

# Majorana Neutrinos and their Properties

S. T. Petcov

SISSA/INFN, Trieste, Italy,  
IPMU, University of Tokyo, Tokyo, Japan and  
INRNE, Bulgarian Academy of Sciences, Sofia, Bulgaria

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Compelling Evidences for  $\nu$ -Oscillations: 3- $\nu$  mixing

$$\nu_{lL} = \sum_{j=1}^3 U_{lj} \nu_{jL} \quad l = e, \mu, \tau.$$

## Three Neutrino Mixing

$$\nu_{lL} = \sum_{j=1}^3 U_{lj} \nu_{jL} .$$

$U$  is the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) neutrino mixing matrix,

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix}$$

- $U$  -  $n \times n$  unitary:

	$n$	2	3	4
mixing angles:	$\frac{1}{2}n(n-1)$	1	3	6

CP-violating phases:

$\nu_j$ - Dirac:	$\frac{1}{2}(n-1)(n-2)$	0	1	3
$\nu_j$ - Majorana:	$\frac{1}{2}n(n-1)$	1	3	6

$n = 3$ : 1 Dirac and

2 additional CP-violating phases, Majorana phases

## Majorana Neutrinos

Can be defined in QFT using fields or states.

Fields:  $\chi_k(x)$  - 4 component (spin 1/2), complex,  $m_k$

Majorana condition:

$$C (\bar{\chi}_k(x))^T = \xi_k \chi_k(x), \quad |\xi_k|^2 = 1$$

- Invariant under proper Lorentz transformations.
- Reduces by 2 the number of components in  $\chi_k(x)$ .

Implications:

$$U(1) : \chi_k(x) \rightarrow e^{i\alpha} \chi_k(x) - \text{impossible}$$

- $\chi_k(x)$  cannot absorb phases.
- $Q_{U(1)} = 0 : Q_{el} = 0, L_l = 0, L = 0, \dots$
- $\chi_k(x)$ : 2 spin states of a spin 1/2 absolutely neutral particle
- $\chi_k \equiv \bar{\chi}_k$

Propagators:  $\Psi(x)$ –Dirac,  $\chi(x)$ –Majorana

$$\langle 0|T(\Psi_\alpha(x)\bar{\Psi}_\beta(y))|0\rangle = S_{\alpha\beta}^F(x-y) ,$$

$$\langle 0|T(\Psi_\alpha(x)\Psi_\beta(y))|0\rangle = 0 , \quad \langle 0|T(\bar{\Psi}_\alpha(x)\bar{\Psi}_\beta(y))|0\rangle = 0 .$$

$$\langle 0|T(\chi_\alpha(x)\bar{\chi}_\beta(y))|0\rangle = S_{\alpha\beta}^F(x-y) ,$$

$$\langle 0|T(\chi_\alpha(x)\chi_\beta(y))|0\rangle = -\xi^* S_{\alpha\kappa}^F(x-y) C_{\kappa\beta} ,$$

$$\langle 0|T(\bar{\chi}_\alpha(x)\bar{\chi}_\beta(y))|0\rangle = \xi C_{\alpha\kappa}^{-1} S_{\kappa\beta}^F(x-y)$$

$$U_{CP} \chi(x) U_{CP}^{-1} = \eta_{CP} \gamma_0 \chi(x') , \quad \eta_{CP} = \pm i .$$

The type of massive neutrinos in a given theory is determined by the type of (effective) mass term  $\mathcal{L}_m^\nu(x)$  neutrinos have, **more precisely**, by the symmetries  $\mathcal{L}_m^\nu(x)$  and the total Lagrangian  $\mathcal{L}(x)$  of the theory have.

- Example: Majorana Mass Term of  $\nu_{lL}(x)$ ,  $l = e, \mu, \tau$

$$\mathcal{L}_M^\nu(x) = -\frac{1}{2} \overline{\nu_{lR}^c}(x) M_{ll} \nu_{lL}(x) + h.c. , \nu_{lR}^c \equiv C (\overline{\nu_{lL}(x)})^\top$$

- If  $M_{ll} \neq 0$ ,  $L_l \neq const.$ ,  $L \neq const.$
- $\nu_{lL}(x)$ -fermions:  $M = M^\top$ , complex.

$$M^{diag} = U^\top M U , U - \text{unitary (congruent transformation)}; U \equiv U_{PMNS}$$

$$\chi_j(x) = U_{jl}^\dagger \nu_{lL}(x) + U_{jl}^* \nu_{lR}^c = C (\overline{\chi_j(x)})^\top , m_j \neq 0$$

CP-invariance:  $M^* = M$ ,  $M$  - real, symmetric.

$$M^{diag} = (m'_1, m'_2, m'_3): m'_j = \rho_j m_j, m_j \geq 0, \rho_j = \pm 1$$

$$\chi_j: m_j \geq 0: \eta_{CP}(\chi_j) = i\rho_j$$

- Dirac Neutrinos: Dirac Mass Term, requires  $\nu_R(x)$  -  $SU(2)_L$  singlet RH  $\nu$  fields

$$\mathcal{L}_D^\nu(x) = -\overline{\nu_{lR}}(x) M_{Dl} \nu_{lL}(x) + h.c. , M_D - \text{complex}$$

- $\mathcal{L}_D^\nu(x)$  conserves  $L$ :  $L = const.$

$$M_D = V M_D^{diag} W^\dagger , V, U - \text{unitary (bi-unitary transformation)}, W \equiv U_{PMNS}$$

Special Properties of the Currents of  $\chi(x)$ –Majorana:

$$\bar{\chi}(x)\gamma_\alpha\chi(x) = 0 : \quad Q_{U(1)} = 0 \quad (Q_{U(1)}(\Psi) \neq 0) ;$$

Has important implications, e.g. for SUSY DM (neutralino) abundance determination (calculation).

$$\bar{\chi}(x)\sigma_{\alpha\beta}\chi(x) = 0 : \quad \mu_\chi = 0 \quad (\mu_\Psi \neq 0)$$

$$\bar{\chi}(x)\sigma_{\alpha\beta}\gamma_5\chi(x) = 0 : \quad d_\chi = 0 \quad (d_\Psi \neq 0, \text{ if } CP \text{ is not conserved})$$

$\chi(x)$  cannot couple to a real photon (field).

$\chi(x)$  couples to a virtual photon through an anapole moment:

$$(g_{\alpha\beta}q^2 - q_\alpha q_\beta)\gamma_\beta\gamma_5 F_a(q^2).$$

Properties of Currents Formed by  $\chi_1(x)$ ,  $\chi_2(x)$ :  $\chi_2 \rightarrow \chi_1 + \gamma$ ,  $\chi_2 \rightarrow \chi_1 \chi_1 \chi_1$ , etc.

$$\bar{\chi}_1(x) \gamma_\alpha (v - a \gamma_5) \chi_2(x) \quad (\bar{\chi}_1(x) \gamma^\alpha (1 - \gamma_5) \chi_1(x), \dots) :$$

- CP is conserved:  $v = 0$  ( $a = 0$ ) if  $\eta_{1CP} = \eta_{2CP}$  ( $\eta_{1CP} = -\eta_{2CP}$ )
- CP is not conserved:  $v \neq 0$ ,  $a \neq 0$

(Has important implications also, e.g. for SUSY neutralino phenomenology:  
 $e^+ + e^- \rightarrow \chi_1 + \chi_2$ ,  $\chi_2 \rightarrow \chi_1 + l^+ + l^-$ , etc.)

$$\bar{\chi}_1(x) \sigma_{\alpha\beta} (\mu_{12} - d_{12} \gamma_5) \chi_2(x) \quad (F^{\alpha\beta}(x)) :$$

- CP is conserved:  $\mu_{12} = 0$  ( $d_{12} = 0$ ) if  $\eta_{1CP} = \eta_{2CP}$  ( $\eta_{1CP} = -\eta_{2CP}$ )
- CP is not conserved:  $\mu_{12} \neq 0$ ,  $d_{12} \neq 0$



## Dirac - Majorana Relation (if any...)

Majorana Mass Term of  $\nu_{lL}(x)$ ,  $l = e, \mu, \tau$ , can lead to Dirac neutrinos with definite mass if it conserves some lepton charge:

$$\mathcal{L}_M^\nu(x) = -\frac{1}{2} \overline{\nu_{lR}^c}(x) M_{ll} \nu_{lL}(x) + h.c. , \quad \nu_{lR}^c \equiv C (\overline{\nu_{lL}(x)})^\top$$

$\mathcal{L}_M^\nu(x)$  conserves, e.g.  $L' = L_e - L_\mu - L_\tau$  if only  $M_{e\mu} = M_{\mu e}, M_{e\tau} = M_{\tau e} \neq 0$   
S.T.P., 1982

• Dirac  $\nu$ ,  $\Psi$ , is equivalent to two Majorana  $\nu$ 's,  $\chi_{1,2}$ , having the same (positive) mass, opposite CP-parities, and which are “maximally mixed”:

$$\Psi(x) = \frac{\chi_1 + \chi_2}{\sqrt{2}}, \quad m_1 = m_2 = m_D > 0, \eta_{jCP} = i\rho_j, \rho_1 = -\rho_2 \quad (C (\overline{\chi_j})^\top = \rho_j \chi_j)$$

$$\text{Example ZKM } \nu: \nu_{eL}(x) = \Psi_L = \frac{\chi_{1L} + \chi_{2L}}{\sqrt{2}}, \quad \nu_{\mu L}(x) = \Psi_L^C = \frac{\chi_{1L} - \chi_{2L}}{\sqrt{2}}$$

• Pseudo-Dirac Neutrino: the symmetry of  $\mathcal{L}_M^\nu(x)$  is not a symmetry of  $\mathcal{L}_{tot}(x)$

Suppose:  $\nu_{eL}(x) = \Psi_L = (\chi_{1L} + \chi_{2L})/\sqrt{2}$ , and to “leading order”  $m_1 = m_2$ , but due to “higher order” corrections  $m_1 \neq m_2$ ,  $|m_2 - m_1| \equiv |\Delta m| \ll m_{1,2}$

All Majorana effects  $\sim \Delta m$

• Suppose:  $m_1 = m_2$ ,  $\rho_1 = -\rho_2$ , but  $\chi_{1,2}$  are not maximally mixed:

$$\nu_{eL}(x) = \chi_{1L} \cos \phi + \chi_{2L} \sin \phi = \Psi_L \cos \phi' + \Psi_L^C \sin \phi'$$

All Majorana effects are  $\sim m_D \cos \phi' \sin \phi'$

Pontecorvo, 1957, 1958:

$$\nu(x) = \frac{\chi_1 + \chi_2}{\sqrt{2}}, \quad m_1 \neq m_2 > 0, \eta_{1CP} = -\eta_{2CP}$$

$\chi_{1,2}$  - Majorana, maximal mixing .

Maki, Nakagawa, Sakata, 1962:

$$\nu_{eL}(x) = \Psi_{1L} \cos \theta_C + \Psi_{2L} \sin \theta_C,$$

$$\nu_{\mu L}(x) = -\Psi_{1L} \sin \theta_C + \Psi_{2L} \cos \theta_C,$$

$\Psi_{1,2}$  - Dirac (composite),  $\theta_C$ - the Cabibbo angle .

# PMNS Matrix: Standard Parametrization

$$U = V \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\frac{\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{i\frac{\alpha_{31}}{2}} \end{pmatrix}$$

$$V = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

- $s_{ij} \equiv \sin \theta_{ij}$ ,  $c_{ij} \equiv \cos \theta_{ij}$ ,  $\theta_{ij} = [0, \frac{\pi}{2}]$ ,
- $\delta$  - Dirac CP-violation phase,  $\delta = [0, 2\pi]$ ,
- $\alpha_{21}$ ,  $\alpha_{31}$  - the two Majorana CP-violation phases.
- $\Delta m_{\odot}^2 \equiv \Delta m_{21}^2 \cong 7.65 \times 10^{-5} \text{ eV}^2 > 0$ ,  $\sin^2 \theta_{12} \cong 0.305$ ,  $\cos 2\theta_{12} \gtrsim 0.26$  ( $2\sigma$ ),
- $|\Delta m_{\text{atm}}^2| \equiv |\Delta m_{31}^2| \cong 2.4$  (2.5)  $\times 10^{-3} \text{ eV}^2$ ,  $\sin^2 2\theta_{23} \cong 1$ ,
- $\theta_{13}$  - the CHOOZ angle:  $\sin^2 \theta_{13} < 0.033$  (0.056 (0.063))  $2\sigma$  ( $3\sigma$ ).

