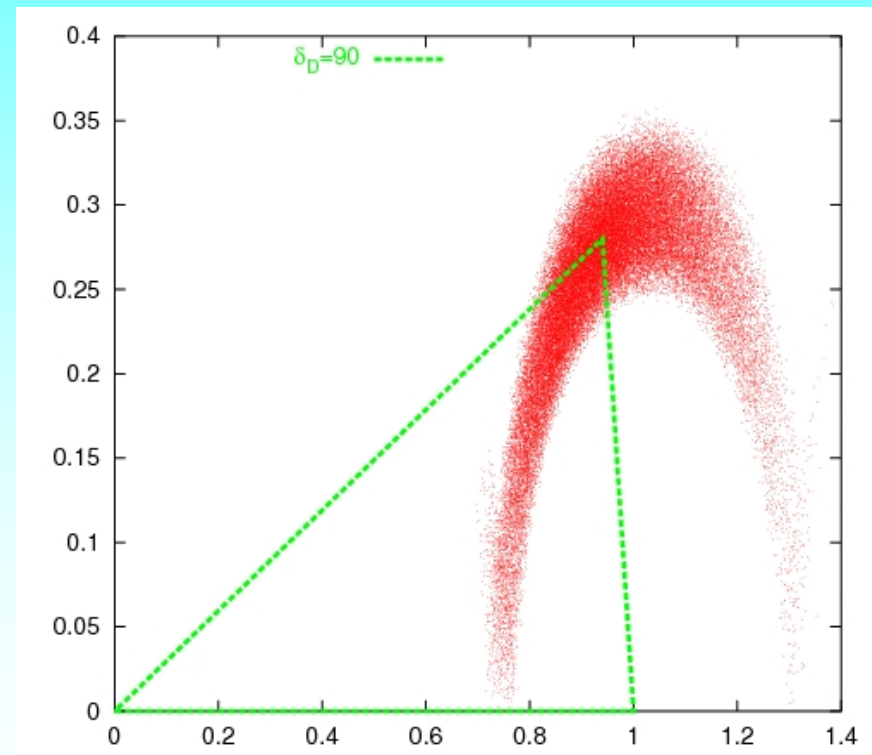
# Leptonic unitarity triangle

$e\mu$  - triangle  $\sin \theta_{13} = 0.15$

today