A.Bandyopadhyay, S.Choubey, S.Goswami, S.T.P., D.P.Roy, arXiv:0804.4857;

T. Schwetz *et al.*, arXiv:0808.2016

## Neutrino Oscillation Parameters

parameter	bf	1 $\sigma$ acc.	2 $\sigma$ range	3 $\sigma$ range
$\Delta m_{21}^2$ [ $10^{-5}$ eV $^2$ ]	7.65	3%	7.25 – 8.11	7.05 – 8.34
$ \Delta m_{31}^2 $ [ $10^{-3}$ eV $^2$ ]	2.4	5%	2.18 – 2.64	2.07 – 2.75
$\sin^2 \theta_{12}$	0.304	7%	0.27 – 0.35	0.25 – 0.37
$\sin^2 \theta_{23}$	0.50	14%	0.39 – 0.63	0.36 – 0.67
$\sin^2 \theta_{13}$	$0.01^{+0.016}_{-0.011}$	–	$\leq 0.040$	$\leq 0.056$

Best fit values (bf), relative accuracies at 1 $\sigma$ , and 2 $\sigma$  and 3 $\sigma$  allowed ranges of three-flavor neutrino oscillation parameters from a combined analysis of global data.

T. Schwetz *et al.*, arXiv:0808.2016[hep-ph]

- $\text{sgn}(\Delta m_{\text{atm}}^2) = \text{sgn}(\Delta m_{31}^2)$  not determined

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2 > 0, \quad \text{normal mass ordering}$$

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{32}^2 < 0, \quad \text{inverted mass ordering}$$

Convention:  $m_1 < m_2 < m_3$  - **NMO**,  $m_3 < m_1 < m_2$  - **IMO**

$$m_1 \ll m_2 < m_3, \quad \text{NH,}$$

$$m_3 \ll m_1 < m_2, \quad \text{IH,}$$

$$m_1 \cong m_2 \cong m_3, \quad m_{1,2,3}^2 \gg \Delta m_{\text{atm}}^2, \quad \text{QD; } m_j \gtrsim 0.10 \text{ eV.}$$

- Dirac phase  $\delta$ :  $\nu_l \leftrightarrow \nu_{l'}, \bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}, l \neq l'$ ;  $A_{\text{CP}}^{(l,l')} \propto J_{\text{CP}} \propto \sin \theta_{13} \sin \delta$

- Majorana phases  $\alpha_{21}, \alpha_{31}$ :

–  $\nu_l \leftrightarrow \nu_{l'}, \bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}$  not sensitive;

S.M. Bilenky, J. Hosek, S.T.P., 1980;  
P. Langacker, S.T.P., G. Steigman, S. Toshev, 1987

–  $|\langle m \rangle|$  in  $(\beta\beta)_{0\nu}$ -decay depends on  $\alpha_{21}, \alpha_{31}$ ;

–  $\Gamma(\mu \rightarrow e + \gamma)$  etc. in SUSY theories depend on  $\alpha_{21,31}$ ;

– BAU, leptogenesis scenario:  $\alpha_{21,31}$  !

# Absolute Neutrino Mass Measurements

The Troitzk and Mainz  ${}^3\text{H}$   $\beta$ -decay experiments

$$m_{\nu_e} < 2.3 \text{ eV} \quad (95\% \text{ C.L.})$$

There are prospects to reach sensitivity

$$\text{KATRIN :} \quad m_{\nu_e} \sim 0.2 \text{ eV}$$

Cosmological and astrophysical data: the WMAP result combined with data from large scale structure surveys (2dFGRS, SDSS)

$$\sum_j m_j \equiv \Sigma < (0.4 - 1.4) \text{ eV}$$

The WMAP and future PLANCK experiments can be sensitive to

$$\sum_j m_j \cong 0.4 \text{ eV}$$

Data on weak lensing of galaxies by large scale structure, combined with data from the WMAP and PLANCK experiments may allow to determine

$$\sum_j m_j : \quad \delta \cong 0.04 \text{ eV.}$$

$\nu_{\odot}$ ,  $\Delta m_{\text{atm}}^2$ , CHOOZ Data:

- $\theta_{12} = \theta_{\odot} \sim \frac{\pi}{6}$ ,  $\theta_{23} = \theta_{\text{atm}} \cong \frac{\pi}{4}$ ,  $\theta_{13} < \frac{\pi}{12}$

$$U_{\text{PMNS}} \cong \begin{pmatrix} \frac{\sqrt{3}}{2} & \frac{1}{2} & \epsilon \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & \frac{1}{\sqrt{2}} \\ \frac{1}{2\sqrt{2}} & -\frac{\sqrt{3}}{2\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix} .$$

Very different from the CKM-matrix!

- $\cos \theta_{12} \cong \cos\left(\frac{\pi}{4} - \frac{\pi}{12}\right) = \frac{1}{\sqrt{2}}(1 + \lambda)$ ,  $\sin \theta_{12} \cong \frac{1}{\sqrt{2}}(1 - \lambda)$ ,
- $\lambda \cong (0.20 - 0.25)$ :  $\theta_{\odot} + \theta_c = \pi/4$  ?

Natural Possibility:

$$U = U_{\text{lep}}^\dagger(\lambda) U_{\text{bim(tri)}}$$

with

$$U_{\text{bim}} = \begin{pmatrix} \frac{1}{\sqrt{2}} & \frac{1}{\sqrt{2}} & 0 \\ -\frac{1}{2} & \frac{1}{2} & \frac{1}{\sqrt{2}} \\ \frac{1}{2} & -\frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix} \quad U_{\text{tri}} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \sqrt{\frac{1}{3}} & 0 \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & -\sqrt{\frac{1}{2}} \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & \sqrt{\frac{1}{2}} \end{pmatrix}.$$

- $U_{\text{lep}}^\dagger(\lambda)$  - from diagonalization of the  $l^-$  mass matrix,
- $U_{\text{bim(tri)}}$  - from diagonalization of the  $\nu$ -mass matrix

Further,  $\Delta m_{\odot}^2 \ll |\Delta m_{\text{atm}}^2|$ .

- $U_{\text{bim}}$  can be associated with a symmetry:

$$L' = L_e - L_\mu - L_\tau$$

S.T.P., 1982

- $U_{\text{bim(tri)}}$  can be associated with a  $\mu - \tau$  symmetry of  $M_\nu$   
T. Fukuyama, H. Nishiura, 1997; R.N. Mohapatra, S. Nussinov, 1999;...

These symmetries cannot be exact.



In the case of conserved  $L' = L_e - L_\mu - L_\tau$ :

$$M = \begin{pmatrix} 0 & M_{e\mu} & M_{e\tau} \\ M_{e\mu} & 0 & 0 \\ M_{e\tau} & 0 & 0 \end{pmatrix}$$

$$\theta_{12} = \pi/4, \theta_{13} = 0, \tan \theta_{23} = M_{e\tau}/M_{e\mu},$$

$m_3 = 0$  - spectrum with IH,  $m_1 = m_2$ ,  $\chi_{1,2}$  - equivalent to one Dirac  $\nu$ ,  $\Psi$ .

Adding  $L'$ -breaking term, e.g.  $M_{ee}$ ,  $|M_{ee}|/\sqrt{M_{e\mu}^2 + M_{e\tau}^2} \sim 0.01$ , leads to  $m_1 \neq m_2$  compatible with  $\Delta m_{\odot}^2$ .

## Rephasing Invariants Associated with CPVP

Dirac phase  $\delta$ :

$$J_{CP} = \text{Im} \{ U_{e1} U_{\mu 2} U_{e2}^* U_{\mu 1}^* \} .$$

C. Jarlskog, 1985 (for quarks)

CP-, T- violation effects in neutrino oscillations

P. Krastev, S.T.P., 1988

Majorana phases  $\alpha_{21}, \alpha_{31}$ :

$$S_1 = \text{Im} \{ U_{e1} U_{e3}^* \}, \quad S_2 = \text{Im} \{ U_{e2} U_{e3}^* \} \quad (\text{not unique}); \quad \text{or}$$
$$S'_1 = \text{Im} \{ U_{\tau 1} U_{\tau 2}^* \}, \quad S'_2 = \text{Im} \{ U_{\tau 2} U_{\tau 3}^* \}$$

J.F. Nieves and P. Pal, 1987, 2001

G.C. Branco et al., 1986

J.A. Aguilar-Saavedra and G.C. Branco, 2000

**CP-violation:** both  $\text{Im} \{ U_{e1} U_{e3}^* \} \neq 0$  and  $\text{Re} \{ U_{e1} U_{e3}^* \} \neq 0$  .

$S_1, S_2$  appear in  $|\langle m \rangle|$  in  $(\beta\beta)_{0\nu}$ -decay.

In general,  $J_{CP}, S_1$  and  $S_2$  are independent.

## Dirac CP-Nonconservation: $\delta$ in $U_{\text{PMNS}}$

Observable manifestations in

$$\nu_l \leftrightarrow \nu_{l'} , \quad \bar{\nu}_l \leftrightarrow \bar{\nu}_{l'} , \quad l, l' = e, \mu, \tau$$

- not sensitive to Majorana CPVP  $\alpha_{21}, \alpha_{31}$

S.M. Bilenky, J. Hosek, S.T.P., 1980;  
P. Langacker et al., 1987

$$A(\nu_l \leftrightarrow \nu_{l'}) = U_{l'j} e^{-i(E_j t - p_j x)} U_{jl}^\dagger$$

$$U = VP : P_j e^{-i(E_j t - p_j x)} P_j^* = e^{-i(E_j t - p_j x)}$$

$P$  - diagonal matrix of Majorana phases.

The result is valid also in the case of oscillations in matter.

## Future Progress

- Determination of the nature - Dirac or Majorana, of  $\nu_j$  .
- Determination of  $\text{sgn}(\Delta m_{\text{atm}}^2)$ , type of  $\nu$ - mass spectrum

$$m_1 \ll m_2 < m_3, \quad \text{NH,}$$

$$m_3 \ll m_1 < m_2, \quad \text{IH,}$$

$$m_1 \cong m_2 \cong m_3, \quad m_{1,2,3}^2 \gg \Delta m_{\text{atm}}^2, \quad \text{QD; } m_j \gtrsim 0.10 \text{ eV.}$$

- Determining, or obtaining significant constraints on, the absolute scale of  $\nu_j$ -masses, or  $\min(m_j)$ .
- Status of the CP-symmetry in the lepton sector: violated due to  $\delta$  (Dirac), and/or due to  $\alpha_{21}, \alpha_{31}$  (Majorana)?
- Measurement of, or improving by at least a factor of (5 - 10) the existing upper limit on,  $\sin^2 \theta_{13}$ .
- High precision determination of  $\Delta m_{\odot}^2, \theta_{\odot}, \Delta m_{\text{atm}}^2, \theta_{\text{atm}}$ .
- Searching for possible manifestations, other than  $\nu_l$ -oscillations, of the non-conservation of  $L_l, l = e, \mu, \tau$ , such as  $\mu \rightarrow e + \gamma, \tau \rightarrow \mu + \gamma$ , etc. decays.

- Understanding at fundamental level the mechanism giving rise to the  $\nu$ - masses and mixing and to the  $L_l$ -non-conservation. Includes understanding
  - the origin of the observed patterns of  $\nu$ -mixing and  $\nu$ -masses ;
  - the physical origin of  $CPV$  phases in  $U_{PMNS}$  ;
  - Are the observed patterns of  $\nu$ -mixing and of  $\Delta m_{21,31}^2$  related to the existence of a new symmetry?
  - Is there any relations between  $q$ -mixing and  $\nu$ - mixing? Is  $\theta_{12} + \theta_c = \pi/4$  ?
  - Is  $\theta_{23} = \pi/4$ , or  $\theta_{23} > \pi/4$  or else  $\theta_{23} < \pi/4$ ?
  - Is there any correlation between the values of  $CPV$  phases and of mixing angles in  $U_{PMNS}$ ?
- Progress in the theory of  $\nu$ -mixing might lead to a better understanding of the origin of the BAU.
  - Can the Majorana and/or Dirac CPVP in  $U_{PMNS}$  be the leptogenesis CPV parameters at the origin of BAU?

If  $\nu_j$  – Majorana particles,  $U_{\text{PMNS}}$  contains (3- $\nu$  mixing)

$\delta$ -Dirac,  $\alpha_{21}$ ,  $\alpha_{31}$  - Majorana physical CPV phases

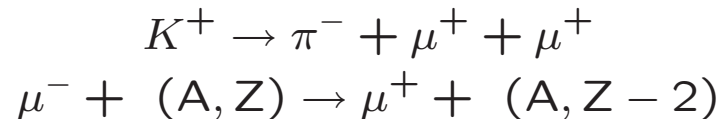
$\nu$ -oscillations  $\nu_l \leftrightarrow \nu_{l'}$ ,  $\bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}$ ,  $l, l' = e, \mu, \tau$ ,

- are not sensitive to the nature of  $\nu_j$ ,

S.M. Bilenky et al., 1980;  
P. Langacker et al., 1987

- provide information on  $\Delta m_{jk}^2 = m_j^2 - m_k^2$ , but not on the absolute values of  $\nu_j$  masses.

The Majorana nature of  $\nu_j$  can manifest itself in the existence of  $\Delta L = \pm 2$  processes:



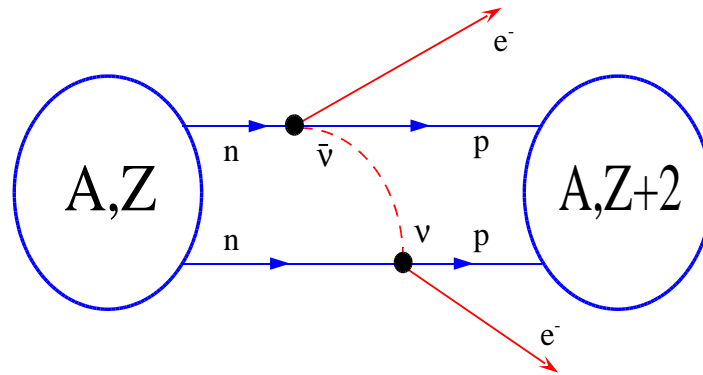
The process most sensitive to the possible Majorana nature of  $\nu_j$  -  $(\beta\beta)_{0\nu}$ -decay



of even-even nuclei,  $^{48}\text{Ca}$ ,  $^{76}\text{Ge}$ ,  $^{82}\text{Se}$ ,  $^{100}\text{Mo}$ ,  $^{116}\text{Cd}$ ,  $^{130}\text{Te}$ ,  $^{136}\text{Xe}$ ,  $^{150}\text{Nd}$ .

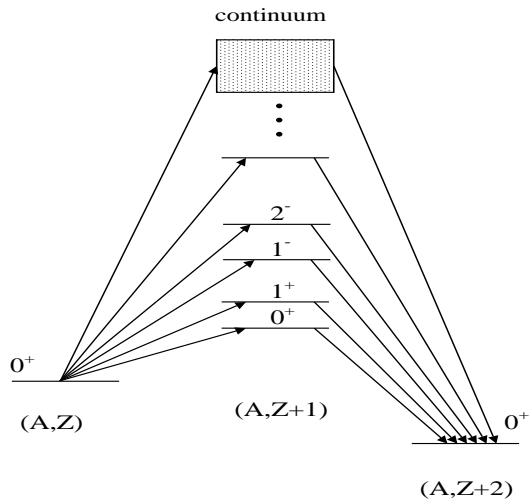
$2n$  from  $(A, Z)$  exchange a virtual Majorana  $\nu_j$  (via the CC weak interaction) and transform into  $2p$  of  $(A, Z+2)$  and two free  $e^-$ .

# Nuclear $0\nu\beta\beta$ -decay



strong in-medium modification of the basic process

$$dd \rightarrow uue^-e^-(\bar{\nu}_e\bar{\nu}_e)$$



virtual excitation  
of states of all multipolarities  
in  $(A, Z+1)$  nucleus

## $(\beta\beta)_{0\nu}$ –Decay Experiments:

- Majorana nature of  $\nu_j$
- Type of  $\nu$ –mass spectrum (NH, IH, QD)
- Absolute neutrino mass scale

${}^3\text{H}$   $\beta$ -decay , cosmology:  $m_{\nu}$  (QD, IH)

- CPV due to Majorana CPV phases

## $\nu_j$ – Dirac or Majorana particles, fundamental problem

$\nu_j$ –Dirac: conserved lepton charge exists,  $L = L_e + L_\mu + L_\tau$ ,  $\nu_j \neq \bar{\nu}_j$

$\nu_j$ –Majorana: no lepton charge is **exactly** conserved,  $\nu_j \equiv \bar{\nu}_j$

The observed patterns of  $\nu$ –mixing and of  $\Delta m_{\text{atm}}^2$  and  $\Delta m_{\odot}^2$  can be related to Majorana  $\nu_j$  and an **approximate** symmetry:

$$L' = L_e - L_\mu - L_\tau$$

S.T.P., 1982

See-saw mechanism:  $\nu_j$ – Majorana

Establishing that  $\nu_j$  are Majorana particles would be as important as the discovery of  $\nu$ – oscillations.



$$A(\beta\beta)_{0\nu} \sim \langle m \rangle M(A,Z), \quad M(A,Z) - \text{NME},$$

$$\begin{aligned} |\langle m \rangle| &= |m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 e^{i\alpha_{21}} + m_3 |U_{e3}|^2 e^{i\alpha'_{31}}| \\ &= |m_1 c_{12}^2 c_{13}^2 + m_2 s_{12}^2 c_{13}^2 e^{i\alpha_{21}} + m_3 s_{13}^2 e^{i\alpha'_{31}}|, \quad \theta_{12} \equiv \theta_{\odot}, \theta_{13} - \text{CHOOZ} \end{aligned}$$

$\alpha_{21}, \alpha_{31}$  - the two Majorana CPVP of the PMNS matrix;  $\alpha'_{31} \equiv \alpha_{31} - 2\delta$

**CP-invariance:**  $\alpha_{21} = 0, \pm\pi, \alpha_{31} = 0, \pm\pi;$

$$\eta_{21} \equiv e^{i\alpha_{21}} = \pm 1, \quad \eta_{31} \equiv e^{i\alpha_{31}} = \pm 1$$

**relative CP-parities of  $\nu_1$  and  $\nu_2$ , and of  $\nu_1$  and  $\nu_3$  .**

L. Wolfenstein, 1981;

S.M. Bilenky, N. Nedelcheva, S.T.P., 1984;

B. Kayser, 1984.

$$|\langle m \rangle| : m_j, \theta_\odot \equiv \theta_{12}, \theta_{13}, \alpha_{21,31}$$

$m_{1,2,3}$  - in terms of  $\min(m_j)$ ,  $\Delta m_{\text{atm}}^2$ ,  $\Delta m_\odot^2$

S.T.P., A.Yu. Smirnov, 1994

Convention:  $m_1 < m_2 < m_3$  - **NMO**,  $m_3 < m_1 < m_2$  - **IMO**

$$\Delta m_\odot^2 \equiv \Delta m_{21}^2, \quad m_2 = \sqrt{m_1^2 + \Delta m_\odot^2},$$

while either

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2 > 0, \quad m_3 = \sqrt{m_1^2 + \Delta m_{\text{atm}}^2}, \quad \text{normal mass ordering, or}$$

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{32}^2 < 0, \quad m_1 = \sqrt{m_3^2 + |\Delta m_{\text{atm}}^2| - \Delta m_\odot^2}, \quad \text{inverted mass ordering}$$

The neutrino mass spectrum –

*Normal hierarchical (NH)* if  $m_1 \ll m_2 \ll m_3$ ,

*Inverted hierarchical (IH)* if  $m_3 \ll m_1 \cong m_2$ ,

*Quasi-degenerate (QD)* if  $m_1 \cong m_2 \cong m_3 = m$ ,  $m_j^2 \gg |\Delta m_{\text{atm}}^2|$ ;  $m_j \gtrsim 0.1$  eV

Given  $|\Delta m_{\text{atm}}^2|$ ,  $\Delta m_\odot^2$ ,  $\theta_\odot$ ,  $\theta_{13}$ ,

$$|\langle m \rangle| = |\langle m \rangle| (m_{\text{min}}, \alpha_{21}, \alpha_{31}; S), \quad S = \text{NO(NH), IO(IH)}.$$

$$A(\beta\beta)_{0\nu} \sim \langle m \rangle M(A,Z), \quad M(A,Z) - \text{NME},$$

$$|\langle m \rangle| \cong \left| \sqrt{\Delta m_{\odot}^2} \sin^2 \theta_{12} e^{i\alpha} + \sqrt{\Delta m_{31}^2} \sin^2 \theta_{13} e^{i\beta} \right|, \quad m_1 \ll m_2 \ll m_3 \text{ (NH)},$$

$$|\langle m \rangle| \cong \sqrt{m_3^2 + \Delta m_{13}^2} |\cos^2 \theta_{12} + e^{i\alpha} \sin^2 \theta_{12}|, \quad m_3 < (\ll) m_1 < m_2 \text{ (IH)},$$

$$|\langle m \rangle| \cong m |\cos^2 \theta_{12} + e^{i\alpha} \sin^2 \theta_{12}|, \quad m_{1,2,3} \cong m \gtrsim 0.10 \text{ eV (QD)},$$

$$\theta_{12} \equiv \theta_{\odot}, \theta_{13} \text{-CHOOZ}; \quad \alpha \equiv \alpha_{21}, \beta + 2\delta \equiv \alpha_{31}.$$

**CP-invariance:**  $\alpha = 0, \pm\pi, \beta_M = 0, \pm\pi;$

$$|\langle m \rangle| \lesssim 5 \times 10^{-3} \text{ eV, NH};$$

$$\sqrt{\Delta m_{13}^2} \cos 2\theta_{12} \cong 0.013 \text{ eV} \lesssim |\langle m \rangle| \lesssim \sqrt{\Delta m_{13}^2} \cong 0.055 \text{ eV, IH};$$

$$m \cos 2\theta_{12} \lesssim |\langle m \rangle| \lesssim m, \quad m \gtrsim 0.10 \text{ eV, QD}.$$

## Solar neutrino and KamLAND data:

$\cos 2\theta_{\odot} = 0.0$  excluded at  $> 6$  s.d.

Best fit value:  $\cos 2\theta_{\odot} \simeq 0.40$

$\cos 2\theta_{\odot} \gtrsim 0.28$ , 95% C.L.

Normal hierarchical spectrum:

$$(|\langle m \rangle|)_{\max} \lesssim 0.005 \text{ eV}$$

Inverted hierarchical spectrum:

$$(|\langle m \rangle|)_{\min} \simeq \sqrt{|\Delta m_{\text{atm}}^2|} \cos 2\theta_{\odot} \cos^2 \theta_{13} \gtrsim 0.01 \text{ eV}$$

$$(|\langle m \rangle|)_{\max} \simeq \sqrt{|\Delta m_{\text{atm}}^2|} \cos^2 \theta_{13} \lesssim 0.055 \text{ eV}$$

Quasi-degenerate spectrum:

$$(|\langle m \rangle|)_{\min} \simeq m (\cos 2\theta_{\odot} \cos^2 \theta_{13} - \sin^2 \theta_{13}) \gtrsim 0.03 \text{ eV}$$

Best sensitivity: Heidelberg-Moscow  $^{76}\text{Ge}$  experiment.

Claim for a positive signal at  $> 3\sigma$ :

H. Klapdor-Kleingrothaus et al., PL B586 (2004),

$|\langle m \rangle| = (0.1 - 0.9) \text{ eV}$  (99.73% C.L.).