in future

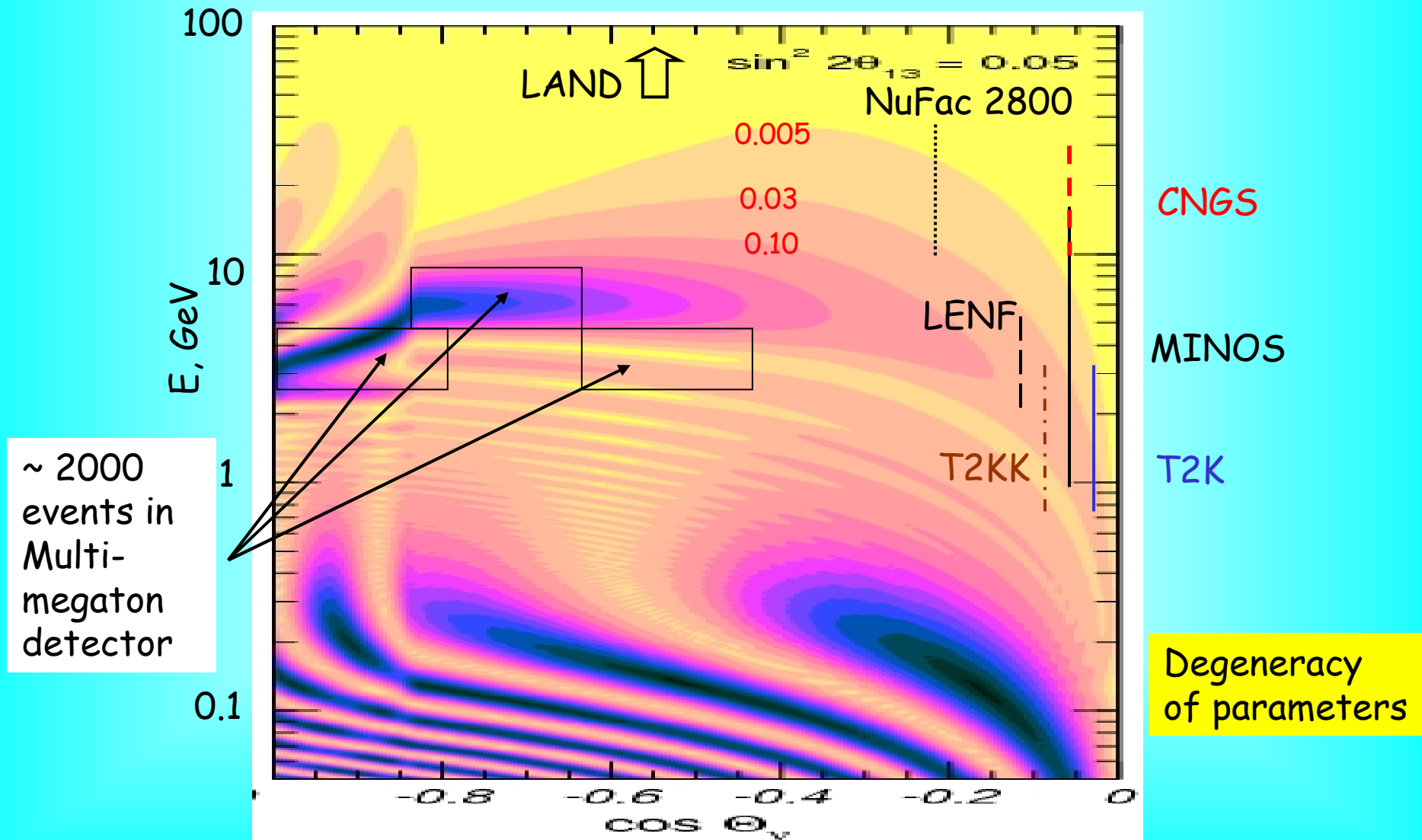


- illustration
- method to measure  $\delta$ , test of unitarity?