IGEX  $^{76}\text{Ge}$ :  $|\langle m \rangle| < (0.33 - 1.35) \text{ eV}$  (90% C.L.).

Taking data - NEMO3 ( $^{100}\text{Mo}$ ), CUORICINO ( $^{130}\text{Te}$ ):

$|\langle m \rangle| < (0.7 - 1.2) \text{ eV}$ ,  $|\langle m \rangle| < (0.18 - 0.90) \text{ eV}$  (90% C.L.).

Large number of projects:  $|\langle m \rangle| \sim (0.01 - 0.05) \text{ eV}$

CUORE -  $^{130}\text{Te}$ ;

GERDA -  $^{76}\text{Ge}$ ;

SuperNEMO -  $^{82}\text{Se}, \dots$ ;

COBRA -  $^{116}\text{Cd}$ ;

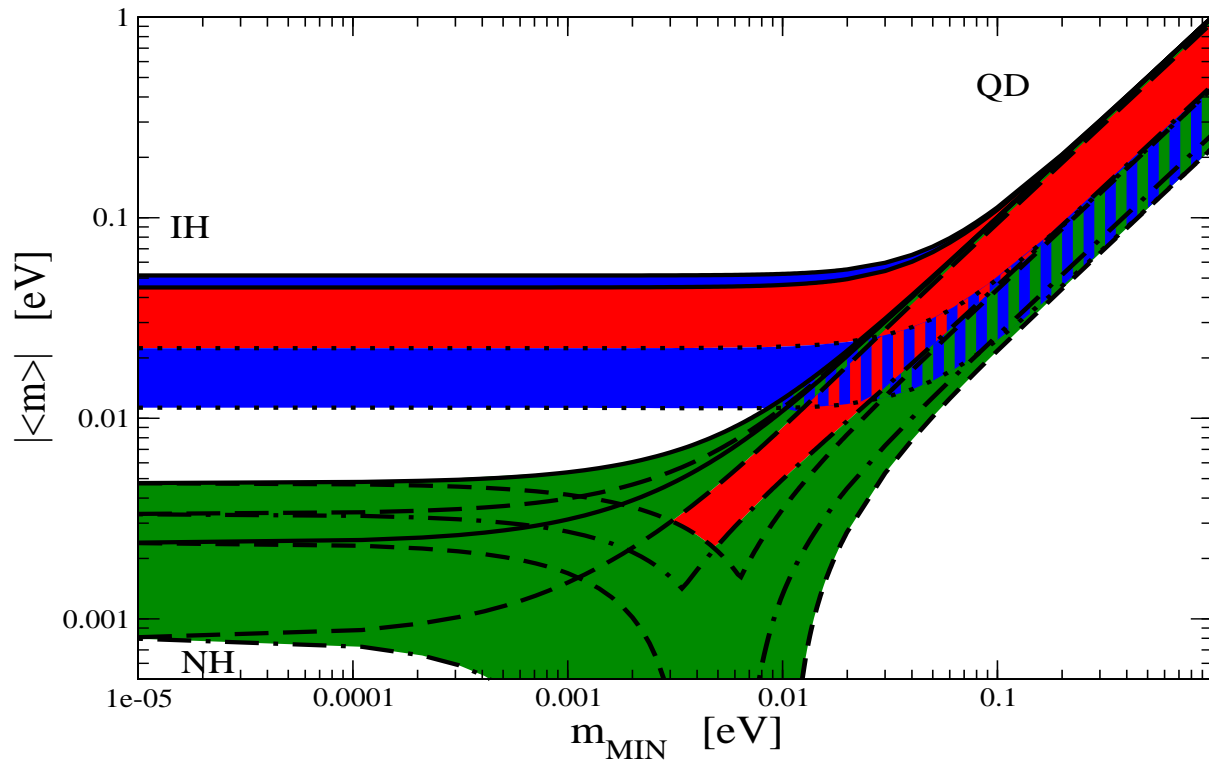
EXO -  $^{136}\text{Xe}$ ;

MAJORANA -  $^{76}\text{Ge}$ ;

MOON -  $^{100}\text{Mo}$ ;

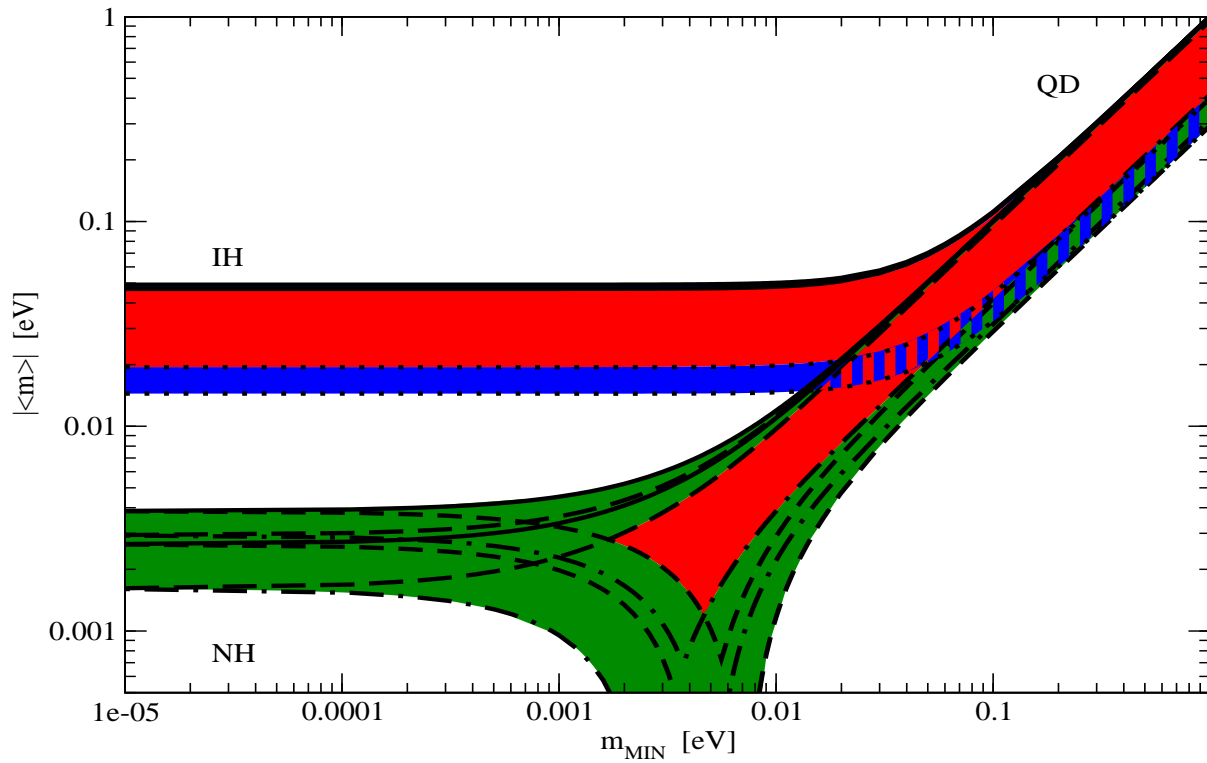
CANDLES -  $^{48}\text{Ca}$ ;

XMASS -  $^{136}\text{Xe}$ .



S. Pascoli, S.T.P., 2007

The current  $2\sigma$  ranges of values of the parameters used.



$\sin^2 \theta_{13} = 0.01 \pm 0.006$ ;  $1\sigma(\Delta m_{\odot}^2) = 2\%$ ,  $1\sigma(\sin^2 \theta_{\odot}) = 4\%$ ,  $1\sigma(|\Delta m_{\text{atm}}^2|) = 2\%$ ;

$2\sigma(|\langle m \rangle|)$  used.

## Majorana CPV Phases and $|\langle m \rangle|$

CPV can be established provided

- $|\langle m \rangle|$  measured with  $\Delta \lesssim 15\%$  ;
- $\Delta m_{\text{atm}}^2$  (IH) or  $m_0$  (QD) measured with  $\delta \lesssim 10\%$  ;
- $\xi \lesssim 1.5$  ;
- $\alpha_{21}$  (QD): in the interval  $\sim [\frac{\pi}{4} - \frac{3\pi}{4}]$ , or  $\sim [\frac{5\pi}{4} - \frac{3\pi}{2}]$  ;
- $\tan^2 \theta_{\odot} \gtrsim 0.40$  .

S. Pascoli, S.T.P., W. Rodejohann, 2002

S. Pascoli, S.T.P., L. Wolfenstein, 2002

S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226

No “No-go for detecting CP-Violation via  $(\beta\beta)_{0\nu}$ -decay”

V. Barger *et al.*, 2002



# Absolute Neutrino Mass Measurements

The Troitzk and Mainz  ${}^3\text{H}$   $\beta$ -decay experiments

$$m_{\nu_e} < 2.3 \text{ eV} \quad (95\% \text{ C.L.})$$

There are prospects to reach sensitivity

$$\text{KATRIN :} \quad m_{\nu_e} \sim 0.2 \text{ eV}$$

Cosmological and astrophysical data: the WMAP result combined with data from large scale structure surveys (2dFGRS, SDSS)

$$\sum_j m_j \equiv \Sigma < (0.4 - 1.7) \text{ eV}$$

The WMAP and future PLANCK experiments can be sensitive to

$$\sum_j m_j \cong 0.4 \text{ eV}$$

Data on weak lensing of galaxies by large scale structure, combined with data from the WMAP and PLANCK experiments may allow to determine

$$\sum_j m_j : \quad \delta \cong 0.04 \text{ eV.}$$

# $M_\nu$ from the See-Saw Mechanism

P. Minkowski, 1977.

M. Gell-Mann, P. Ramond, R. Slansky, 1979;

T. Yanagida, 1979;

R. Mohapatra, G. Senjanovic, 1980.

- Explains the smallness of  $\nu$ -masses.
- Through **leptogenesis theory** links the  $\nu$ -mass generation to the generation of baryon asymmetry of the Universe  $Y_B$ .

S. Fukugita, T. Yanagida, 1986.

- In SUSY GUT's with see-saw mechanism of  $\nu$ -mass generation, the LFV decays

$$\mu \rightarrow e + \gamma, \quad \tau \rightarrow \mu + \gamma, \quad \tau \rightarrow e + \gamma, \quad \text{etc.}$$

are predicted to take place with rates within the reach of present and future experiments.

F. Borzumati, A. Masiero, 1986.

- The  $\nu_j$  are **Majorana particles**;  $(\beta\beta)_{0\nu}$ -decay is allowed.

**See-Saw:** Dirac  $\nu$ -mass  $m_D$  + Majorana mass  $M_R$  for  $N_R$

# The See-Saw Lagrangian

$$\mathcal{L}^{\text{lep}}(x) = \mathcal{L}_{\text{CC}}(x) + \mathcal{L}_Y(x) + \mathcal{L}_M^{\text{N}}(x),$$

$$\mathcal{L}_{\text{CC}} = -\frac{g}{\sqrt{2}} \bar{l}_L(x) \gamma_\alpha \nu_{lL}(x) W^{\alpha\dagger}(x) + h.c.,$$

$$\mathcal{L}_Y(x) = \lambda_{il} \bar{N}_{iR}(x) H^\dagger(x) \psi_{lL}(x) + Y_l H^c(x) \bar{l}_R(x) \psi_{lL}(x) + h.c.,$$

$$\mathcal{L}_M^{\text{N}}(x) = -\frac{1}{2} M_i \bar{N}_i(x) N_i(x).$$

$\psi_{lL}$  - LH doublet,  $\psi_{lL}^T = (\nu_{lL} \ l_L)$ ,  $l_R$  - RH singlet,  $H$  - Higgs doublet.

Basis:  $M_R = (M_1, M_2, M_3)$ ;  $D_N \equiv \text{diag}(M_1, M_2, M_3)$ ,  $D_\nu \equiv \text{diag}(m_1, m_2, m_3)$ .

$m_D$  generated by the Yukawa interaction:

$$-\mathcal{L}_Y^\nu = \lambda_{il} \bar{N}_{iR} H^\dagger(x) \psi_{lL}(x), \quad v = 174 \text{ GeV}, \quad v \lambda = m_D - \text{complex}$$

For  $M_R$  - sufficiently large,

$$m_\nu \simeq v^2 \lambda^T M_R^{-1} \lambda = U_{\text{PMNS}}^* m_\nu^{\text{diag}} U_{\text{PMNS}}^\dagger.$$

$Y_\nu \equiv \lambda = \sqrt{D_N} R \sqrt{D_\nu} (U_{\text{PMNS}})^\dagger / v_u$ , all at  $M_R$ ;  $R$ -complex,  $R^T R = 1$ .

J.A. Casas and A. Ibarra, 2001

In GUTs,  $M_R < M_X$ ,  $M_X \sim 10^{16}$  GeV;

in GUTs, e.g.,  $M_R = (10^9, 10^{12}, 10^{15})$  GeV,  $m_D \sim 1$  GeV.

## The CP-Invariance Constraints

Assume:  $C(\bar{\nu}_j)^T = \nu_j$ ,  $C(\bar{N}_k)^T = N_k$ ,  $j, k = 1, 2, 3$ .

The CP-symmetry transformation:

$$\begin{aligned} U_{\text{CP}} N_j(x) U_{\text{CP}}^\dagger &= \eta_j^{\text{NCP}} \gamma_0 N_j(x'), \quad \eta_j^{\text{NCP}} = i\rho_j^N = \pm i, \\ U_{\text{CP}} \nu_k(x) U_{\text{CP}}^\dagger &= \eta_k^{\nu\text{CP}} \gamma_0 \nu_k(x'), \quad \eta_k^{\nu\text{CP}} = i\rho_k^\nu = \pm i. \end{aligned}$$

CP-invariance:

$$\lambda_{jl}^* = \lambda_{jl} (\eta_j^{\text{NCP}})^* \eta^l \eta^{H*}, \quad j = 1, 2, 3, \quad l = e, \mu, \tau,$$

Convenient choice:  $\eta^l = i$ ,  $\eta^H = 1$  ( $\eta^W = 1$ ):

$$\lambda_{jl}^* = \lambda_{jl} \rho_j^N, \quad \rho_j^N = \pm 1,$$

$$U_{lj}^* = U_{lj} \rho_j^\nu, \quad \rho_j^\nu = \pm 1,$$

$$R_{jk}^* = R_{jk} \rho_j^N \rho_k^\nu, \quad j, k = 1, 2, 3, \quad l = e, \mu, \tau,$$

$\lambda_{jl}$ ,  $U_{lj}$ ,  $R_{jk}$  - either real or purely imaginary.

Relevant quantity:

$$P_{jkml} \equiv R_{jk} R_{jm} U_{lk}^* U_{lm}, \quad k \neq m,$$

$$\text{CP: } P_{jkml}^* = P_{jkml} (\rho_j^N)^2 (\rho_k^\nu)^2 (\rho_m^\nu)^2 = P_{jkml}, \quad \text{Im}(P_{jkml}) = 0.$$

$$P_{jkml} \equiv R_{jk} R_{jm} U_{lk}^* U_{lm}, \quad k \neq m,$$

$$CP: \quad P_{jkml}^* = P_{jkml} (\rho_j^N)^2 (\rho_k^\nu)^2 (\rho_m^\nu)^2 = P_{jkml}, \quad \text{Im}(P_{jkml}) = 0.$$

Consider NH  $N_j$ , NH  $\nu_k$ :  $P_{123\tau} = R_{12} R_{13} U_{\tau 2}^* U_{\tau 3}$

Suppose, CP-invariance holds at low  $E$ :  $\delta = 0$ ,  $\alpha_{21} = \pi$ ,  $\alpha_{31} = 0$ .

Thus,  $U_{\tau 2}^* U_{\tau 3}$  - purely imaginary.

Then real  $R_{12} R_{13}$  corresponds to CP-violation at "high"  $E$ .

# Leptogenesis

$$Y_B = \frac{n_B - n_{\bar{B}}}{S} \sim 8.6 \times 10^{-11} \quad (n_\gamma: \sim 6.3 \times 10^{-10})$$

$$Y_B \cong -10^{-2} \quad \varepsilon \kappa$$

W. Buchmüller, M. Plümacher, 1998;

W. Buchmüller, P. Di Bari, M. Plümacher, 2004

$\kappa$ - efficiency factor;  $\kappa \sim 10^{-1} - 10^{-3}$ :  $\varepsilon \gtrsim 10^{-7}$ .

$\varepsilon$ :  $CP$ -,  $L$ - violating asymmetry generated in out of equilibrium  $N_{Rj}$ -decays in the early Universe,

$$\varepsilon_1 = \frac{\Gamma(N_1 \rightarrow \Phi^- \ell^+) - \Gamma(N_1 \rightarrow \Phi^+ \ell^-)}{\Gamma(N_1 \rightarrow \Phi^- \ell^+) + \Gamma(N_1 \rightarrow \Phi^+ \ell^-)}$$

M.A. Luty, 1992;

L. Covi, E. Roulet and F. Vissani, 1996;

M. Flanz *et al.*, 1996;

M. Plümacher, 1997;

A. Pilaftsis, 1997.

$\kappa = \kappa(\tilde{m})$ ,  $\tilde{m}$  - determines the rate of wash-out processes:

$\Phi^+ + \ell^- \rightarrow N_1$ ,  $\ell^- + \Phi^+ \rightarrow \Phi^- + \ell^+$ , etc.

W. Buchmüller, P. Di Bari and M. Plümacher, 2002;

G. F. Giudice *et al.*, 2004

# Low Energy Leptonic CPV and Leptogenesis

Assume:  $M_1 \ll M_2 \ll M_3$

Individual asymmetries:

$$\varepsilon_{1l} = -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left( \sum_{j,k} m_j^{1/2} m_k^{3/2} U_{lj}^* U_{lk} R_{1j} R_{1k} \right)}{\sum_j m_j |R_{1j}|^2}, \quad v = 174 \text{ GeV}$$

$$\widetilde{m}_l \equiv \frac{|\lambda_{1l}|^2 v^2}{M_1} = \left| \sum_k R_{1k} m_k^{1/2} U_{lk}^* \right|^2, \quad l = e, \mu, \tau.$$

The “One-Flavor” Regime:  $M_1 \sim T > 10^{12}$  GeV;  $Y_{e,\mu,\tau}$  - “small”

Boltzmann eqn. for  $n(N_1)$  and  $\Delta L = \Delta(L_e + L_\mu + L_\tau)$ .

$Y_l H^c(x) \bar{l}_R(x) \psi_{lL}$ - out of equilibrium at  $T \sim M_1$ .

One-flavor approximation:  $M_1 \sim T > 10^{12}$  GeV

$$\varepsilon_1 = \sum_l \varepsilon_{1l} = -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left( \sum_{j,k} m_j^2 R_{1j}^2 \right)}{\sum_k m_k |R_{1k}|^2},$$

$$\widetilde{m}_1 = \sum_l \widetilde{m}_l = \sum_k m_k |R_{1k}|^2.$$

Two-Flavour Regime:  $10^9 \text{ GeV} \lesssim M_1 \sim T \lesssim 10^{12} \text{ GeV}$

At  $M_1 \sim T \lesssim 10^{12} \text{ GeV}$ :  $Y_\tau$  - in equilibrium,  $Y_{e,\mu}$  - not;

wash-out dynamics changes:  $\tau_R^-, \tau_L^+$

$N_1 \rightarrow (\lambda_{1e} e_L^- + \lambda_{1\mu} \mu_L^- + \lambda_{1\tau} \tau_L^-) + \Phi^+$ ;  $(\lambda_{1e} e_L^- + \lambda_{1\mu} \mu_L^- + \lambda_{1\tau} \tau_L^-) + \Phi^+ \rightarrow N_1$ ;

$\tau_L^- + \Phi^0 \rightarrow \tau_R^-$ ,  $\tau_L^- + \tau_L^+ \rightarrow N_1 + \nu_L$ , etc.

$\varepsilon_{1\tau}$  and  $(\varepsilon_{1e} + \varepsilon_{1\mu}) \equiv \varepsilon_2$  evolve independently.

Thus, at  $M_1 \sim 10^9 - 10^{12} \text{ GeV}$ :  $L_\tau, \Delta L_\tau$  - distinguishable;

$L_e + L_\mu, \Delta(L_e + L_\mu)$  - distinguishable;

$L_e, L_\mu, \Delta L_e, \Delta L_\mu$  - individually not distinguishable.

Three-Flavour Regime:  $M_1 \sim T < 10^9 \text{ GeV}$

At  $M_1 \sim T \sim 10^9 \text{ GeV}$ :  $Y_\tau, Y_\mu$  - in equilibrium,  $Y_e$  - not.

$\varepsilon_{1\tau}, \varepsilon_{1e}$  and  $\varepsilon_{1\mu}$  evolve independently.

A. Abada et al., 2006; E. Nardi et al., 2006

A. Abada et al., 2006



## Individual asymmetries:

Assume:  $M_1 \ll M_2 \ll M_3$ ,  $10^9 \lesssim M_1 (\sim T) \lesssim 10^{12}$  GeV,

$$\epsilon_{1l} = -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left( \sum_{j,k} m_j^{1/2} m_k^{3/2} U_{lj}^* U_{lk} R_{1j} R_{1k} \right)}{\sum_j m_j |R_{1j}|^2}$$

$$\widetilde{m}_l \equiv \frac{|\lambda_{1l}|^2 v^2}{M_1} = \left| \sum_k R_{1k} m_k^{1/2} U_{lk}^* \right|^2, \quad l = e, \mu, \tau.$$

The baryon asymmetry is

$$Y_B \simeq -\frac{12}{37g_*} \left( \epsilon_2 \eta \left( \frac{417}{589} \widetilde{m}_2 \right) + \epsilon_\tau \eta \left( \frac{390}{589} \widetilde{m}_\tau \right) \right),$$

$$\eta(\widetilde{m}_l) \simeq \left( \left( \frac{\widetilde{m}_l}{8.25 \times 10^{-3} \text{ eV}} \right)^{-1} + \left( \frac{0.2 \times 10^{-3} \text{ eV}}{\widetilde{m}_l} \right)^{-1.16} \right)^{-1}.$$

$$Y_B = -(12/37) (Y_2 + Y_\tau),$$

$$Y_2 = Y_{e+\mu}, \quad \epsilon_2 = \epsilon_{1e} + \epsilon_{1\mu}, \quad \widetilde{m}_2 = \widetilde{m}_{1e} + \widetilde{m}_{1\mu}$$

A. Abada et al., 2006; E. Nardi et al., 2006

A. Abada et al., 2006

Real (Purely Imaginary)  $R$ :  $\varepsilon_{1l} \neq 0$ , CPV from  $U$

$$\varepsilon_{1e} + \varepsilon_{1\mu} + \varepsilon_{1\tau} = \varepsilon_2 + \varepsilon_{1\tau} = 0,$$

$$\begin{aligned} \varepsilon_{1\tau} &= -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left( \sum_{j,k} m_j^{1/2} m_k^{3/2} U_{\tau j}^* U_{\tau k} R_{1j} R_{1k} \right)}{\sum_j m_j |R_{1j}|^2} \\ &= -\frac{3M_1}{16\pi v^2} \frac{\sum_{j,k>j} m_j^{1/2} m_k^{1/2} (m_k - m_j) R_{1j} R_{1k} \text{Im} (U_{\tau j}^* U_{\tau k})}{\sum_j m_j |R_{1j}|^2}, R_{1j} R_{1k} = \pm |R_{1j} R_{1k}|, \\ &= \mp \frac{3M_1}{16\pi v^2} \frac{\sum_{j,k>j} m_j^{1/2} m_k^{1/2} (m_k + m_j) |R_{1j} R_{1k}| \text{Re} (U_{\tau j}^* U_{\tau k})}{\sum_j m_j |R_{1j}|^2}, R_{1j} R_{1k} = \pm i |R_{1j} R_{1k}| \end{aligned}$$