*Y. Farzan*

# Another approach?

Large atmospheric neutrino detectors

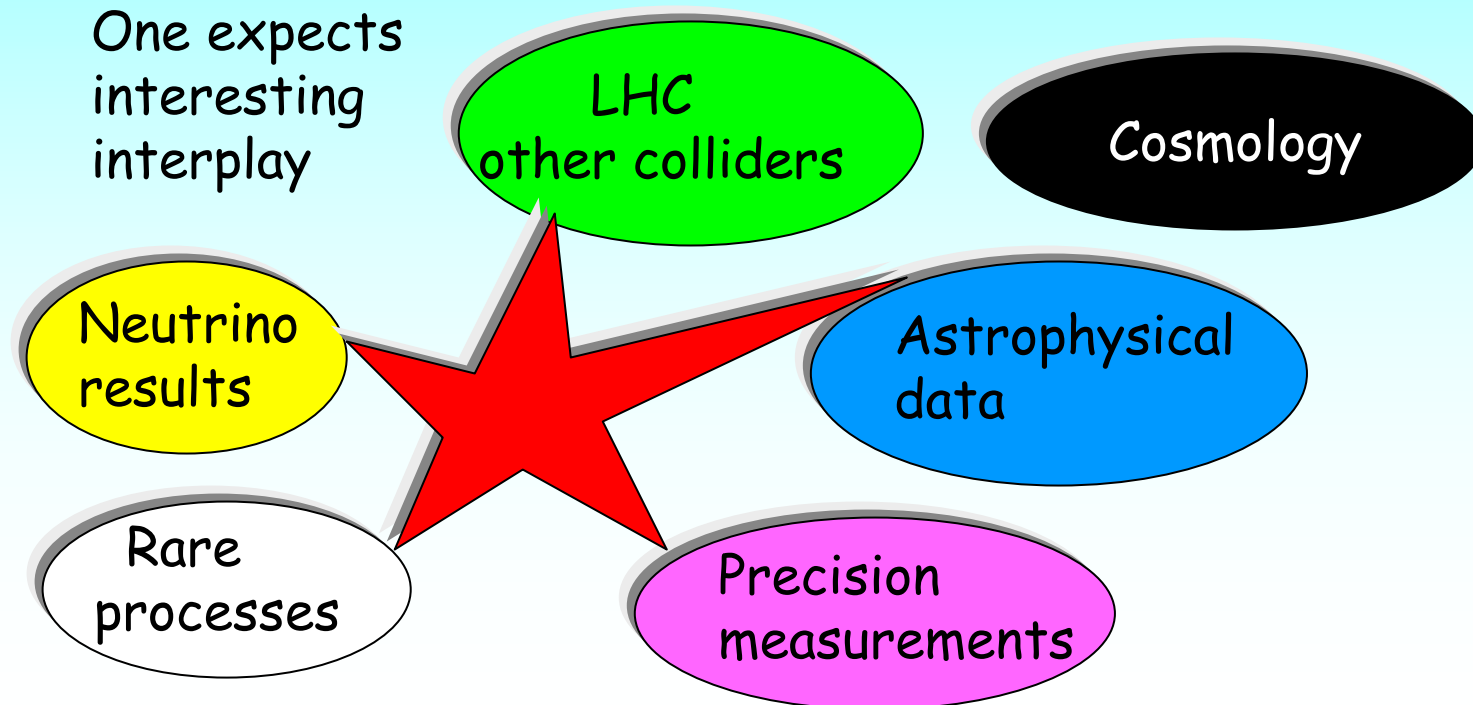


# Searches for new physics beyond standard scenario

improving bounds ...

Precision measurements  
in neutrino physics

One expects  
interesting  
interplay



# Detection of neutrino burst from Galactic supernova

Eventually...

May have very strong impact on  
Neutrino physics,  
Astrophysics  
Particle physics

Plausible: we discover  
something completely  
new.

Shed some light on

- explosion mechanism
- neutrino propagation and properties
- theory of conversion
- nucleosynthesis

# Cosmic neutrinos

Searches, measurements of  $\nu$ -fluxes

Further study of  $\nu$ - $\gamma$  connection.  
Implication of EM radiation data

New developments related to establishing the GZK cut-off and evidences of AGN as sources of the Cosmic rays

Cosmogenic,  
GZK- neutrinos

Astrophysics

Determination of neutrino parameters

# Future which we don't know

Supposed to be prophetic?

- According to citation index - 45 max

# ...but can imagine

Programs of previous neutrino conferences

Projecting from the past

Logic of the field

Seeing new experiments

New neutrino physics

History of neutrinos

**neutrinos are unpredictable!**

Looking at other fields

Out of blue

# Neutrinos & LHC

Expectations range from

Identification of the mechanism of neutrino mass generation

e.g. if the Higgs triplet with terascale mass and small VEV generates neutrino mass and mixing

to

Practically nothing

with conclusion that some EW scale mechanisms with certain values of parameters are excluded



# New experimental techniques

$\beta$ -beams

Neutrino factories with rotating  
(changing direction) beams

New neutrino  
sources

Multi-Megaton detectors

with flavor and charge  
identification  
low energy threshold

Higher precision - new physics  
open new horizons

Table-top experiments  
with sources

Radioactive nuclei  
Metastable atoms  
for neutrino detection

Mossbauer effect

Coherence

Detection of very  
weak signals

New methods of  
decrease of backgrounds

superconductivity  
Cryogenic detectors

# Toward the neutrino technologies

some proposed long time ago  
now less speculative  
now we know much more

not unique, multiple use

Monitoring of nuclear reactors

Tomography of the Earth

- absorption
- oscillation

Geo-neutrinos

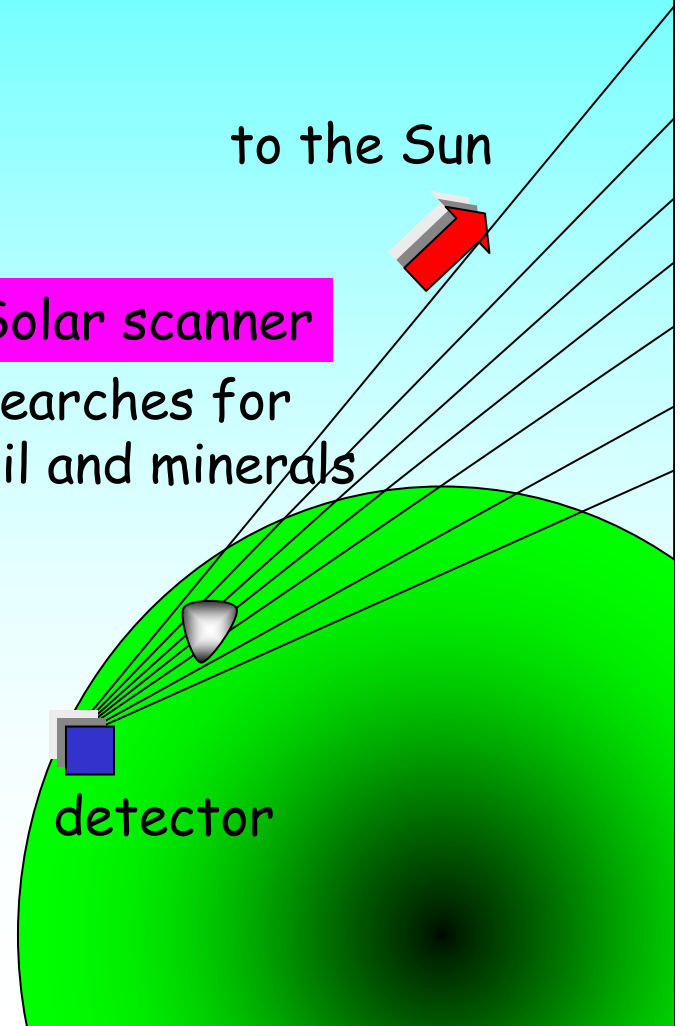
Mossbauer effect for neutrinos

Neutrino as a probe...

Neutrino communication systems  
Galactic communication

Solar scanner  
searches for  
oil and minerals

to the Sun



*J. Learnd,  
S. Pakvasa  
A. Zee*

# Neutrino structure of the Universe

Some work has already been done

Clumping of neutrinos depending on their masses

Neutrino halos, neutrino stars

Possible new interactions  
accelerons



Neutrino condensates

Superfluidity

Detection of relic neutrinos

Using metastable atoms  
and nuclei

*J. I Kapusta  
J R Bhatt  
U. Sarkar*

*M. Yoshimura  
P. Vogel et al*

# Summary

Neutrino physics is in the transition phase

Substantial territory is already ``captured" which can be described as the standard neutrino scenario

Tests of the standard scenario and searches for ``physics beyond" are the main motivations for further studies

Precision measurement and exploration of extreme conditions (energies, densities, distances) will open new horizons

# and what emerges?

Unclear implications of results to fundamental theory

- origin of neutrino mass and mixing
- existence of flavor symmetries, unification etc.

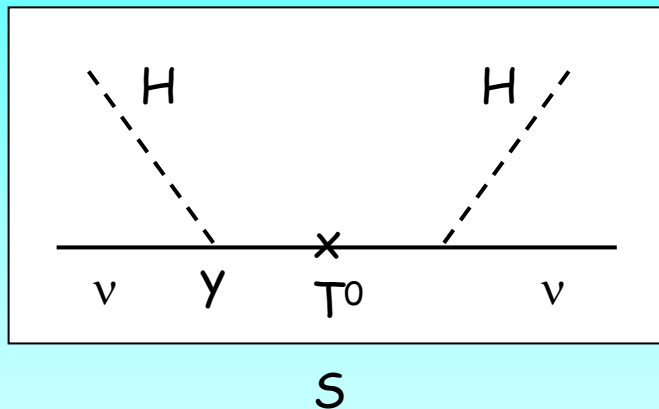
The question what should be done to have progress in understanding neutrino mass and mixing is already and will be a driving force of future developments

LHC and other HE experiments may clarify the situation

In spite of these problems we can start to think seriously about applications of neutrinos and neutrino technologies

# RH neutrinos at LHC

■ Type III seesaw:



SU(2) triplet:

$$T = \begin{pmatrix} T^+ \\ T^0 \\ T^- \end{pmatrix} \quad \begin{matrix} \gamma \sim 10^{-6} \\ M_T \sim 100 \text{ GeV} \end{matrix}$$