S. Pascoli, S.T.P., A. Riotto, 2006.

CP-Violation:  $\text{Im} (U_{\tau j}^* U_{\tau k}) \neq 0$ ,  $\text{Re} (U_{\tau j}^* U_{\tau k}) \neq 0$ ;

$$Y_B = -\frac{12}{37} \frac{\varepsilon_{1\tau}}{g_*} \left( \eta \left( \frac{390}{589} \widetilde{m}_\tau \right) - \eta \left( \frac{417}{589} \widetilde{m}_2 \right) \right)$$

$m_1 \ll m_2 \ll m_3$ ,  $M_1 \ll M_{2,3}$ ;  $R_{12}R_{13}$  – real;  $m_1 \cong 0$ ,  $R_{11} \cong 0$  ( $N_3$  decoupling)

$$\varepsilon_{1\tau} = - \frac{3M_1 \sqrt{\Delta m_{31}^2}}{16\pi v^2} \left( \frac{\Delta m_{\odot}^2}{\Delta m_{31}^2} \right)^{\frac{1}{4}} \frac{|R_{12}R_{13}|}{\left( \frac{\Delta m_{\odot}^2}{\Delta m_{31}^2} \right)^{\frac{1}{2}} |R_{12}|^2 + |R_{13}|^2} \\ \times \left( 1 - \frac{\sqrt{\Delta m_{\odot}^2}}{\sqrt{\Delta m_{31}^2}} \right) \text{Im} (U_{\tau 2}^* U_{\tau 3})$$

$$\text{Im} (U_{\tau 2}^* U_{\tau 3}) = -c_{13} \left[ c_{23}s_{23}c_{12} \sin \left( \frac{\alpha_{32}}{2} \right) - c_{23}^2 s_{12}s_{13} \sin \left( \delta - \frac{\alpha_{32}}{2} \right) \right]$$

$\alpha_{32} = \pi$ ,  $\delta = 0$ :  $\text{Re}(U_{\tau 2}^* U_{\tau 3}) = 0$ , CPV due to  $R$

S. Pascoli, S.T.P., A. Riotto, 2006.

$$M_1 \ll M_2 \ll M_3, m_1 \ll m_2 \ll m_3 \text{ (NH)}$$

## Dirac CP-violation

$$\alpha_{32} = 0 \text{ (} 2\pi \text{)}, \beta_{23} = \pi \text{ (} 0 \text{)}; \quad \beta_{23} \equiv \beta_{12} + \beta_{13} \equiv \arg(R_{12}R_{13}).$$

$$|R_{12}|^2 \cong 0.85, \quad |R_{13}|^2 = 1 - |R_{12}|^2 \cong 0.15 - \text{maximise } |\epsilon_\tau| \text{ and } |Y_B|:$$

$$|Y_B| \cong 2.8 \times 10^{-13} |\sin \delta| \left( \frac{s_{13}}{0.2} \right) \left( \frac{M_1}{10^9 \text{ GeV}} \right).$$

$$|Y_B| \gtrsim 8 \times 10^{-11}, \quad M_1 \lesssim 5 \times 10^{11} \text{ GeV imply}$$

$$|\sin \theta_{13} \sin \delta| \gtrsim 0.11, \quad \sin \theta_{13} \gtrsim 0.11.$$

The lower limit corresponds to

$$|J_{\text{CP}}| \gtrsim 2.4 \times 10^{-2}$$

FOR  $\alpha_{32} = 0 \text{ (} 2\pi \text{)}, \beta_{23} = 0 \text{ (} \pi \text{)}$ :

$$|\sin \theta_{13} \sin \delta| \gtrsim 0.09, \quad \sin \theta_{13} \gtrsim 0.09; \quad |J_{\text{CP}}| \gtrsim 2.0 \times 10^{-2}$$

$$M_1 \ll M_2 \ll M_3, m_1 \ll m_2 \ll m_3 \text{ (NH)}$$

Majorana CP-violation

$$\delta = 0, \text{ real } R_{12}, R_{13} (\beta_{23} = \pi (0));$$

$$\alpha_{32} \cong \pi/2, \quad |R_{12}|^2 \cong 0.85, \quad |R_{13}|^2 = 1 - |R_{12}|^2 \cong 0.15 - \text{maximise } |\epsilon_\tau| \text{ and } |Y_B|:$$

$$|Y_B| \cong 2 \times 10^{-12} \left( \frac{\sqrt{\Delta m_{31}^2}}{0.05 \text{ eV}} \right) \left( \frac{M_1}{10^9 \text{ GeV}} \right).$$

We get  $|Y_B| \gtrsim 8 \times 10^{-11}$ , for  $M_1 \gtrsim 3.6 \times 10^{10} \text{ GeV}$ , or  $|\sin \alpha_{32}/2| \gtrsim 0.15$

$$M_1 \ll M_2 \ll M_3, m_3 \ll m_1 < m_2 \text{ (IH)}$$

$m_3 \cong 0, R_{13} \cong 0$  ( $N_3$  decoupling): impossible to reproduce  $Y_B^{obs}$  for real  $R_{11}R_{12}$ ;

$|Y_B|$  suppressed by the additional factor  $\Delta m_{\odot}^2/|\Delta m_{31}^2| \cong 0.03$ .

Purely imaginary  $R_{11}R_{12}$ : no (additional) suppression

Dirac CP-violation

$$\alpha_{21} = \pi; R_{11}R_{12} = i\kappa|R_{11}R_{12}|, \kappa = 1;$$

$|R_{11}| \cong 1.07, |R_{12}|^2 = |R_{11}|^2 - 1, |R_{12}| \cong 0.38$  - maximise  $|\epsilon_\tau|$  and  $|Y_B|$ :

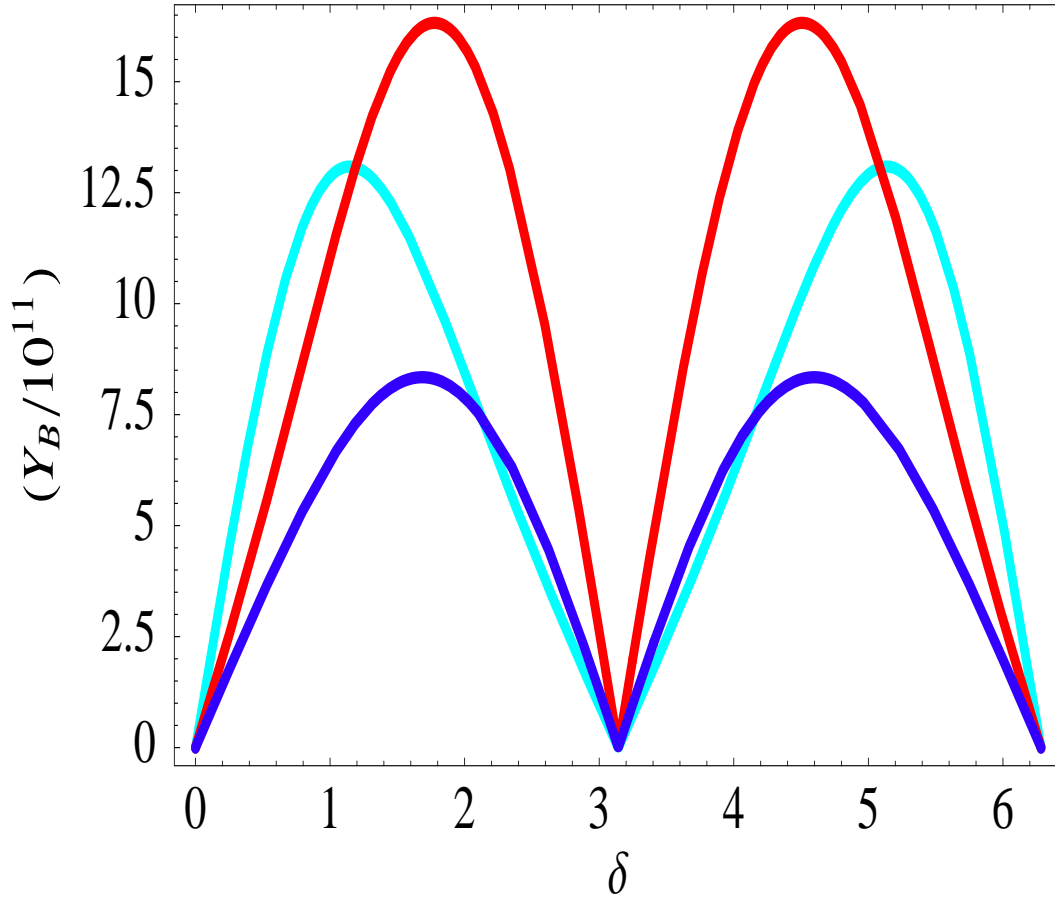
$$|Y_B| \cong 8.1 \times 10^{-12} |s_{13} \sin \delta| \left( \frac{M_1}{10^9 \text{ GeV}} \right).$$

$|Y_B| \gtrsim 8 \times 10^{-11}, M_1 \lesssim 5 \times 10^{11} \text{ GeV}$  imply

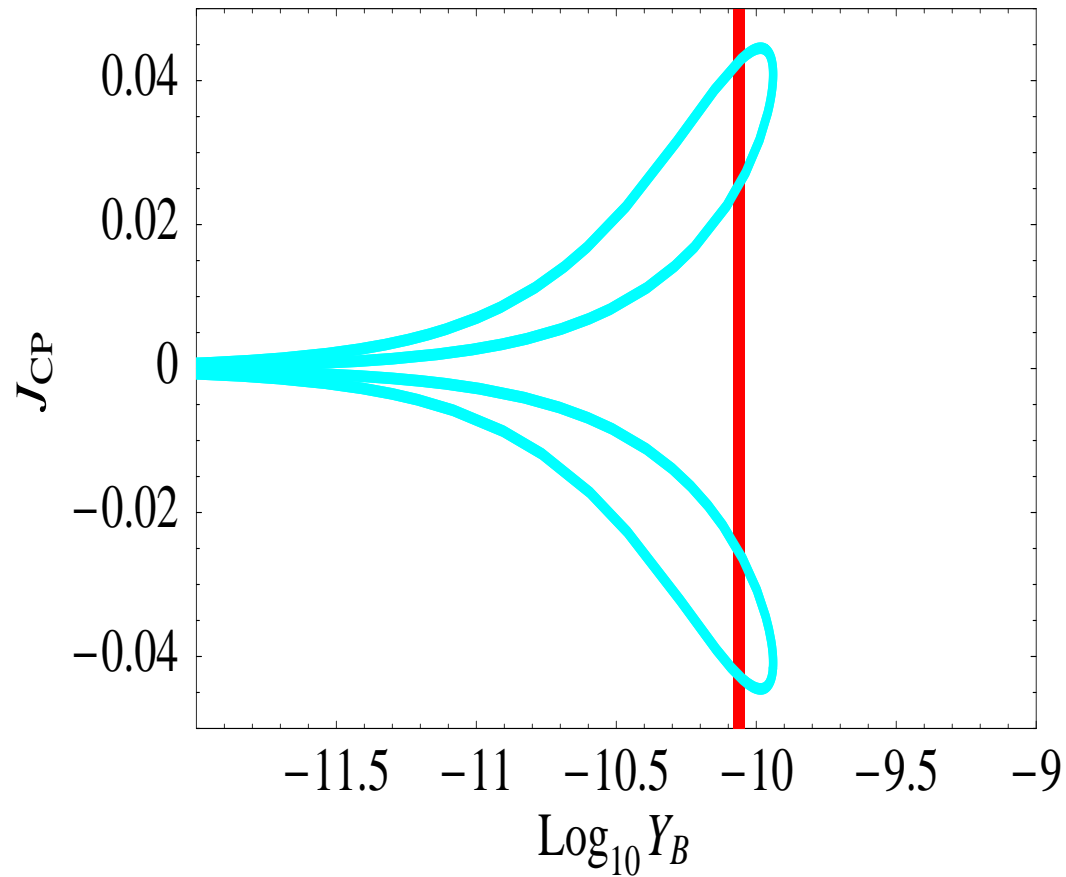
$$|\sin \theta_{13} \sin \delta| \gtrsim 0.02, \quad \sin \theta_{13} \gtrsim 0.02.$$