*B. Bajc  
G. Senjanovic  
M. Nemevsek  
P. Filiviez-Perez*

■ In SU(5):  $24_F$  T, S

$$\gamma \bar{5}_F 24_F 5_H + \gamma'/M \bar{5}_F 24_F 24_H 5_H$$

Type I + Type III  
one usual neutrino is mass less

■ **LHC:** EW production:  $W^{+*} \rightarrow T^+ T^0$ ,  $Z^* \rightarrow T^0 T^0$

Decay:  $T^0 \rightarrow W l$ ,  $T^0 \rightarrow Z \nu$ ,  $T^0 \rightarrow T^+ l^- \nu$

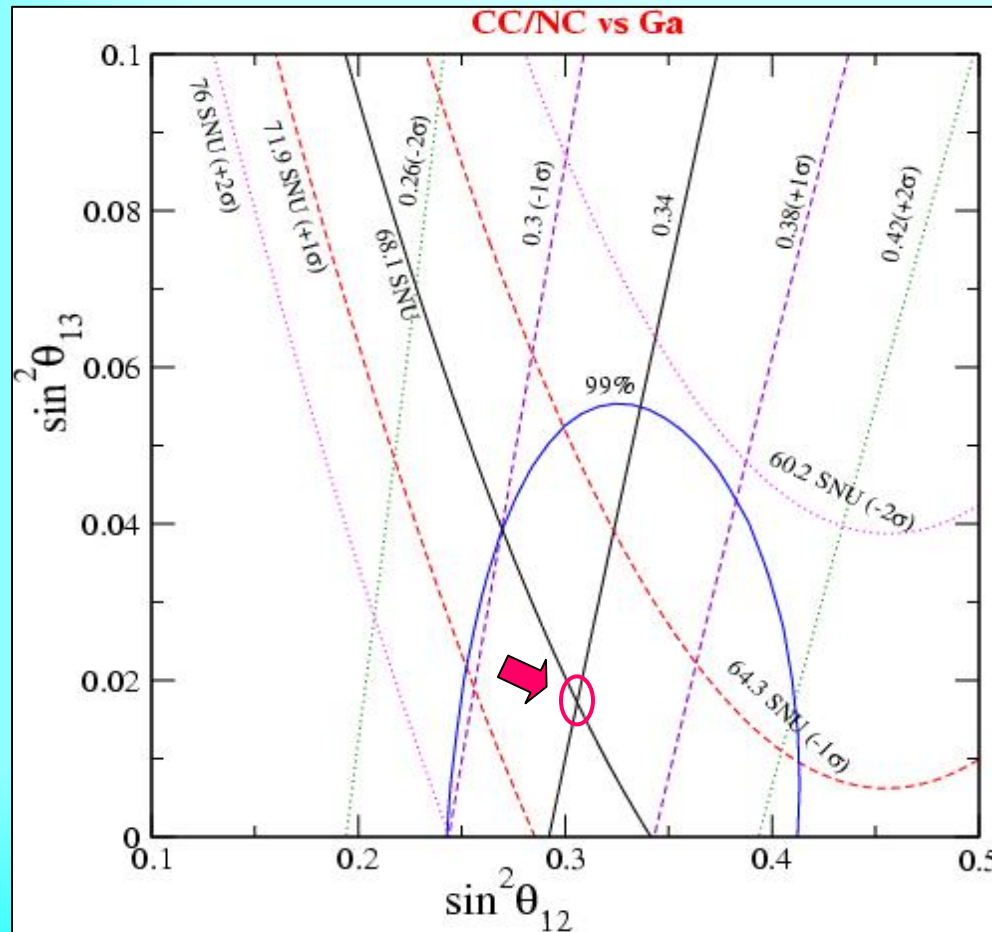
$$\tau \sim 10^{-16} - 10^{-13} \text{ sec}$$

$$\Gamma \sim (\text{mixing})^2$$

# Theta 1-3

Solar neutrinos: degeneracy of 1-2 and 1-3 mixing

*S. Goswami, A.S.*



$$\sin^2 \theta_{13} = 0.017 \pm 0.26$$

# The best scenario

Minimal number of assumptions:

Assumption 1:

$$\cos^2\theta_{13} = 1$$

Assumption 2:

$$1/\sin^2\theta_{23} = 2$$

Assumption 3:

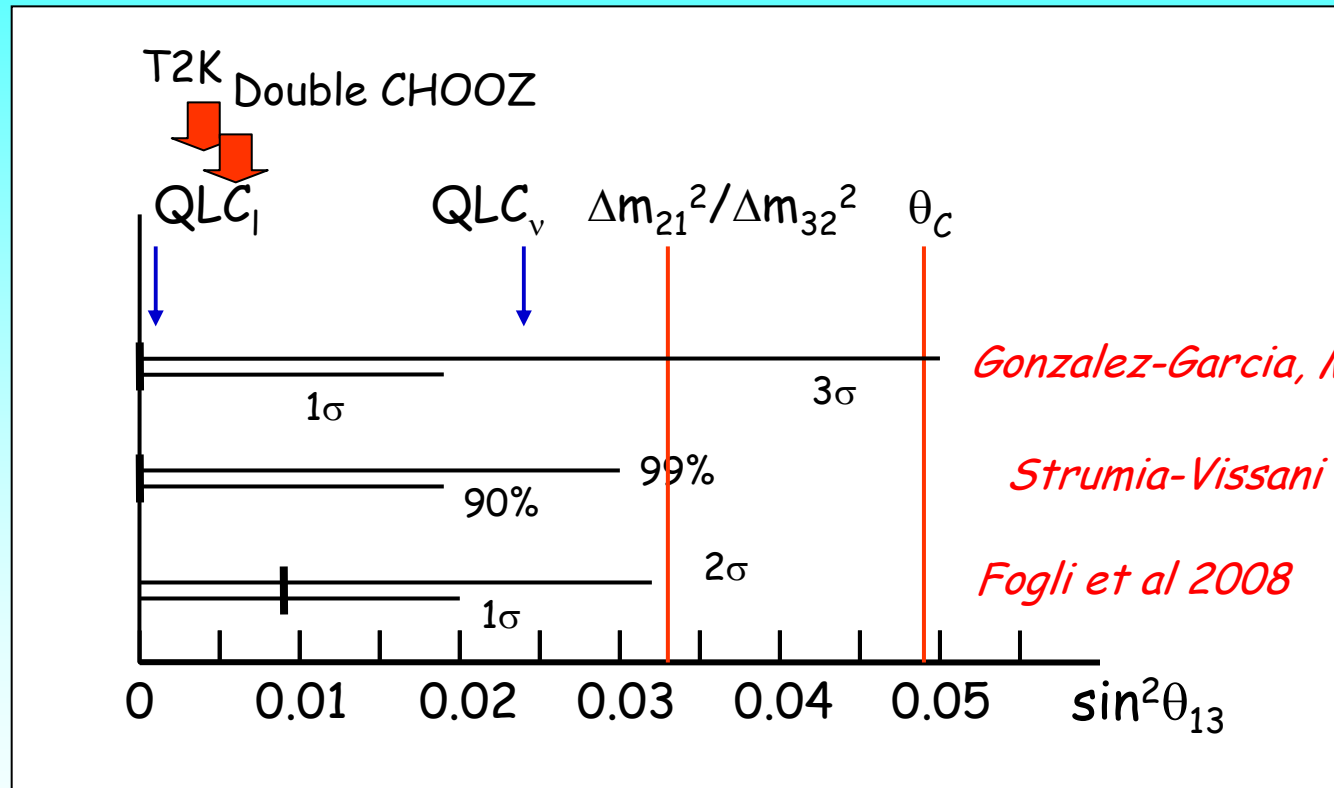
$$1/\sin^2\theta_{12} = 3$$

Plus possible small corrections...

Any model with smaller number of assumptions?



# 1-3 mixing

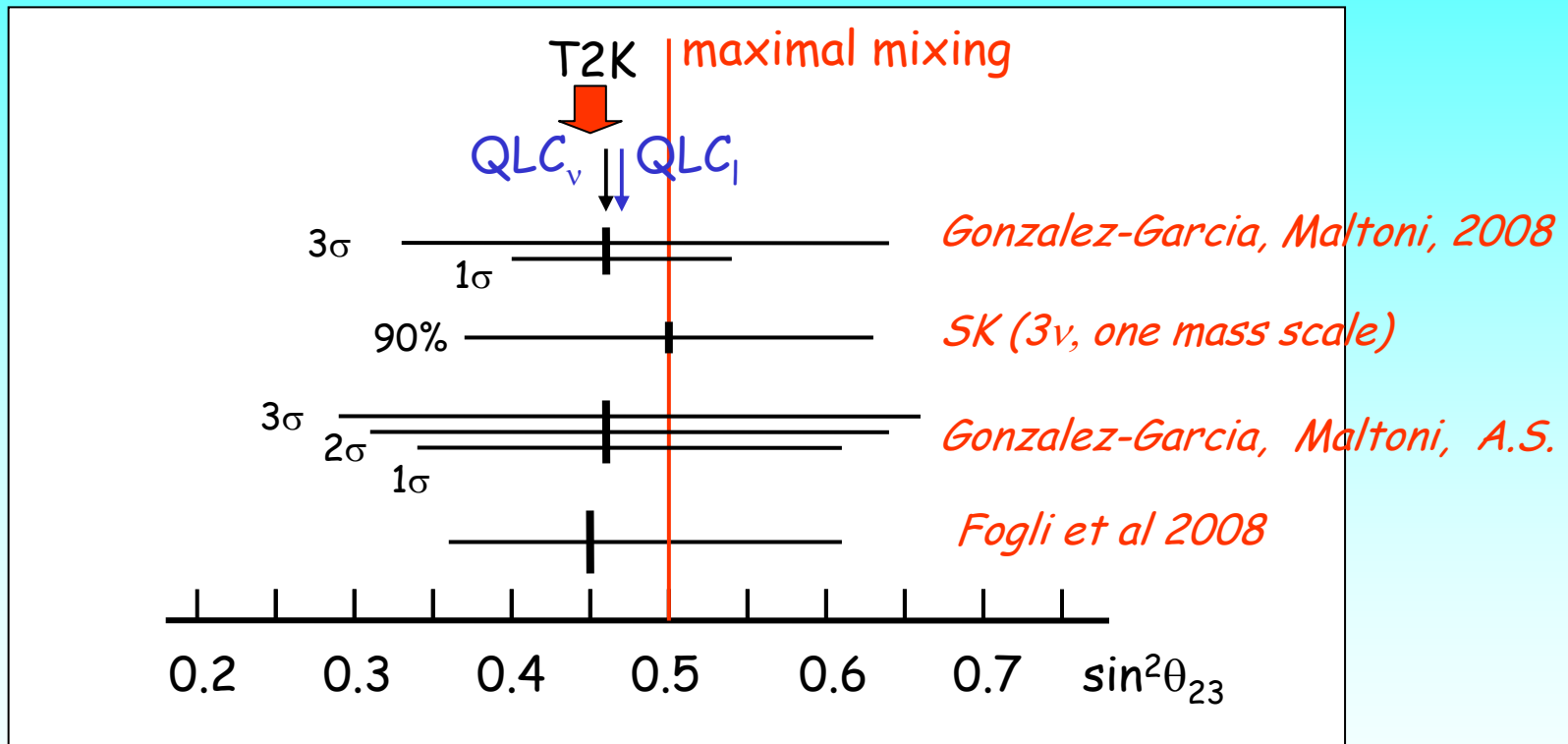


\* Non-zero central value (Fogli, et al): Atmospheric neutrinos, SK spectrum of multi-GeV e-like events

\* MINOS lead to stronger the bound on 1-3 mixing (G-G, M.)

# 2-3 mixing

SK:  $\sin^2 2\theta_{23} > 0.93$ , 90% C.L.



- \* in agreement with maximal, though all complete 3v - analyses show shift
- \* shift of the bfp from maximal is small
- \* still large deviation is allowed:  $(0.5 - \sin^2 \theta_{23}) / \sin^2 \theta_{23} \sim 40\%$   $2\sigma$

# Resume

No unique and convincing explanation of neutrino mass has been found

We are still in the explorative phase

There is no simple "one step" explanation/solution  
large number of assumptions

Inclusion of quark sector usually requires further  
complication of models *Grand unification?*

Perturbation approach:  
do not try to explain everything at once

Lowest order  
(symmetry)

Perturbations may  
have no symmetry  
or different symmetry

*Ambiguity to identify  
what is zero order*