The lower limit corresponds to

$$|J_{CP}| \gtrsim 4.6 \times 10^{-3}$$



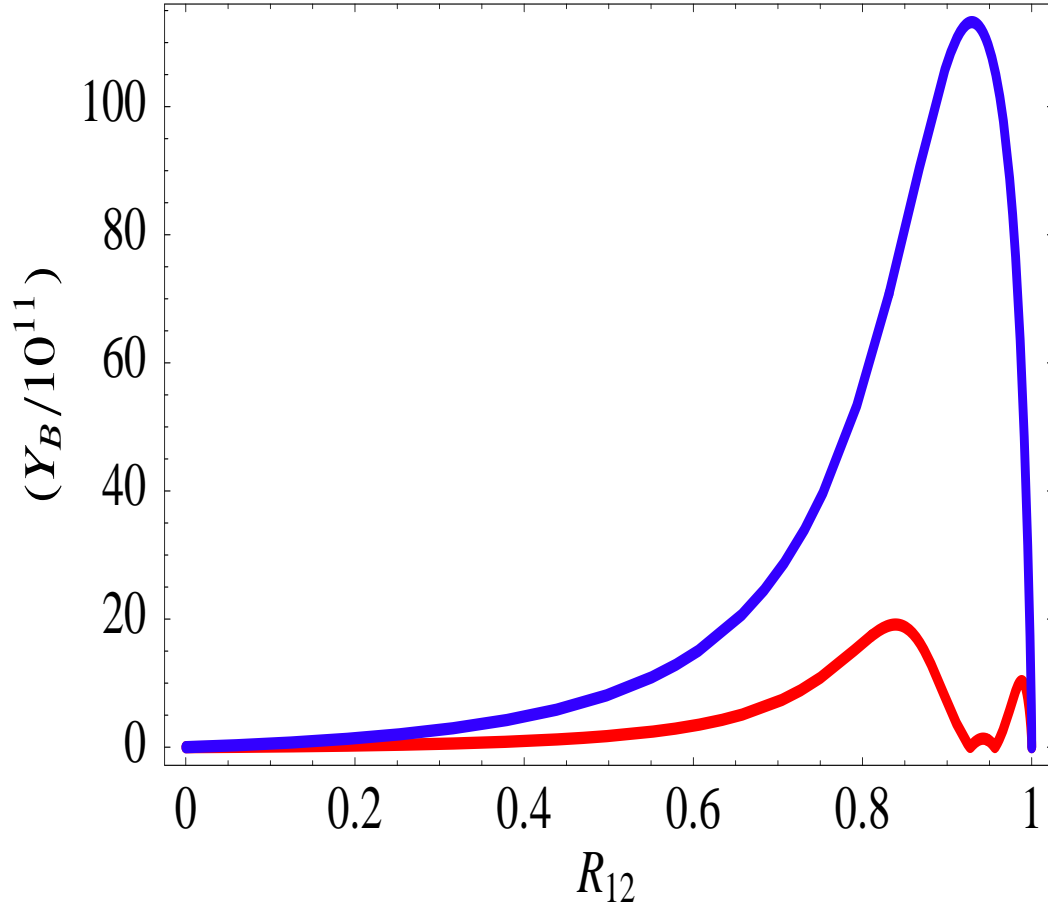
$M_1 \ll M_2 \ll M_3$ ,  $m_1 \ll m_2 \ll m_3$ ; Dirac CP-violation,  $\alpha_{32} = 0$ ;  $2\pi$ ;  
 real  $R_{12}$ ,  $R_{13}$ ,  $|R_{12}|^2 + |R_{13}|^2 = 1$ ,  $|R_{12}| = 0.86$ ,  $|R_{13}| = 0.51$ ,  $\text{sign}(R_{12}R_{13}) = +1$ ;  
 i)  $\alpha_{32} = 0$  ( $\kappa' = +1$ ),  $s_{13} = 0.2$  (red line) and  $s_{13} = 0.1$  (dark blue line);  
 ii)  $\alpha_{32} = 2\pi$  ( $\kappa' = -1$ ),  $s_{13} = 0.2$  (light blue line);  
 $M_1 = 5 \times 10^{11}$  GeV.



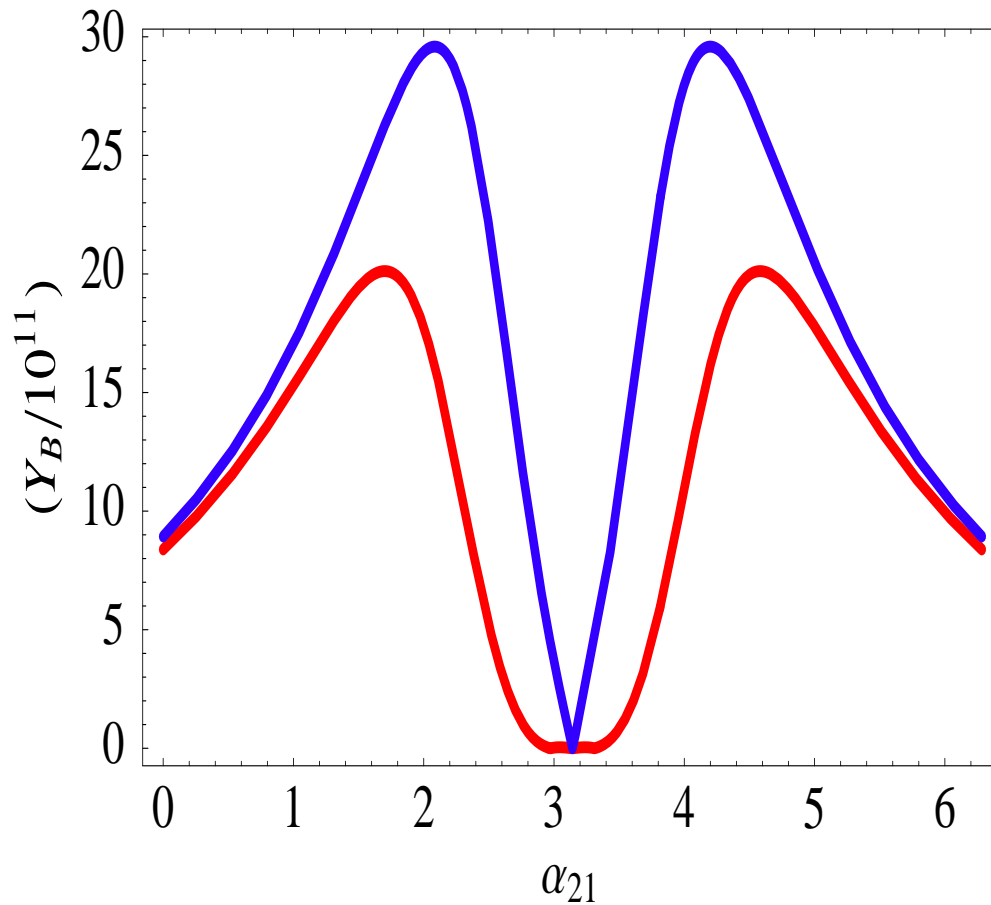
$M_1 \ll M_2 \ll M_3, m_1 \ll m_2 \ll m_3; M_1 = 5 \times 10^{11} \text{ GeV};$   
 Dirac CP-violation,  $\alpha_{32} = 0 \text{ (} 2\pi \text{)}$ ;  
 $|R_{12}| = 0.86, |R_{13}| = 0.51, \text{sign}(R_{12}R_{13}) = +1 \text{ (-1)} \text{ (}\beta_{23} = 0 \text{ (}\pi\text{), } \kappa' = +1\text{)}$ ;  
 The red region denotes the  $2\sigma$  allowed range of  $Y_B$ .

S. Pascoli, S.T.P., A. Riotto, 2006.

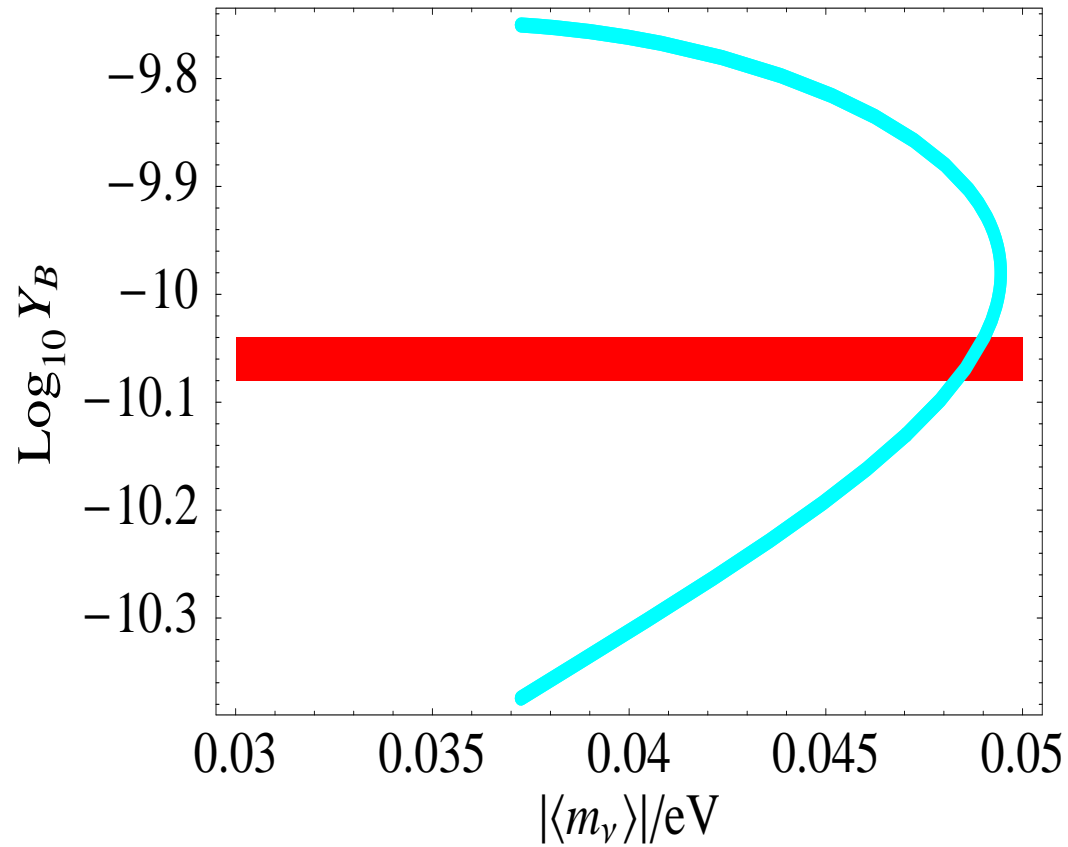




$M_1 \ll M_2 \ll M_3, m_1 \ll m_2 \ll m_3; M_1 = 5 \times 10^{11} \text{ GeV};$   
 real  $R_{12}, R_{13}, \text{sign}(R_{12}R_{13}) = +1, R_{12}^2 + R_{13}^2 = 1, s_{13} = 0.20;$   
 a) Majorana CP-violation (blue line),  $\delta = 0$  and  $\alpha_{32} = \pi/2$  ( $\kappa = +1$ );  
 b) Dirac CP-violation (red line),  $\delta = \pi/2$  and  $\alpha_{32} = 0$  ( $\kappa' = +1$ );  
 $\Delta m_{\odot}^2, \sin^2 \theta_{12}, \Delta m_{31}^2, \sin^2 2\theta_{23}$  - fixed at their best fit values.

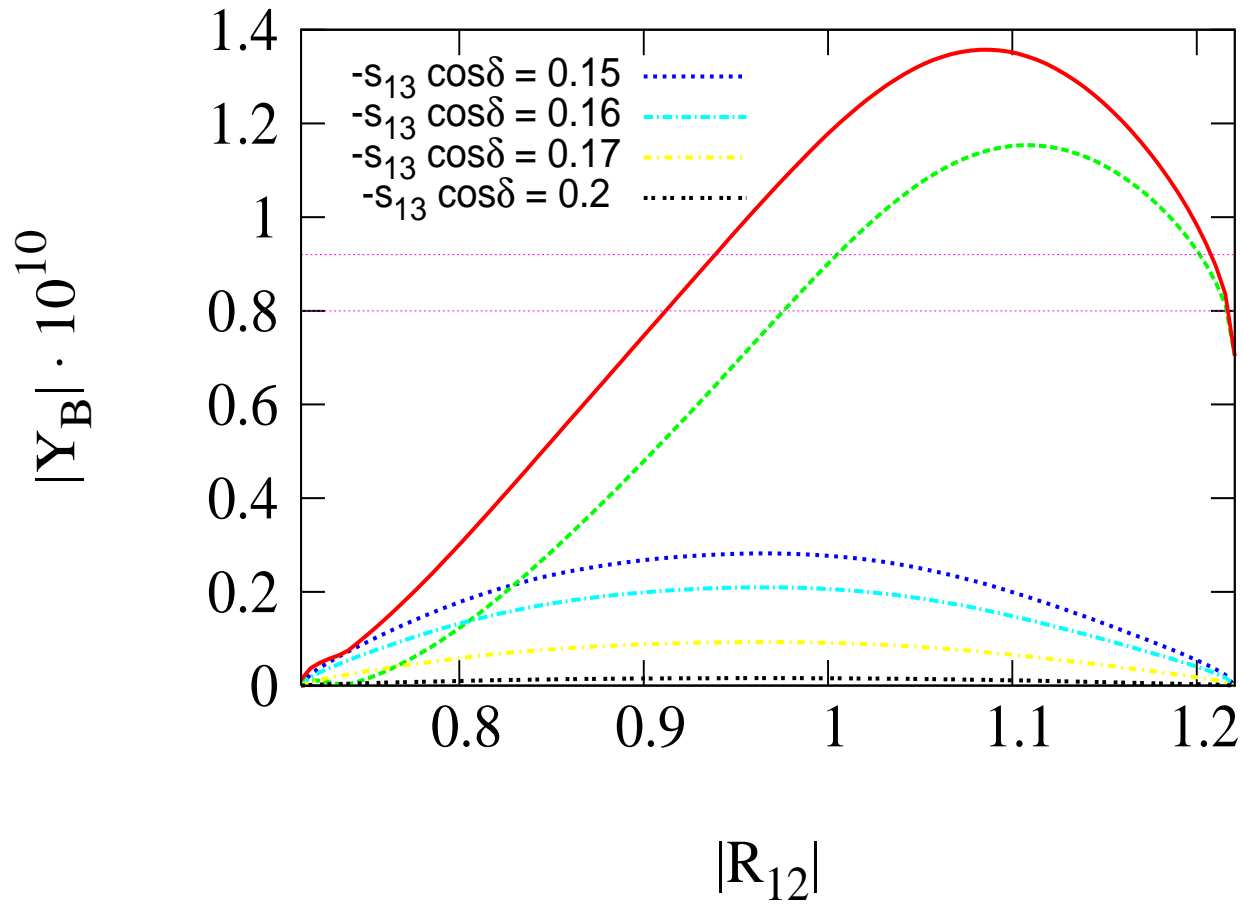


$M_1 \ll M_2 \ll M_3$ ,  $m_3 \ll m_1 < m_2$ ;  $M_1 = 2 \times 10^{11}$  GeV;  
 Majorana CP-violation,  $\delta = 0$ ;  
 purely imaginary  $R_{11}R_{12} = i\kappa|R_{11}R_{12}|$ ,  $\kappa = -1$ ,  $|R_{11}|^2 - |R_{12}|^2 = 1$ ,  $|R_{11}| = 1.2$ ;  
 $s_{13} = 0$  (blue line) and 0.2 (red line).

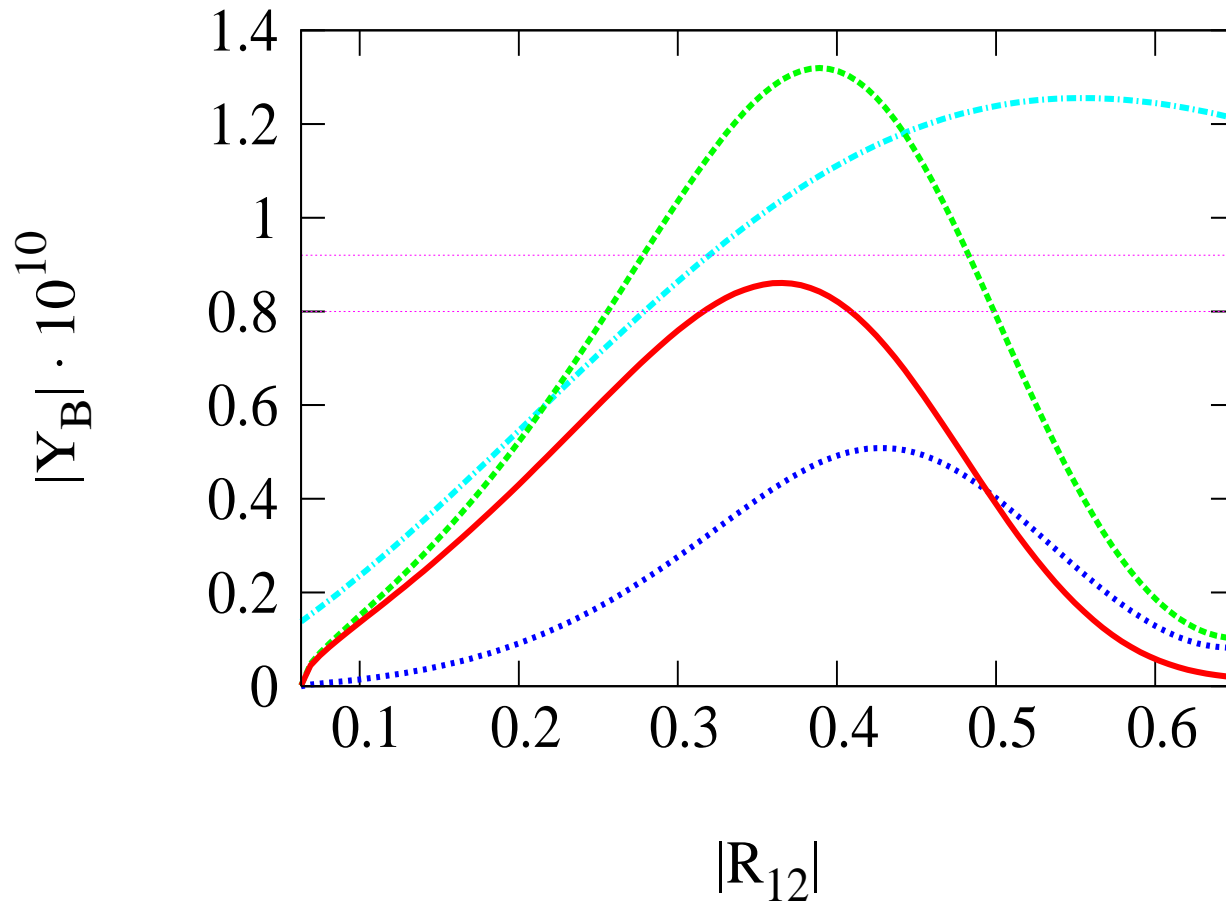


$M_1 \ll M_2 \ll M_3$ ,  $m_3 \ll m_1 < m_2$ ;  $M_1 = 2 \times 10^{11}$  GeV;  
 Majorana CP-violation,  $\delta = 0$ ,  $s_{13} = 0$ ;  
 purely imaginary  $R_{11}R_{12} = i\kappa|R_{11}R_{12}|$ ,  $\kappa = +1$   $|R_{11}|^2 - |R_{12}|^2 = 1$ ,  $|R_{11}| = 1.05$ .  
 The Majorana phase  $\alpha_{21}$  is varied in the interval  $[-\pi/2, \pi/2]$ .

S. Pascoli, S.T.P., A. Riotto, 2006.



$m_3 \ll m_1 < m_2$  (IH),  $R_{13} = 0$ , Majorana and  $R$ -matrix CPV ,  
 $\alpha_{21} = \pi/2$ ,  $(-s_{13} \cos \delta) = 0.15$ ,  $|R_{11}| = 1.2$ ,  $M_1 = 10^{11}$  GeV;  
 $|Y_B^0 A_{HE}|$  ( $R$  CPV, blue),  $|Y_B^0 A_{MIX}|$  ( $U$  CPV, green), total  $|Y_B|$  (red line).



$m_3 \ll m_1 < m_2$  (IH),  $R_{13} = 0$ , Majorana and  $R$ -matrix CPV ,  
 $\alpha_{21} = \pi/2$ ,  $s_{13} = 0$ ,  $|R_{11}| \cong 1$ ,  $M_1 = 10^{11}$  GeV;  
 $|Y_B^0 A_{\text{HE}}|$  ( $R$  CPV, blue),  $|Y_B^0 A_{\text{MIX}}|$  ( $U$  CPV, green), total  $|Y_B|$  (red line) .  
 Light-blue line: CP-conserving  $R$ ,  $R_{11}R_{12} \equiv ik|R_{11}R_{12}|$ ,  $k = -1$   $|R_{11}|^2 - |R_{12}|^2 = 1$ .

# Low Energy Leptonic CPV and Leptogenesis: Summary

Leptogenesis: see-saw mechanism;  $N_j$  - heavy RH  $\nu$ 's;

$N_j, \nu_k$  - Majorana particles

$N_j$ :  $M_1 \ll M_2 \ll M_3$

The observed value of the baryon asymmetry of the Universe can be generated

A. CP-violation due to the Dirac phase  $\delta$  in  $U_{\text{PMNS}}$ , no other sources of CPV (Majorana phases in  $U_{\text{PMNS}}$  equal to 0, etc.); requires  $M_1 \gtrsim 10^{11}$  GeV.

$m_1 \ll m_2 \ll m_3$  (NH):

$$|\sin \theta_{13} \sin \delta| \gtrsim 0.09, \quad \sin \theta_{13} \gtrsim 0.09; \quad |J_{\text{CP}}| \gtrsim 2.0 \times 10^{-2}$$

$m_3 \ll m_1 < m_2$  (IH):

$$|\sin \theta_{13} \sin \delta| \gtrsim 0.02, \quad \sin \theta_{13} \gtrsim 0.02; \quad |J_{\text{CP}}| \gtrsim 4.6 \times 10^{-3}$$

B. CP-violation due to the Majorana phases in  $U_{\text{PMNS}}$ , no other sources of CPV (Dirac phase in  $U_{\text{PMNS}}$  equal to 0, etc.); requires  $M_1 \gtrsim 3.5 \times 10^{10}$  GeV.

C. CP-violation due to both Dirac and Majorana phases in  $U_{\text{PMNS}}$ .

D.  $Y_B$  can depend non-trivially on  $\min(m_j) \sim (10^{-5} - 10^{-2})$  eV.

S. Pascoli, S.T.P., A. Riotto, 2006 (A-C);  
E. Molinaro, S.T.P., T. Shindou, Y. Takahashi, 2007 (D).

## Conclusions

Determining the nature - Dirac or Majorana, of massive neutrinos is of fundamental importance for understanding the origin of neutrino masses.

- Can establish the Majorana nature of  $\nu_j$
- Can provide unique information on the  $\nu$  mass spectrum
- Can provide unique information on the absolute scale of  $\nu$  masses
- Can provide information on the Majorana CPV phases

The see-saw mechanism provides a link between  $\nu$ -mass generation and BAU.

Majorana CPV phases in  $U_{\text{PMNS}}$ :  $(\beta\beta)_{0\nu}$ -decay,  $Y_{\text{B}}$ .

Any of the CPV phases in  $U_{\text{PMNS}}$  can be the leptogenesis CPV parameters.

Obtaining information on Dirac and Majorana CPV is a remarkably challenging problem.

Dirac and Majorana CPV may have the same source.

Low energy leptonic CPV can be directly related to the existence of BAU.

Understanding the status of the CP-symmetry in the lepton sector is of fundamental importance.

These results underline further the importance of the experiments aiming to measure the CHOOZ angle  $\theta_{13}$  and of the experimental searches for Dirac and/or Majorana leptonic CP-violation at low energies.